

CSIRO DIVISION OF RADIOPHYSICS

THE AUSTRALIA TELESCOPE

AT/20.1.1/019

Overall Systems and  
Performance  
Tech. Notes & Reports

COMPUTATIONAL LOAD FOR BANDWIDTH SYNTHESIS

G.J. NELSON  
5th March 1986

Following Forster (AT/20.1/006) we see that acceptable bandwidth smearing requires  $\frac{\Delta\nu}{\nu} = 1.73 \times 10^{-3}$  or  $\sim 48$  channels in a 128MHz bandwidth at L band. Alternatively this implies channel bandwidths of  $\sim 2$  MHz at 1.25GHz. Thus if all Stokes' parameters are required then at one time we can observe two bands each 64MHz wide and each containing 32 X 2MHz wide channels. Four separate sets of such observations at different frequencies will cover the whole L band for bandwidth synthesis observations.

In summary we can say that after an integration time of 20 seconds (5 seconds on each pair of 64MHz bands) we will have accumulated  $4 \times 2 \times 4 = 32$  correlation functions per baseline in the worst case. Thus each of  $15 \times 32 = 480$  correlation functions need be transformed into 32 complex frequency components. If the transform times scale as  $N \log N$  then  $480 \times 32$  point complex transforms take only 5% of the time required for the worst case ( $15 \times 8$  192 point complex) transforms to be performed during line observations.

Thus the worst case bandwidth synthesis observations place quite small loads on the correlator array processor.