



Correlators

Max Voronkov | ASKAP Software scientist

Astronomy and Space Science
www.csiro.au

Narrabri - 2nd October 2014
+ some slides from John Tuthill

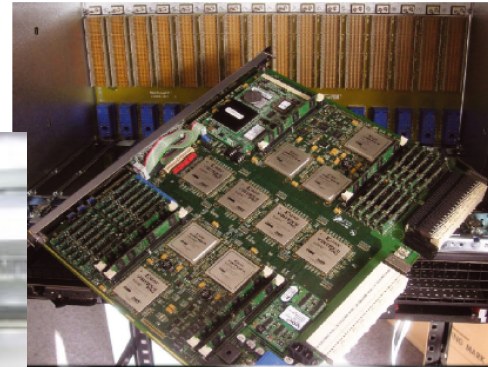


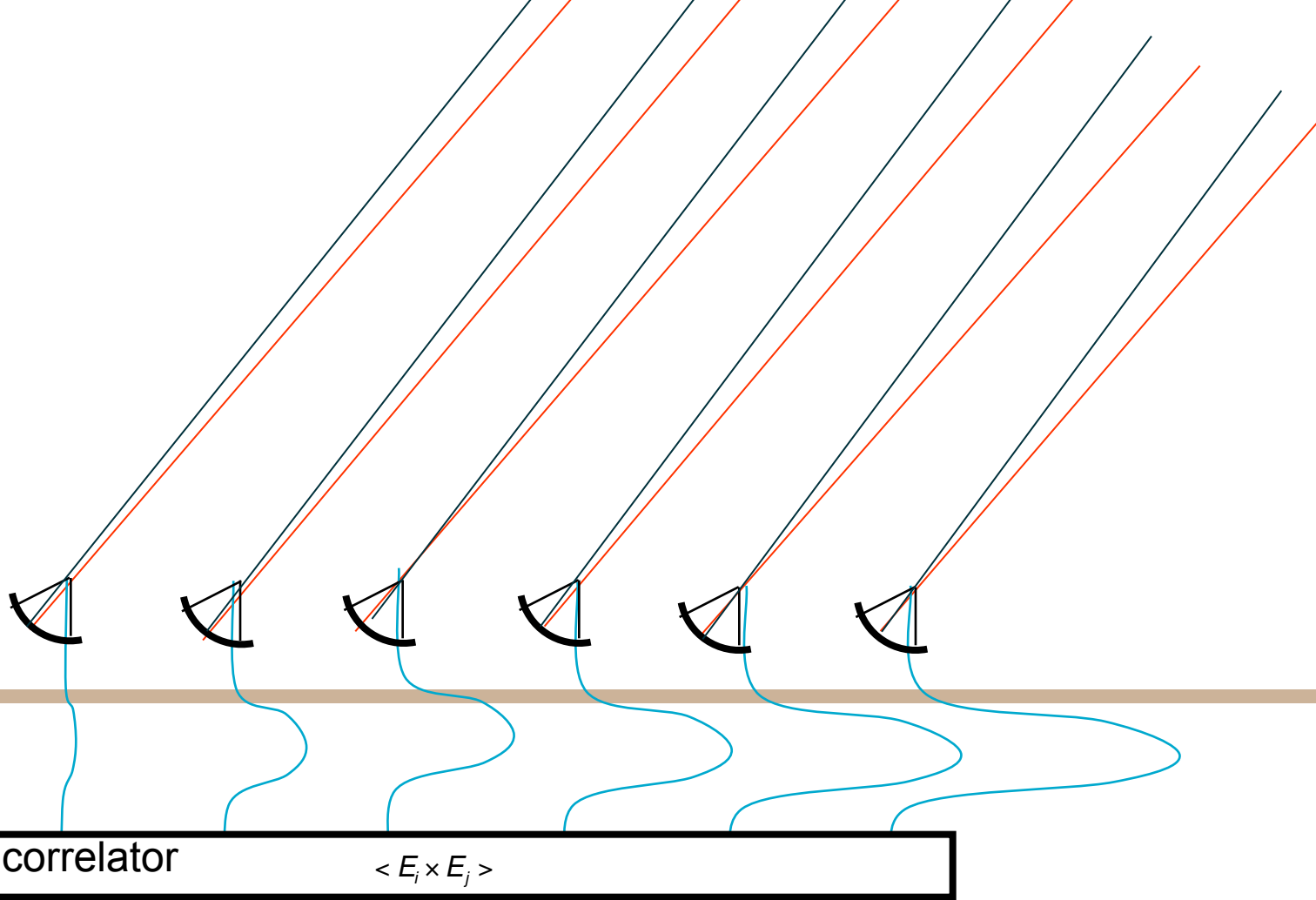
Correlator is the brain
of radio-interferometer
(as it computes visibilities)



”black box” for most science work

May need to know details,
if you start pushing the limits





van Cittert-Zernike theorem

Fourier Transform


$I(r)$

Image credit: Ron Ekers

Simplest possible case

Assuming bandwidth is small enough & we don't care about frequency structure

$$V_{ij} = \frac{1}{N} \sum_{k=1}^N E_i(k) \times E_j^*(k-\tau)$$



Compute correlation directly
(e.g. in software)

Code snippet from BETA-3 (3-baseline) software correlator:

```
IndexType offset1 = itsDelay1;  
IndexType offset2 = itsDelay2;  
IndexType offset3 = itsDelay3;
```

```
for (; (offset1 < size)&&(offset2 < size)&&(offset3 < size); ++offset1, ++offset2, ++offset3) {  
    itsVis12 += *(stream1 + offset1) * conj(*(stream2+offset2));  
    itsVis13 += *(stream1 + offset1) * conj(*(stream3+offset3));  
    itsVis23 += *(stream2 + offset2) * conj(*(stream3+offset3));  
}
```

Non-monochromatic input signal

More intuitive description via continuous formalism:

$$E(t) = \int \mathfrak{s}(\nu) \exp\{2\pi j\nu t\} d\nu$$

spectral representation

Correlation between antenna 1 and 2 data streams for a given lag:

$$\gamma(\tau) = \int E_1(t) E_2^*(t - \tau) dt = \int \underbrace{\mathfrak{s}_1(\nu) \mathfrak{s}_2^*(\nu)} \exp\{2\pi j\nu \tau\} d\nu$$

