

# Large Magellanic Cloud Distance from Cepheid Variables using Least Square Solutions

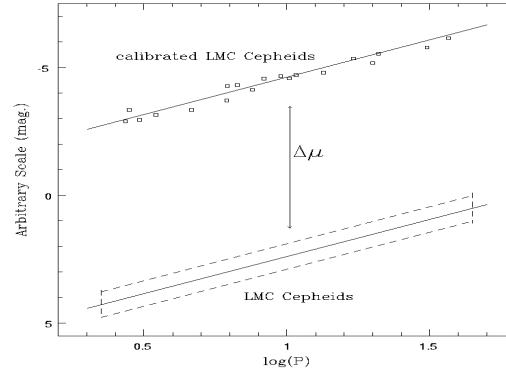
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**ABSTRACT:** We determine the LMC distance using Cepheid variables in the LMC. We combine the individual LMC Cepheid distances obtained from the infrared surface brightness method and a dataset with a large number of LMC Cepheids. Using the standard least squares method, the LMC distance can be found from the ZP offsets of these two samples. We have adopted both a linear P-L relation and a "broken" P-L relation in our calculations. The resulted LMC distance moduli are  $18.478 \pm 0.034$  mag and  $18.486 \pm 0.036$  mag (the quoted errors are random error only), respectively, which are consistent to the adopted 18.50 mag.

## I. Introduction

- Gieren et al (2005, hereafter G05): determined the distance to 13 LMC Cepheids (including 6 short-period Cepheids in LMC cluster NGC 1866) and derive the Cepheid period-luminosity (P-L) relation.
- LMC hosts >600 Cepheids available from the Optical Gravitational Lensing Experiment (OGLE) database.
- By combining these two datasets, it is possible to determine *both* of the LMC distance modulus and P-L relation simultaneously (see **Figure 1**).

**Figure 1:** Illustration of the *least squares method*. Introducing an offset of the ZP,  $\Delta\mu$ , between the calibrated Cepheids from infrared surface brightness method and a large number of OGLE Cepheids will allow the LMC distance and the P-L relation to be solved simultaneously.



## II. Data & Methods

- Data used: (1) G05 (2) OGLE data from Kanbur & Ngeow (2006).
- The least squares take the form:
$$m = \alpha \Delta\mu + a + b \log(P), \alpha = \begin{cases} 0, & \text{if G05} \\ 1, & \text{if OGLE} \end{cases}$$
- Recent studies have suggested the LMC Cepheid P-L relation is non-linear (Kanbur & Ngeow 2004, 2006; Sandage et al 2004).
- We also adopt a "broken" P-L relation with a break period at 10 days.

## III. Results

- For linear P-L relation:

*V*-band:  $\Delta\mu = 18.468 \pm 0.059$ ;  $a = -1.363 \pm 0.068$ ;  $b = -2.761 \pm 0.036$   
*I*-band:  $\Delta\mu = 18.483 \pm 0.041$ ;  $a = -1.863 \pm 0.046$ ;  $b = -2.981 \pm 0.024$

- For non-linear P-L relation:

*V*-band:  $\Delta\mu = 18.485 \pm 0.063$  & *I*-band:  $\Delta\mu = 18.486 \pm 0.043$   
 Long period ( $P > 10$  days):  $a_V = -1.228 \pm 0.208$ ;  $b_V = -2.838 \pm 0.155$   
 $a_I = -1.688 \pm 0.143$ ;  $b_I = -3.093 \pm 0.106$   
 Short period ( $P < 10$  days):  $a_V = -1.278 \pm 0.072$ ;  $b_V = -2.939 \pm 0.061$   
 $a_I = -1.801 \pm 0.049$ ;  $b_I = -3.092 \pm 0.041$

- *V* and *I* band averaged distance modulus is given in the ABSTRACT.
- F-test (Kanbur & Ngeow 2004) results for non-linear P-L relation:  
 $F_V = 7.32$ ;  $F_I = 6.87$   
 (null hypothesis [linear P-L relation] can be rejected with >95% CL if  $F > 3$ ).

**REFERENCE:** (i) Gieren et al 2005, ApJ 627:224 (ii) Kanbur & Ngeow 2004, MNRAS 350:962 (iii) Kanbur & Ngeow 2006, MNRAS 369:705 (iv) Sandage et al 2004, A&A 424:43