A search for water masers in the Magellanic Clouds (C973) Beasley, Claussen, Marvel & Staveley-Smith

AIMS

- Identify more strong masers in LMC & SMC
- Study relationship with HII regions and AGB stars
- Maser luminosity function
- Monitor variability
- Multi-epoch VLBI study of stable masers to measure PROPER MOTION (1 marcsec yr⁻¹) and ANNUAL PARALLAX (40µarcsec) of Magellanic Clouds

Observations

- 6_{16} - 5_{23} water maser transition at 22.235 GHz
- Two epochs Feb 2002 & Nov 2002
- Each epoch 3 x 12-hr obs, spread over 6 days to allow for bad weather, nighttime
- Generally 1-2 hour angles, 2-5 mins integration
- 500 sources
- Typical elevations: 30-60 degrees

Observations (cont)

- Channelization: two types
 - 16 MHz 512 channels (search)
 - 8 MHz + 128 MHz (known sources)
- Mosaic mode to reduce overhead time for snapshots and make use of short coherence times during bad weather
- Phase cal every 10-15 mins

Feb 2002

- Night 1
 - Weather poor (overcast, some rain)
 - Tsys 150-200+ K (varied by 50-80 K)
 - High Tsys + high Opacity = unable to detect 2-10 Jy (peak) masers
 - CACAL on 1934-638 @16 MHz BW poor choice
 - Detection limits: 0.4 km/s, 1-2 mins > 1 Jy

Feb 2002

- Night 2 & 3
 - Better weather clearer, Tsys ~120 K (varied from 80 160 K)
 - CACAL on 0537-441 (7 Jy) 🙂
 - Phase stability good coherence times of many 10's of mins or more at times
 - Good (new) detections @ 1-2+ Jy

Nov 2002

- All nights
 - CACAL on 1921-293 (12 Jy, assumed) ③
 - Also: Mars observed each night
 - Weather good: Tsys 120K, 120K, 60-100K
 - Phase stability excellent: coherence times > 1-2 hours
 (!)
 - System: good performance, a few ACCs died
 - Good detections, 0.4 km/s channel rms ~0.2 Jy after 2-3 minutes

Calibration

- First epoch
 - CACAL: used 0537-441 (7 Jy), 1934-638 (~0.4 Jy)
 - 0454-810, 0100-760 phase cals
 - Mars observation included (not used yet)
- Second epoch
 - CACAL: Used 1921-293 (12 Jy), also assumed it at 12
 Jy to set flux scale (interp. VLA value)
 - 0454-810 (2.5 Jy), 0100-760 (0.75 Jy) phase cals
 - Mars observations included (3", 0.4 Jy)

Data Reduction

- During run observed spectra on SPD in baseline average mode – most detections easily visible after 60 secs.
- Also: AIPS reduction (Beasley's preference).
- No task to implement Mars calibration available in AIPS (Beasley plans to write one...)

Misc. info

- Tsys increase with elevation obvious below 30 degrees poor (Tsys > 150 K)
- Dawn effect phase stability ruined about 1-2 hours after dawn (C1116).
- Mosaic mode occasional mislabeling of source names seen in dataset (i.e. when changing mosaic files?).
- SYSC option in ATLOD doesn't work for 12mm data (?) (or is it for mosaic data?)

Results

- Nine new detections (8 in LMC, 1 in SMC)
- 7 IRAS or MSX-selected HII regions and 2 O-rich AGB stars
- 30+ spectral features
- Flux densities > 0.5-1 Jy
- 2% detection rate



Typical detection: Epoch 2, night 3



Previous Parkes detections

(Scalise & Braz 1982; Whiteoak et al. 1983; Whiteoak & Gardner 1986)

New ATCA detections

22 GHz water masers in LMC Tony Beasley's wish list:

-Easy "Dip" routine to determine tau (exists?), allow you to quantify atmos. part of Tsys

–Mosaic mode where RA/DECs of grid points specified (rather than degree offsets.. boring)

–Monitoring of 1921-293 (+ others) vs. Mars et al. for simple flux calibration, CACAL

-Good cal list (underway..)