Power (cross-correlation) spectrum

Option 1

Correlate streams for a number of lags and Fourier-transform

Option 2

Lag or XF correlator

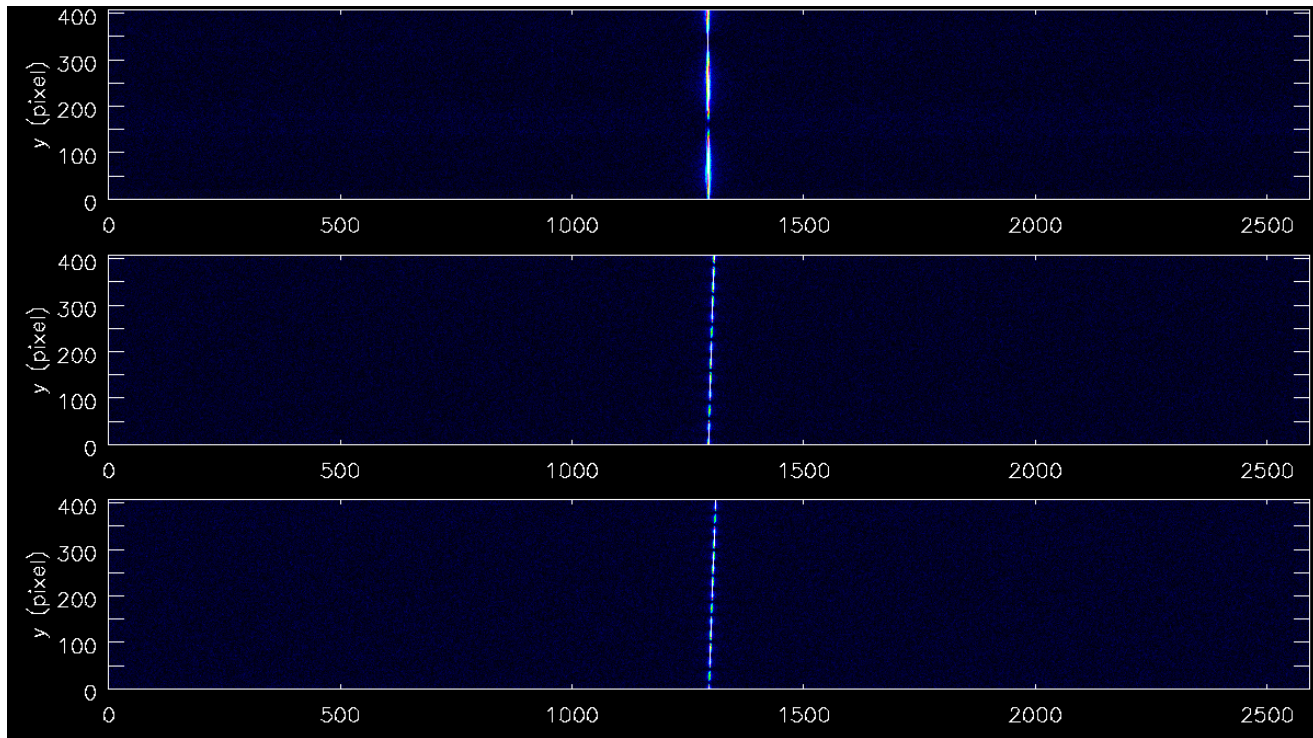
Fourier-transform input streams and cross-multiply

FX-correlator

More on lag domain

Regardless of the correlator architecture, lag spectrum

$$\gamma(\tau) = \int V_{12}(v) \exp\{2\pi jv\tau\} dv \quad \text{is a useful diagnostic tool}$$



AK04-AK05

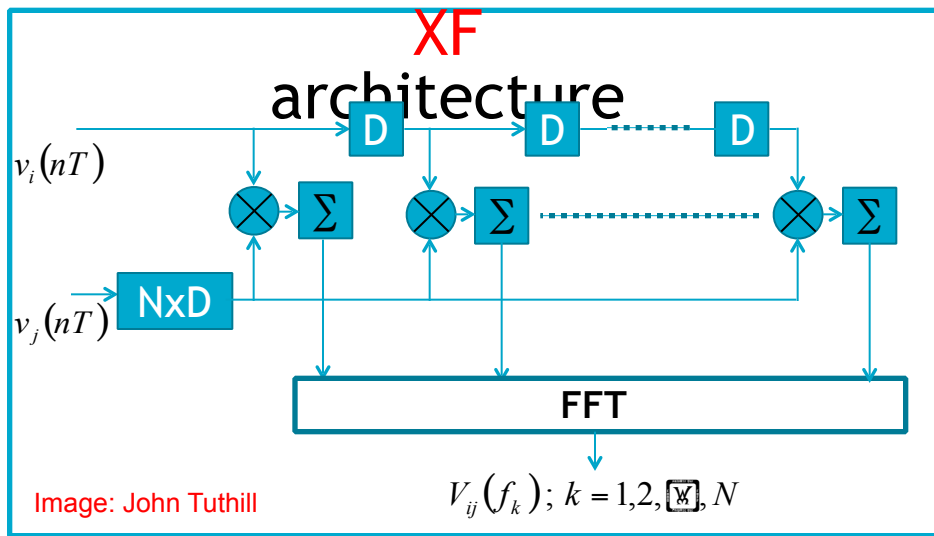
AK04-AK12

AK05-AK12

Lag (1296 pixel is the centre)

Lag (XF) correlator

Old ATCA correlator (pre-CABB) is a good example



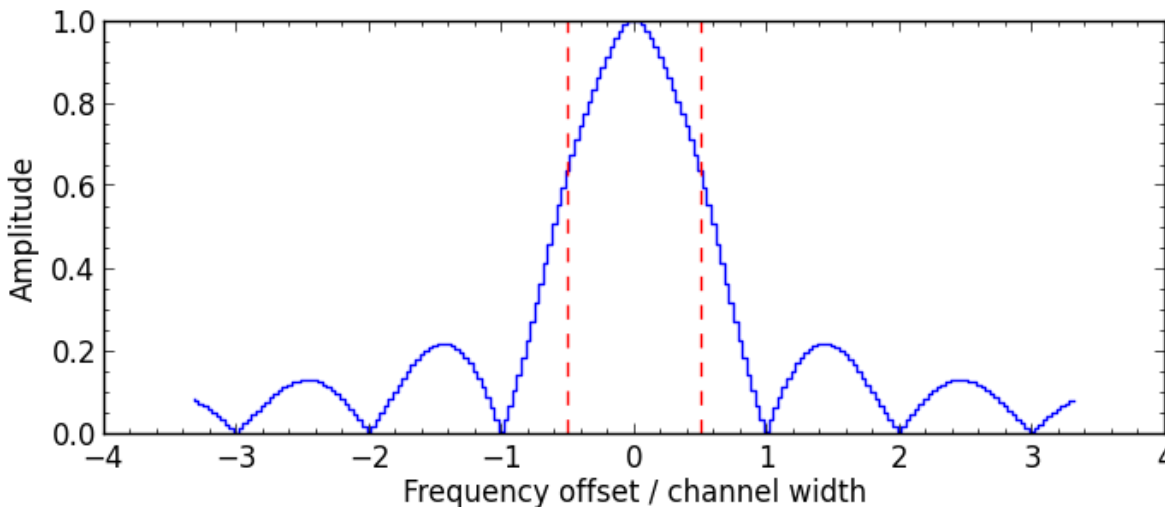
Fractional delays (< 1 sample) need to be corrected in some other way

