


Deconvolution and Multi frequency synthesis

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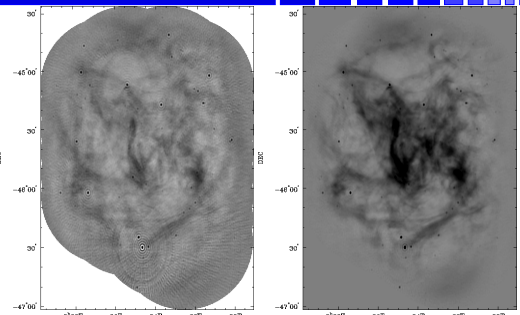
Deconvolution

- Basics (again!)
- Multi-frequency synthesis
- Characteristics of the dirty beam
- Linear deconvolution
- Constraints
- CLEAN
- Maximum entropy
- Restoration
- Multi-frequency deconvolution


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An example of deconvolution



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Filling the Fourier plane

- Use many antennas (6 antennas or more)
- Use Earth rotation (12 h observations)
- Physically move antennas

But the aperture is *NEVER* completely filled

- Limited observing time
- Limited number of antennas
- Various interruptions to the observation
- Min and max baselines

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Multi-frequency synthesis

As (u, v) coordinate is measured in wavelengths, another way of filling the Fourier plane is to observe at multiple wavelengths simultaneously.

Basic imaging relationship

$$V(u, v) = \int I(\ell, m) \exp(i2\pi(u\ell + vm)) d\ell dm$$

Using a “direct Fourier transform” we produce the dirty image

$$I_D(\ell, m) = \sum_k w_k V(u_k, v_k) \exp(-2\pi i(u_k \ell + v_k m))$$

$$= \mathcal{F} \left[\left(\sum_k w_k \delta(u - u_k, v - v_k) \right) \cdot V(u, v) \right]$$

