ATUC June 2016 - ATNF Operations

John Reynolds



CSIR



1

The ATNF



ATNF Operations - organization



Updated 17 May 2016

Staff changes

Welcome

Hayley Bignall (Per) Kevin Ferguson (Per) Eric Bastholm (Per)

Jimi Green (Syd) Steve Ord (Syd) Edward Guntek (Syd)

Steve Hathway (Get)

Farewell & Thanks

Kate Brooks (Per)

Christoph Brem (Get)

Jessica Chapman (Syd)



ASKAP Operations



ASKAP operations - headline items

- Recent ACES results covered by Dave McConnell
- ATNF Operations Plan with ASKAP released
- BETA operations ceased in Feb 2016
- Geraldton staffing close to steady-state levels
- Cut-over to new power station, July 2016
- "FIFO" model for site access working well
- Early Science commencement ~Sep 2016
- Tied-array mode now a new project covered in Tasso's talk



CSIRO's Stage 2 renewable expansion



CSIRC

ASKAP delivery

Integrated schedule

	2014 2015				2016															
	Q1	Q2	Q3 (24	Q1	Q2	JUL A	UG SE	р ост	NOV	DEĊ	JAN	FEB	MAR	APR MAY JUN	JUL A	AUG SEP	ост	NOV	DEC
Mk II PAF Production	Spec verif	ification ication	on n	n Proc	duction of	n of P	AFs 1-8					Productic		luctio	on of DATE 0.20					
													FIOC	Juctio						
Firmware deliverables	De	velop	ment o	of co	rrela	or for	Mk II P	AFs			Fast	/slow	, transi	ients,	zoom modes					
Commissioning & verification	BETA com	missio	Comr	nissi	oning	and e	arly sci	ence w	vith BETA						Commissioning PAFs 7-12	5				
a vernication	M co	Mk II comn	PAF pr nission	otot ing a	type at MRO			Commis PAFs	sioning 1-6	ng					Comm PAF:	ission 13-3	ing 0			
RTC supporting BETA				P/	4Fs 5-6	5 at MRO PAFs 5-12 at MR ASKAP Phase 1 Comple				VRO	Early science with PAFs 1-12									
Prototype Mk II P Mk II PAF correlat	AF at or at	MRO MRO								7.51			comp	here	1		PAFs 1	3-30 a	t MRO	
Product	ion re of Mk	adine II PAF:	ss revi s comp	ews olete																
PAFs 1-4 at MR				MRO																



The Early Science Program

• A wide area continuum survey

(1 MHz and 18.5 kHz resolution, full Stokes, 700-1800 MHz, 6-12 hours per field)

• A few 30 square degree fields studied in neutral hydrogen

(18.5 kHz resolution, 1130-1430 MHz, 120 hours integration time per field (TBC).

• A single deep HI field

(18.5 kHz resolution, 1000-1300 MHz)

- A science program at high spectral resolution (details TBC)
- A science program to probe the variable and transient radio sky (details TBC)

Other ideas requiring very small observing times (e.g. MWA EoR field) may be observed on an "opportunistic" basis as part of commissioning.



Estimated time available for early science

Program duration: 12-14 months ≈ 9400 hours

Fraction for early science $25\% \approx 2300$ hours

Minus ~20% overheads (beamforming, calibration etc.)

Total available on sky ≈ 1800 hours*

Estimated commencement of Early Science operations: September 2016

Contact for further information: Lisa Harvey-smith



Status & Impact



Media impacts

- Traditional media releases
- Social media campaigns (#CSIROspace)
- ATNF website news
- Events/meetings

The media landscape continues to evolve!

- Sunrise Breakfast: live cross (Jul 2015)
- How Ridiculous:145,753 YouTube hits + Sunrise promo
- Eureka Prize finalist in science communications (Lisa Harvey-Smith, mentioned in Hansard)

In planning/negotiations;

- Stargazing Live 2017 (3xBBC, 3xABC, host: Brian Cox)
- Apollo 11 50th anniversary doco (Fairtrade Films)
- The Bachelor (!?)

Impacts;

- Trigger significant traffic to CSIRO media and home pages (6,000+ views for ASKAP HI discovery)
- Strong social media support of space and ATNF's observatories result in significant media pick-up both domestic and internationally



Australian technology behind the world's largest telescope >

Technology developed by CSIRO will sit at the heart of one of the world's biggest ever science projects, following an agreement with China's leading astronomical research organisation. 🖾



05 MAY 2016

Solved! First distance to a 'fast radio burst'>

For the first time a team of scientists has tracked down the location of a fast radio burst (FRB), confirming that these short but spectacular flashes of radio waves originate in the distant universe. 🖾 🗃 🖪



12 FEBRUARY 2016

Aussie innovation helps hunt down gravitational waves >

For the first time in history, astronomers have observed elusive gravitational waves ripples in space time caused by a violent cosmic event taking place in the distant Universe. 🔘



Dark 'noodles' may lurk in the Milky Way >

Invisible structures shaped like noodles, lasagne sheets or hazelnuts could be floating around in our Galaxy radically challenging our understanding of gas conditions in the Milky Way. 🚳



25 SEPTEMBER 2015

Eleven year cosmic search leads to black hole rethink >

One hundred years since Einstein proposed gravitational waves as part of his general theory of relativity, an 11-year search has failed to detect them, casting doubt on our understanding of galaxies and black holes.