$$\gamma(\tau) = \int V_{12}(v) \exp\{2\pi jv\tau\} dv$$

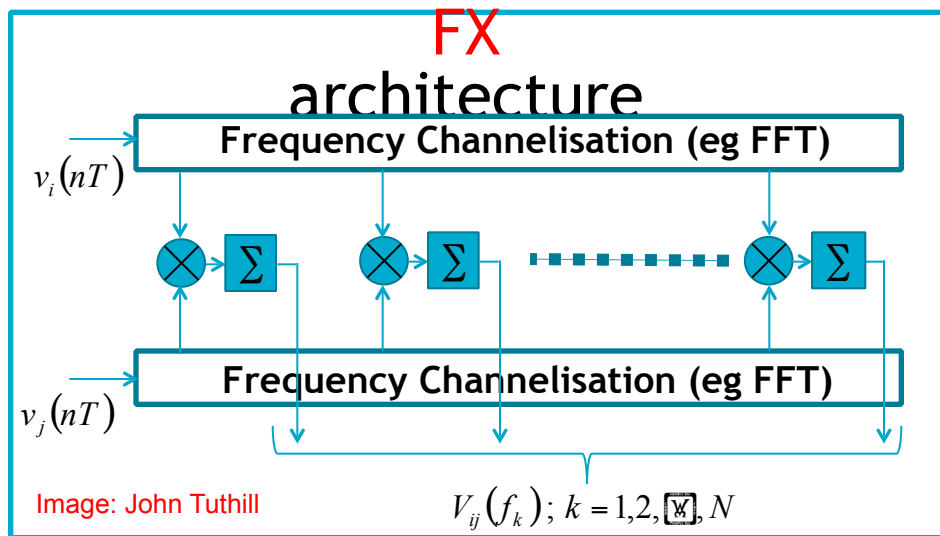
Only finite number of lags can be measured



Convolution with *sinc* in the frequency domain

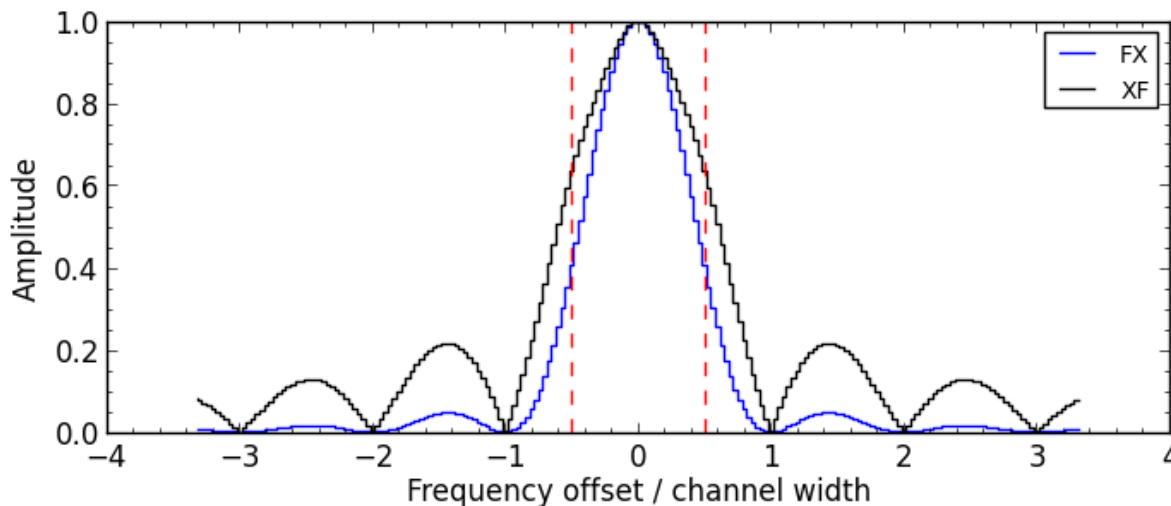


FX correlator



Fractional delays are easy to implement

Output is the product of two Fourier Transforms, each is presented with a finite chunk of data

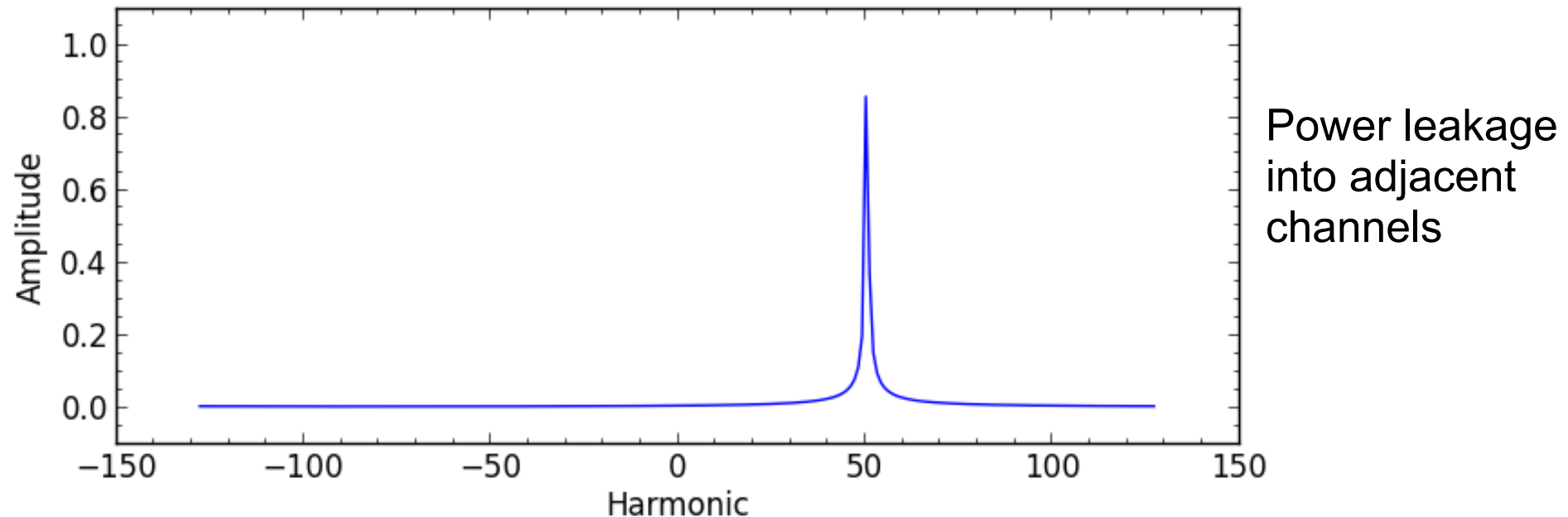


Convolution with sinc^2 in the frequency domain

Can we get a better channel response?