$$= \mathcal{F} [S(u, v) \cdot V(u, v)]$$

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Convolution relationship

Fourier theory tells us that

$$F(u)G(u) \Leftrightarrow f(\ell) * g(\ell)$$

so

$$I_D(\ell, m) = B(\ell, m) * I(\ell, m)$$

where

$$B(\ell, m) = \mathcal{F} \left[\sum_k w_k \delta(u - u_k, v - v_k) \right]$$

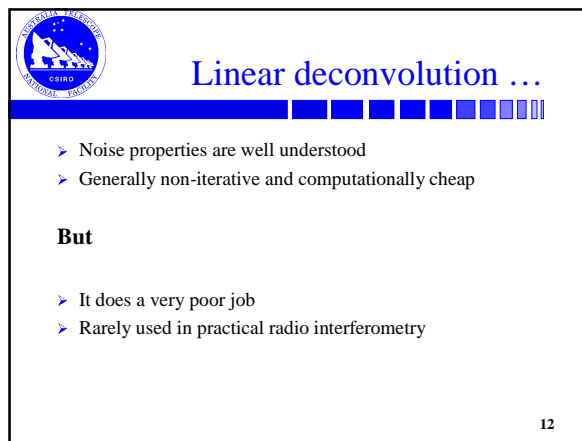
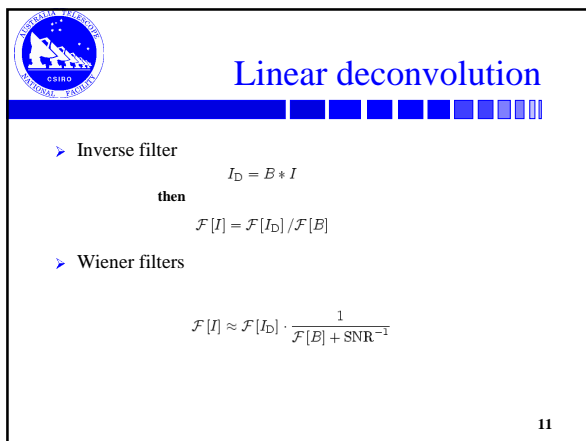
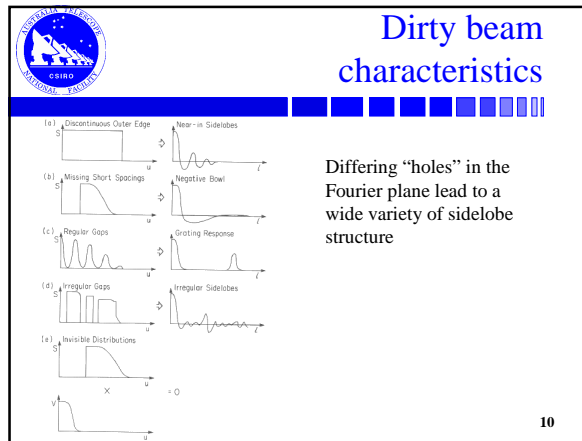
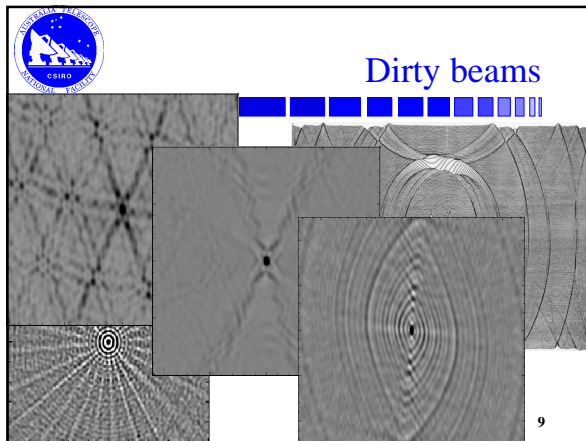
Jargon: The point-spread function is usually called the “beam”.


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Deviations from convolution relationship

- Wide-field effects are usually neglected. These include
 - Time and bandwidth smearing
 - Primary beam effects
 - So-called non-coplanar baseline effects
- Convolution relationship strictly applies only for continuous functions (not a sampled grid of pixels).
- “Aliasing” in the imaging process is also not accounted for.
- Finite extent assumption.

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Non-linear deconvolution

- Linear deconvolution is fundamentally unable to extrapolated unmeasured spatial frequencies.
- A function which is non-zero only in the unsampled part of the Fourier plane is called an *invisible distribution*.
- A good non-linear deconvolution algorithm is one that picks plausible invisible distributions to fill in the Fourier plane.


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Prior Information or Assumptions

- Bounded support (“CLEAN boxes”).
- Positivity
- The sky is mostly empty
- Use a goodness measure to pick “reasonable” solutions.


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CLEAN Algorithm (Högbom, 1974)

- Assumes that the sky can be modelled as a collection of point sources.
- Iteratively decomposes the sky into a collection of point sources.
- In principle, CLEAN is guaranteed to converge, although in practice it can become unstable if pushed too far.
- Generally it is quite a robust algorithm.

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CLEAN algorithm

1. Search for the largest peak in the residual image
2. Assume this is a result of a point source – a component!
3. Subtract off some fraction (“damping factor” or “loop gain”) of the point source.
4. Add that fraction of the point source to a component list.
5. *Iterate*


Iteration stops when the residual is below some cut-off, when a negative component is encountered, or when a fixed number of components are found.

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 **CLEAN implementations**


- There are different implementations of the algorithms (with their individual strengths and weaknesses):
 - Högbom algorithm – the classical one
 - Clark algorithm – faster for large images with many point sources.
 - Cotton-Schwab (“MX”) algorithm – works partially in the visibility domain. Able to cope with extra artefacts. Can be slow.
 - Steer Dewdney Ito algorithm – works best for very extended objects.