Publications



Figure 10: Publications that include data from, or are related to, the Compact Array, Mopra, Parkes, VLBI and ASKAP, grouped by facility for 2011–2015. A small number of papers with data from more than one facility are counted more than once.

Papers and impact for 2014

	No. of papers	No. of citations	Citations/paper
VLA/JVLA	200	1885	9.4
ATCA	70	662	9.5
Parkes	49	588	12.0
LBA	7	76	11.0
Mopra	17	104	6.1
ASKAP	4	26	6.5
Other	72	1075	14.3

Results for ATCA and Parkes (based on ATNF publications lists) examining impact for papers published in 2014, based on citations as given by ADS in the following years.

"Other" – CASS affiliated author(s) using non-CASS facilities



Funding



"Breakthrough Listen" Project

- 25% of Parkes time for 5 years for SETI
- Full cost of operating Parkes for 25% of time (current model)
- No impact on high-ranked proposals
- Uniform LST coverage, time-critical overrides still OK
- Observations commence in October 2016

Science benefits:

- Potential commonalities in new software/hardware
- Includes new versatile backend that will support SETI + other sci.
- Potential for future collaboration / support of new receivers
- Open data and commensal observing



NCRIS funding history

2013-14	CRIS + NCRIS 2013	\$1.4M	
2014-15	CRIS + NCRIS 2013	\$4.5M	
2015-16	NCRIS 2013 + 2015	\$4.8M	(BUT only \$3.4M new)
2016-17	NCRIS 2016	\$3.1M	(BUT only \$2.5M new
2017-	Proposed termination		

2016/17 reduction can be managed

- •Slow down in rollout of PAF systems 31-36
- •Reduce secondments to ASKAP Commissioning & Early Science (ACES)
- •Breakthrough Listen funding balances increase in ASKAP energy costs

Zero NCRIS funding from 2017 poses serious problems for current ATNF operations

•CASS is committed to commissioning and operating ASKAP – little scope for savings
•Contraction of science operations – ATCA and/or Parkes – now become serious options
•Parkes bringing in \$1.5M/y through, BL, has already been streamlined



ATNF Ops – direct costs by facility

Direct costs of facilities for FY15/16:

- ASKAP \$8.3m
- ATCA \$3.5m
- Parkes \$2.2m
- Common items \$1.4m (including LBA)

Have already achieved substantial savings at Parkes, ATCA, Mopra

Recent level of NCRIS funding (~\$4m/yr) comprises around 20% of total costs in ATNF Ops, including overheads and depreciation)



ATNF Ops:

Forecast

ATNF Operations forecast for 2016/17 (direct costs):

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Forecast	2016/17
	(\$m)
Revenue	
Internal	11.9
NCRIS	3.1
Breakthrough	1.1
Other External	0.7
Total	16.8
Expenses	.0
ASKAP	9.0
ATCA	3.6
Parkes	2.3
ATNF Common	1.4
Other (not telescopes)	1.4
Total	17.7

Operating Result

(0.9)



By Telescope (2016-17)

		Other						
2016-17	Labour	Electricity	Operating	Total				
ASKAP	4,368	2,550	2,043	8,961				
ATCA	2,480	309	833	3,622				
Parkes	1,658	221	406	2,285				

Can likely save ~\$1M through;

- •under-spend on ASKAP power
- •delayed deployment of ASKAP 31-36 roll-out





- Recent level of NCRIS funding (~\$4m/yr) comprises around 20% of total costs in ATNF Ops, including overheads and depreciation)
- Zero NCRIS funding from 2017 leaves serious shortfall for current ATNF operations
- CASS is committed to commissioning and operating ASKAP little scope for savings
- Contraction of science operations ATCA and/or Parkes now become serious options
- Parkes bringing in \$1.5M/y through, BL, has already been streamlined (remote observing, reduced config changes, infrastructure upgrades)
- Mopra already self-funded (zero cost to CSIRO)



Other leads being pursued/considered

- Continuing dialogue with AAL re: strategy and timescales
- Possible use of Mopra as downlink for (e-Rosetta, WFIRST)
- Recovery of common costs for running MRO
- Pursue opportunities to sell telescope time on all facilities, possibly in competition with national facility users
- Quantify the benefit to Australia of open access to international astronomy facilities



ATCA – time allocation by base



Figure 4: Compact Array time allocation by all investigators, October 2011–September 2015. Time allocated to each proposal has been divided evenly between all authors on the proposal. For each year the time allocation is for 12 months from October to September.



The ATNF User Community – your input

Obtain broad astronomy community input on the implications of ceasing or substantially decreasing funding for ATCA and Parkes, to inform a decision on a timescale of 6 months.

Decision will involve identifying and evaluating several new scenarios for operating our facilites. In the absence of new funding streams these will very likely include scenarios for significant reduction in National Facility science operations at ATCA and/or Parkes.



We acknowledge the Wajarri Yamatji people as the traditional owners of the Murchison Radio-astronomy Observatory site

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