Let's consider FX architecture as a channeliser + cross-multiplier
(think of a simple correlator described earlier)

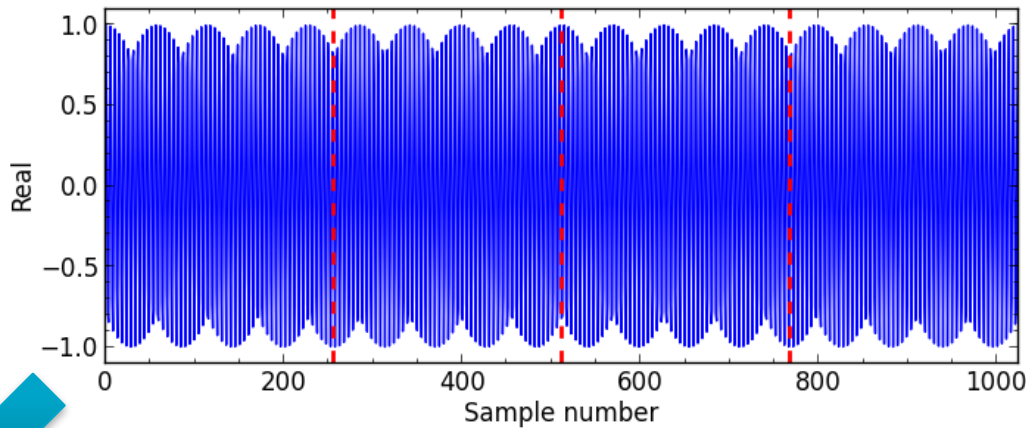
The channel response is determined by the channeliser performance
and FFT is known to be quite bad



FFT of the sine wave with frequency of $50.3 * f_s / 256$

Filtering

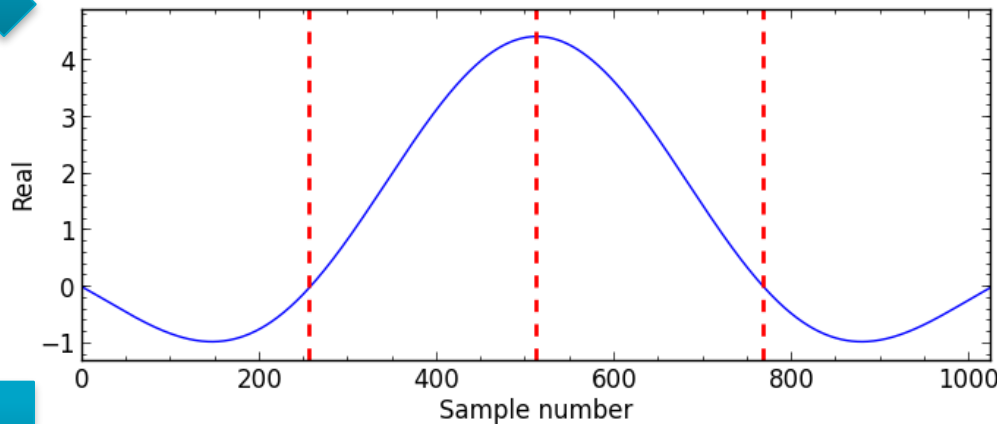
Input data samples



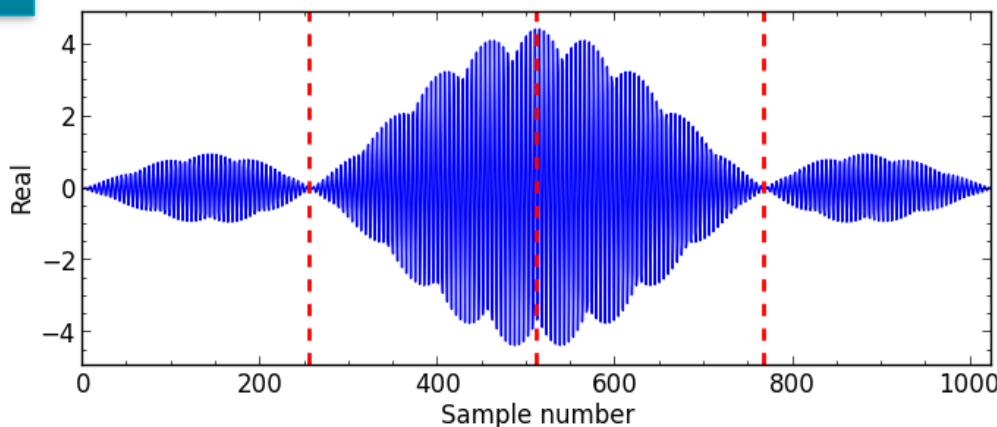
Sine wave with the frequency of $50.3 f_s / 256$

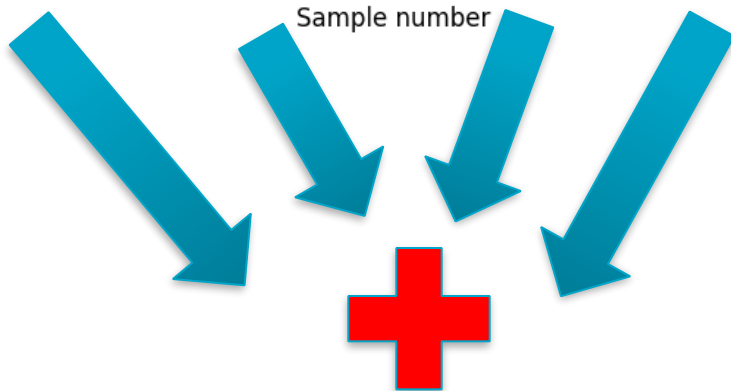
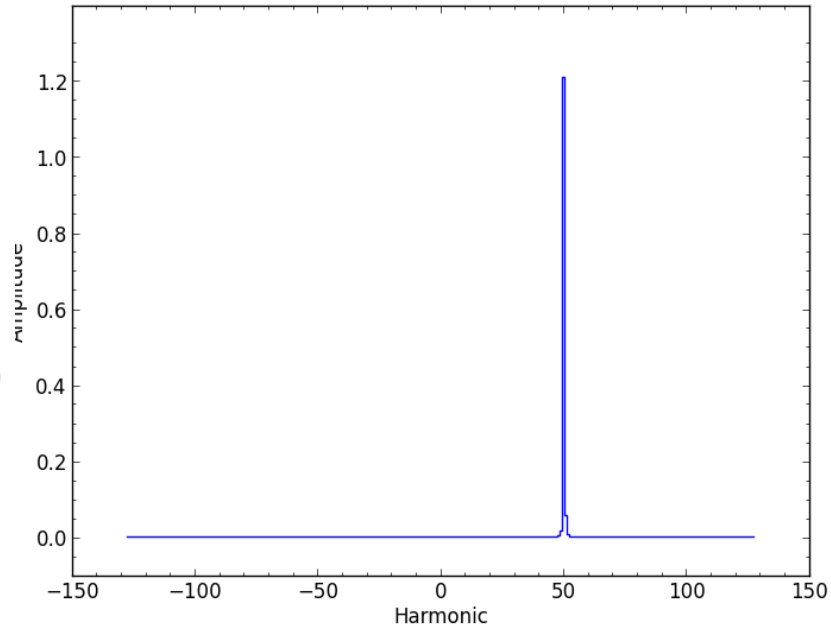
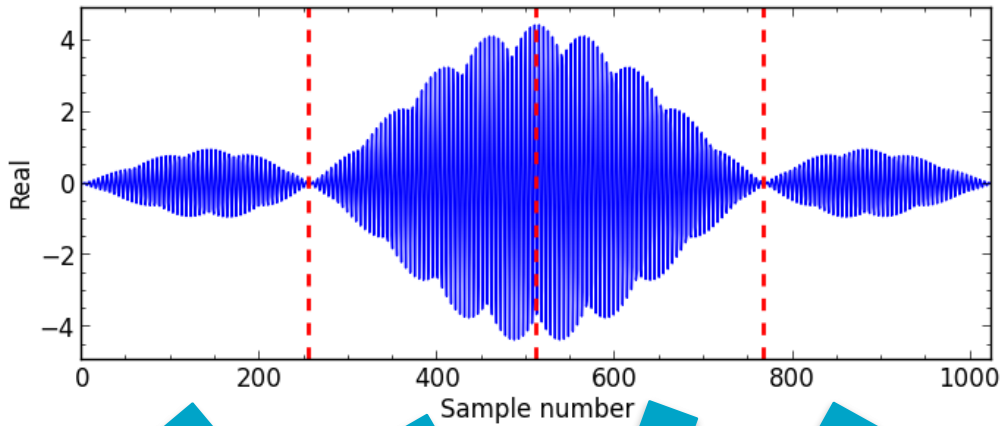
Filter

