The 20 GHz Pilot Survey - What did we find?

Elaine M. Sadler (on behalf of the 20 GHz Survey team)

• Preliminary analysis of scans and 18 GHz images from Sept/Oct 2002 ATCA sessions

• Do we see a 'new' source population?

• What kinds of astrophysical objects are these?

(Follows on from Roberto Ricci's talk...)

The 20 GHz data set

• 226 detected (5 σ) sources above ~60 mJy, Dec strip at -60 to -70, RA 0h to 24h

• More than half lie within 10° of the Galactic Plane or in the LMC (HII regions, PNe etc.)

• Work so far: cross-matching and identifying 82 sources at high Galactic latitude.

• Cross-match with 843 MHz SUMSS catalogue to get two-point spectral indices

Optical ID s from NED and Cosmos

WBC-SUMSS cross-match



• All (*|b|*>10°) sources detected at 18 GHz are also present in the 843 MHz SUMSS catalogue (and well above the 6 mJy limit)

• The 843 MHz and 18 GHz flux densities are essentially uncorrelated

No 'new' sources yet!

Radio spectral indices



• The sample is dominated by sources with flat/inverted radio spectra.

• No obvious sign that α varies with flux density

• Wide range in α

Optical identification



• Both ATCA and SUMSS have good positional accuracy (~1 arcsec).

• Cross-match with NED for existing IDs

• Other optical IDs via Cosmos cat.

Optical ID rate

High optical ID rate for 18 GHz sources,
(80/82 have a candidate DSS ID within 8 arcsec,
versus ~35% for low-freq surveys)

Optical IDs for 82 18 GHz sources:

- 13 galaxies
- 21 catalogued QSOs

38 new candidate QSOs (~20 likely to be genuine)

10 faint objects or blank fields

PKS 0313-660 - a known OSO (z=0.636)



PMNJ 0150-6044 - a new QSO candidate



IRAS 23074-5957 - a galaxy at z=0.142



20 GHz source populations



Most optical counterparts are bright!



Optical follow-up of most of the 18 GHz sources needs only a 2mclass telescope.

Summary of results ...

• Most of the extragalactic 18 GHz sources appear to be QSOs

• Also some distant radio galaxies, and interesting rarer populations such as IRAS galaxies

• Most optical counterparts are bright, so follow-up spectroscopy should be easy

• A substantial population of Galactic and LMC sources is also present