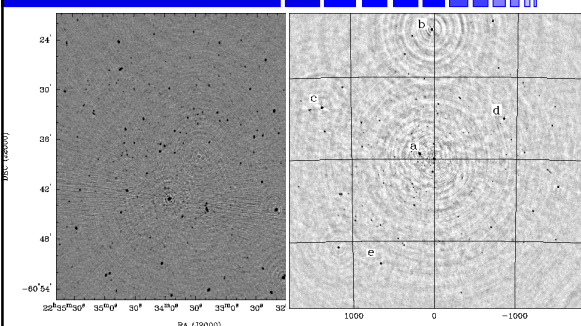
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 **Strengths/weaknesses**


- CLEAN is good for fields of sources which are unresolved or just resolved.
- Generally quite robust in the face of many defects.
- CLEAN is very poor for very extended objects:
 - Slow!
 - Corrugation instability.
 - CLEAN poorly estimates broad structure (short spacings). The result is the so-called “negative bowl” effect.
- CLEAN’s procedural definition makes it difficult to analyse.

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 **Examples of CLEANed images**



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 **Bayesian Statistics and Maximum Entropy**

- Two basic views of probability theory:
 - Views probability distribution function as a measure of the relative frequency of an outcome.
 - Views probability distribution function as a reflection of our uncertainty.
- Principle of maximum entropy:

Of all the possible probability distributions which are consistent with the available information, the one that has the maximum entropy is most likely the correct one.
- Maximum entropy image deconvolution:

Of all the possible images consistent with the observed data, the one that has the maximum entropy is most likely to be the correct one.

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Maximum entropy


- Of all the possible images, pick that one which maximises some goodness measure called “entropy”.

$$\mathcal{H} = \sum f(I(\ell, m))$$

- The most popular choice is the entropy function

$$f(I) = -I \log(I/eM)$$


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Maximum entropy

- The solution is generally constrained so that a χ^2 measure is consistent: i.e. the χ^2 measure is consistent with the expected noise level.
- Integrated flux constraint can be included.
- “CLEAN box” constraint is readily added.
- The default image, M , can be chosen to be a uniform value, or can be set to some prior expectation of the source.
- Solution image must be positive-valued.


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Strengths/weaknesses

- Fourier extrapolation tends to be more conservative than CLEAN.
- Tends to work better for images with a large amount of extended emission.
- Tends to be faster for large images (> 1024x1024 pixels).
- Susceptible to analysis.
- Depends more critically on its control parameters (e.g. noise variance and integrated flux).
- More likely to blow up on poorly calibrated data, or data that violates the convolution relationship in some way.
- Poorly deconvolves point sources.

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CLEAN vs MEM

The answer is image dependent:

- “High quality” data, extended emission, large images
⇒ Maximum entropy
- “Poor quality” data, confused fields, point sources
⇒ CLEAN

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“Restoration” Step


- CLEAN and MEM “super-resolve”, and the high spatial frequencies can be of poor quality (particularly CLEAN).

Solution: Downweight the high spatial frequencies by convolving with a gaussian “CLEAN beam”.

$$\hat{I}_R = \hat{I} * B_G + \underbrace{(\hat{I}_D - \hat{I} * B)}_{\text{residuals}}$$

- The CLEAN beam usually has the same FWHM as the main lobe of the dirty beam.


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Why include the residuals?

- The residuals give an easy way of seeing how believable the features in an image are.
- The residuals still contain emission from sources that have not been CLEANed out.


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Multi-frequency deconvolution

- Multi-frequency synthesis uses observations at many frequencies to prove the Fourier plane coverage.
Problem: *Source structure is a function of frequency.*
- For modest spread in fractional bandwidth (< 15%), and modest dynamic range (< 500), the errors caused by source structure varying with frequency can be ignored.
- When this is not the case, a multi-frequency deconvolution algorithm can be used to eliminate the resultant errors.

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Multi-frequency deconvolution algorithm

- The algorithm models the spectral variation at each pixel as a constant and a linearly varying component with frequency.
- The response to the constant part of this variation is just the normal dirty beam.
- The response to the linearly-varying component can be represented by a second response function. The dirty image is the sum of the responses to the constant and varying components.
- A joint deconvolution, simultaneously solving for the two components can be performed.
- In *Miriad*, this is the so-called mfclean algorithm.

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