(Sinc in this particular example)



Filtered data samples





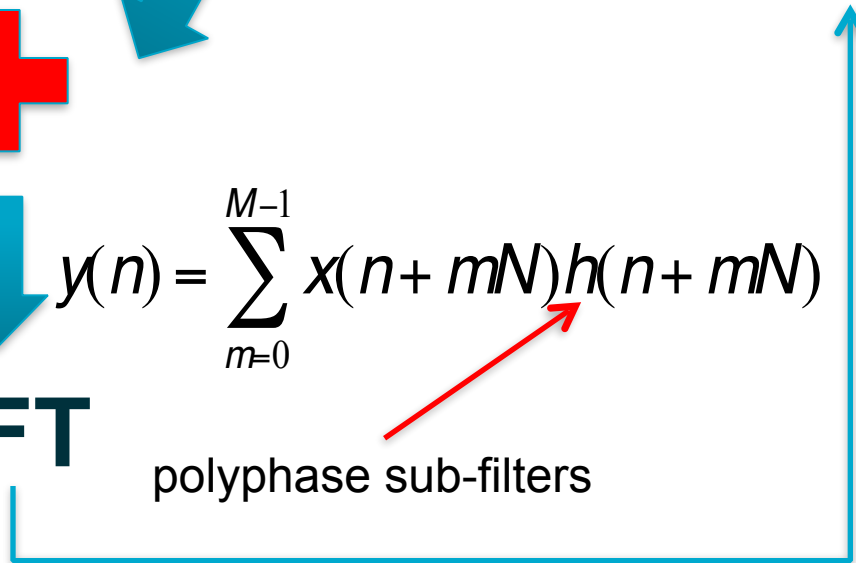
In this example:

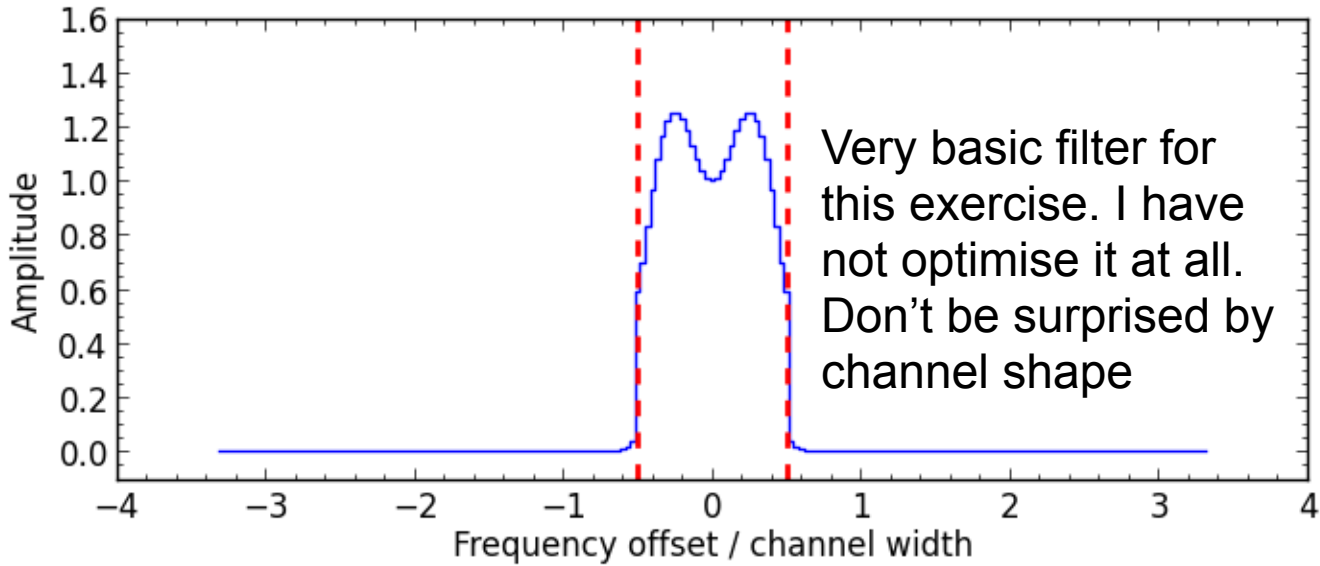
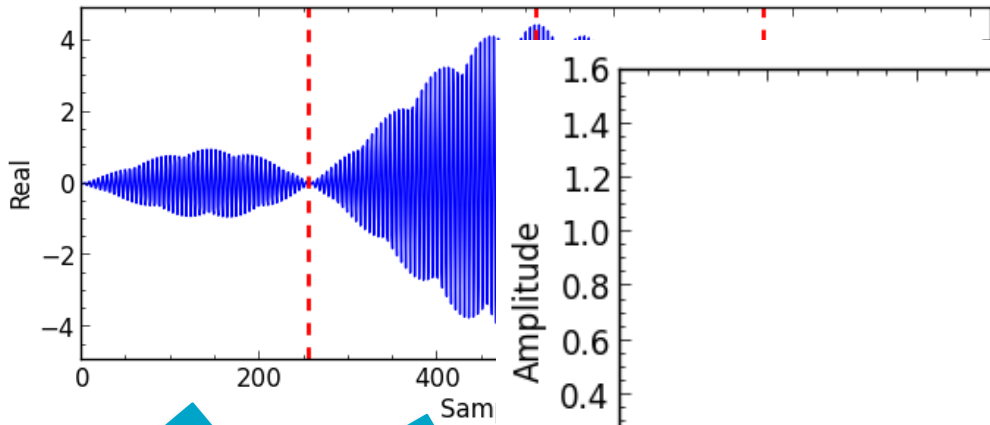
$N = 256$
 $M = 4$

$$y(n) = \sum_{m=0}^{M-1} x(n + mN)h(n + mN)$$

FFT

polyphase sub-filters





In this example:

$N = 256$
 $M = 4$

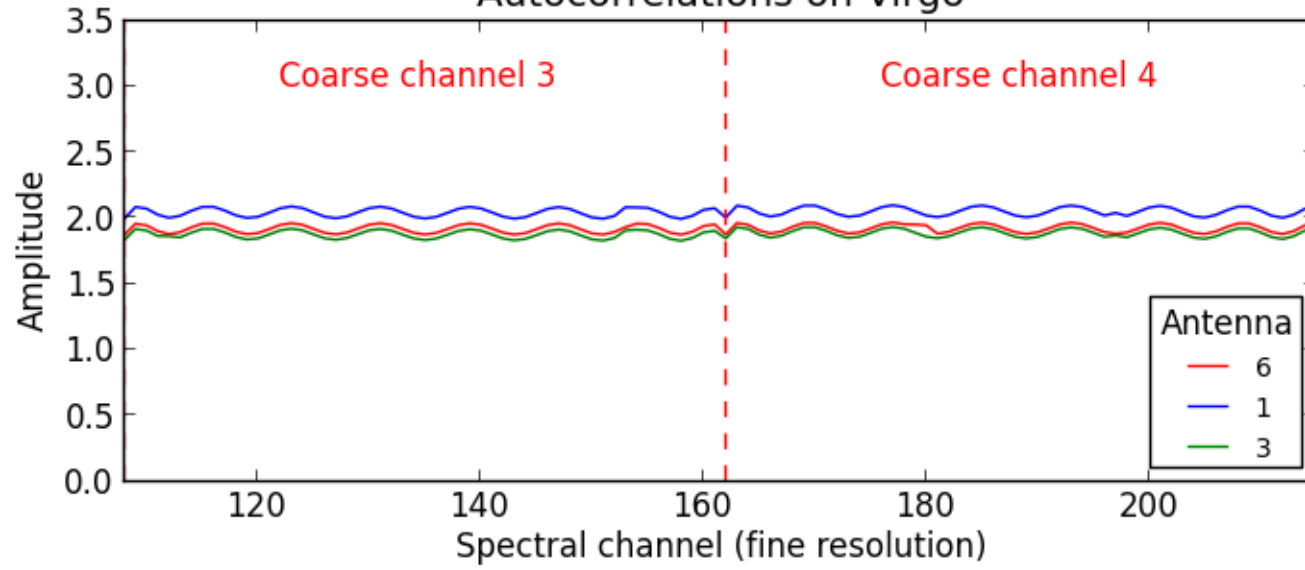


$$y(n) = \sum_{m=0}^{M-1} x(n + mN)h(n + mN)$$

FFT

polyphase sub-filters

Autocorrelations on Virgo



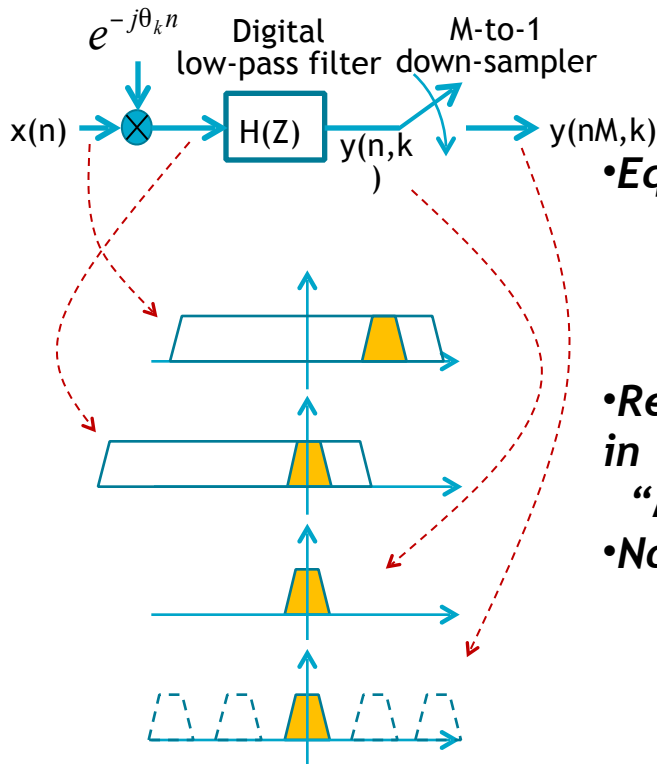
**But filter does
cause a ripple
(can be corrected)**

← Early BETA tests



Polyphase decomposition: engineer's view

Standard single-channel down converter



•Equivalency Theorem

•Exchange mixer and low-pass filter with a band-pass filter and a mixer.

•Re-write the band-pass filter in

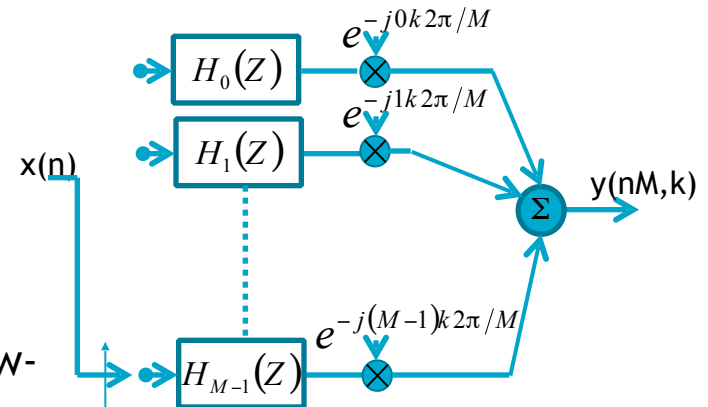
“M-path form”

•Noble Identity

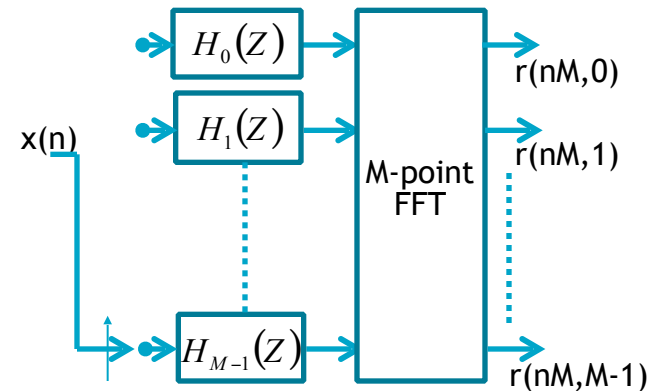
•Move a down-sampler back through a digital filter

$$H(Z^M) (\downarrow M) = (\downarrow M) H(Z)$$

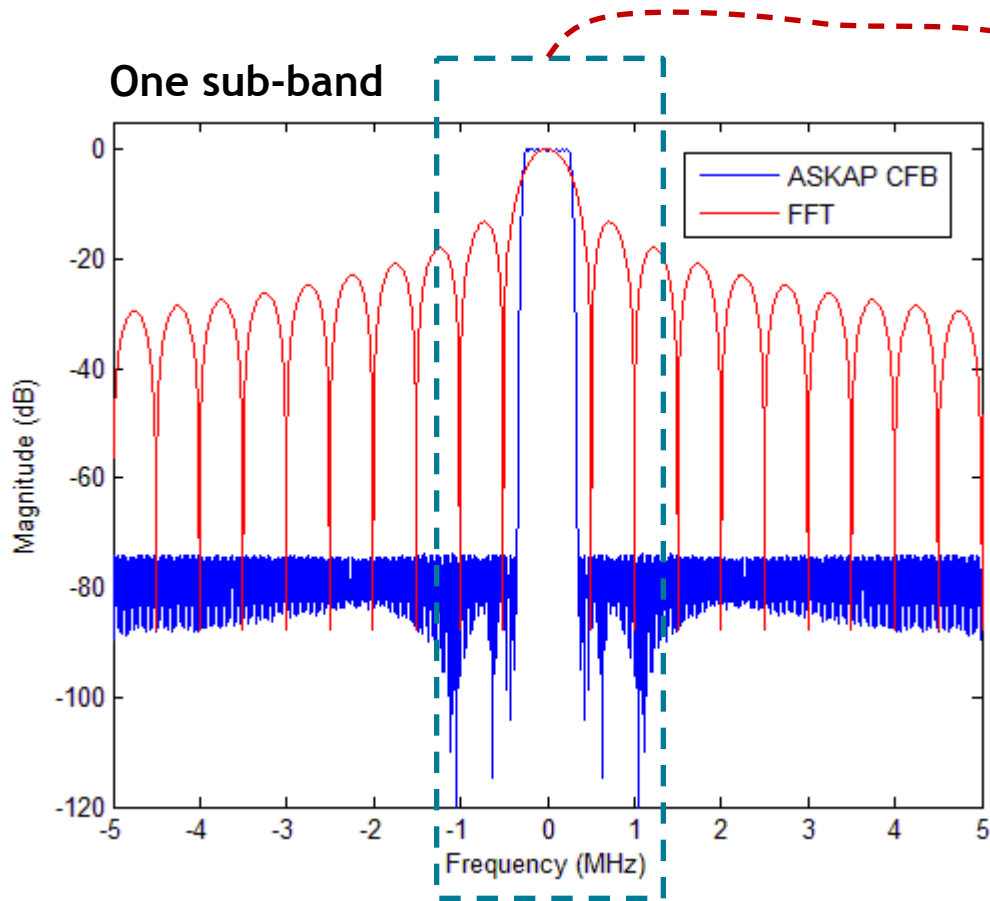
M-path Polyphase down converter



M-path Polyphase channeliser

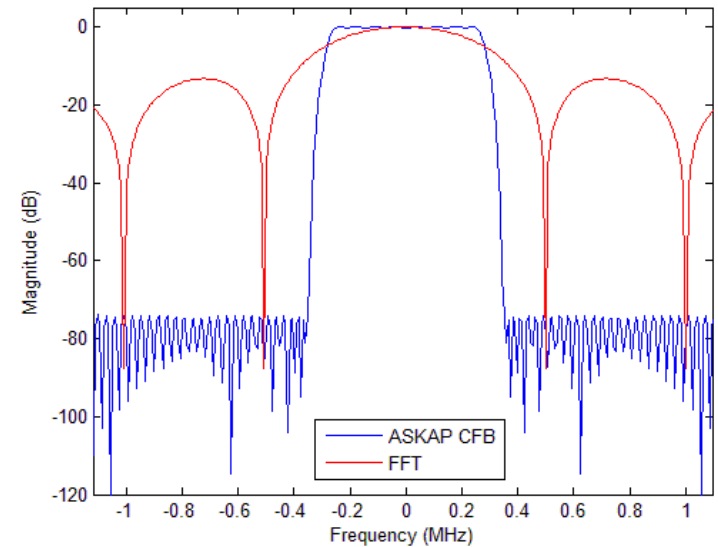


Filterbanks: FFT vs Polyphase Filters

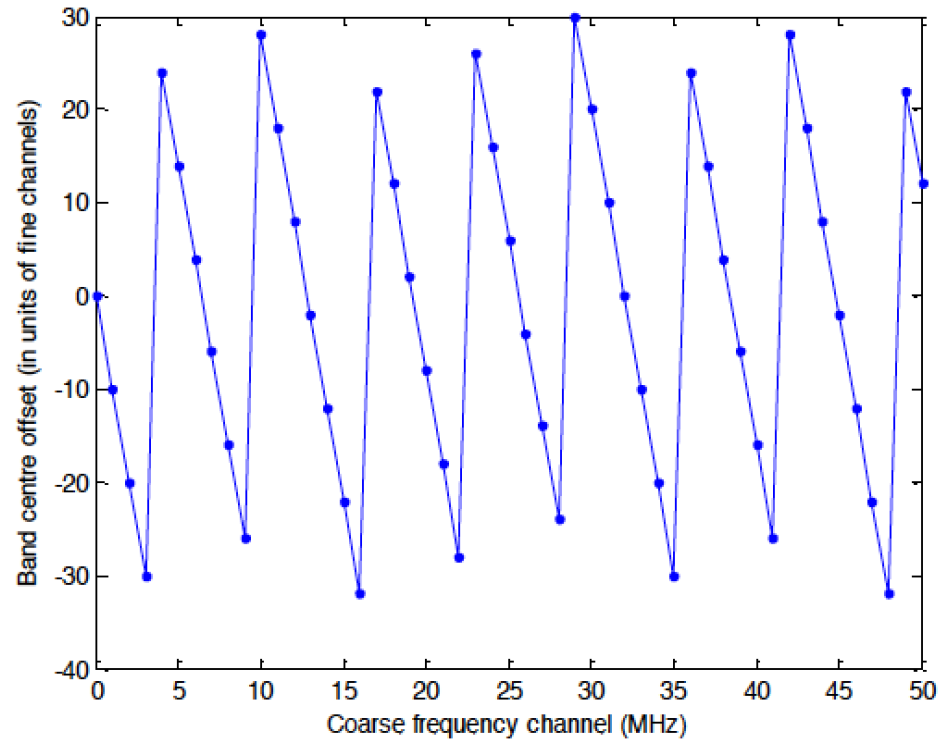
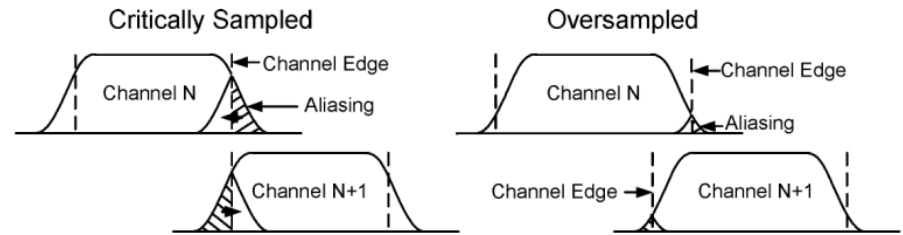
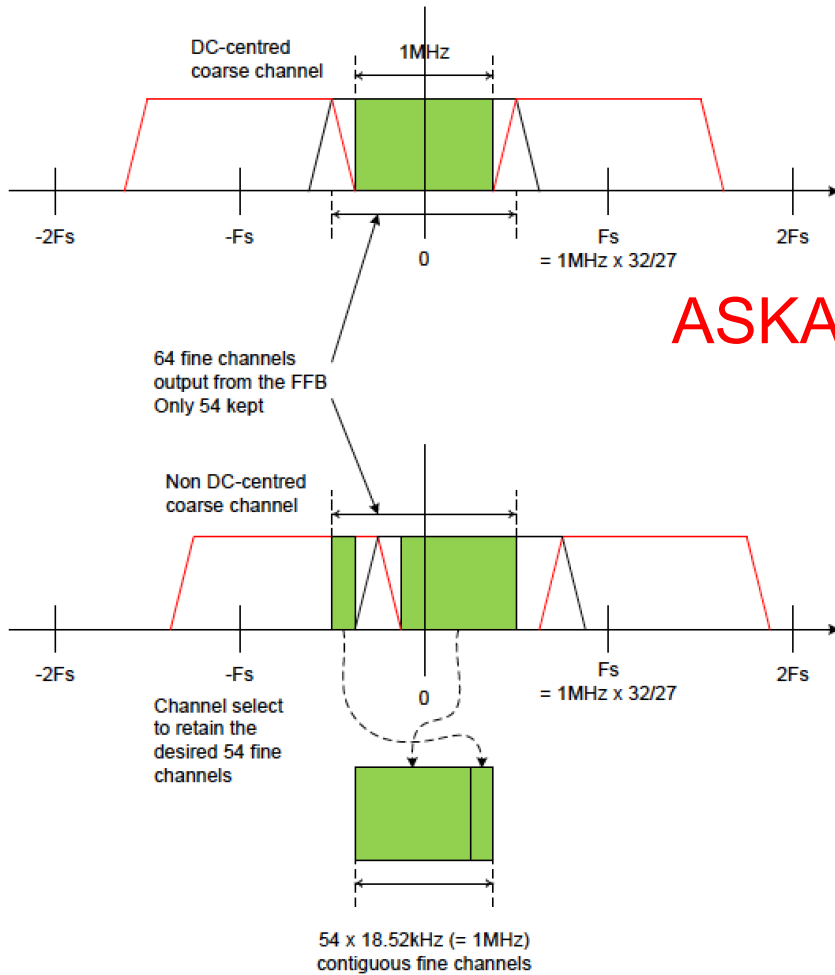


768-point FFT

12,288-tap polyphase filter + 768-point FFT



A few words on oversampled PFBs



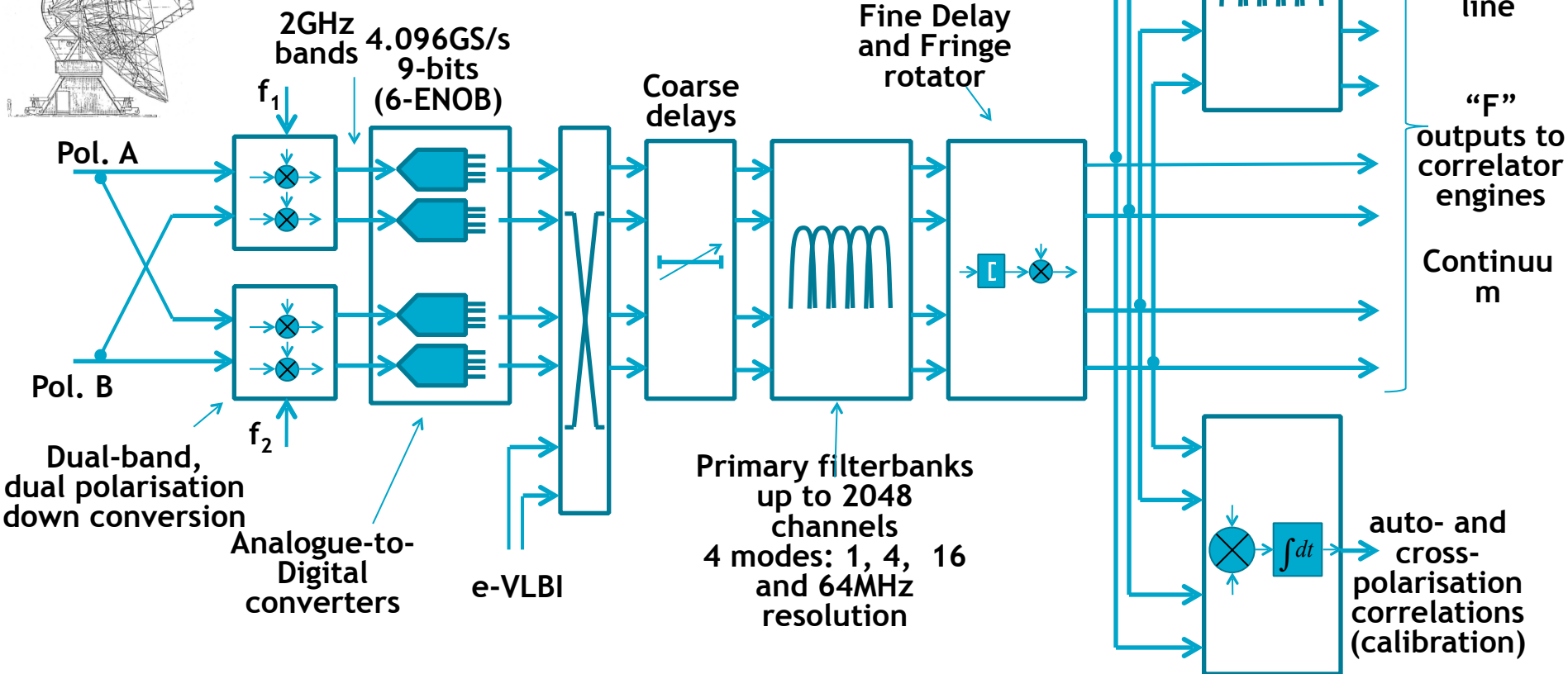
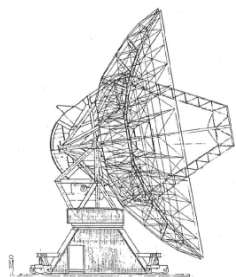
ATCA + CABB see Wilson et al. (2011, MNRAS, 416, 832)



Compact Array Broadband Backend (CABB)

Per antenna

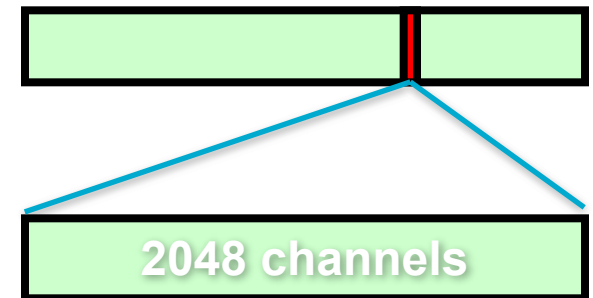
Secondary filterbanks
16 overlapping windows
2048 channels/window
(resolution depends on
primary filterbank mode)



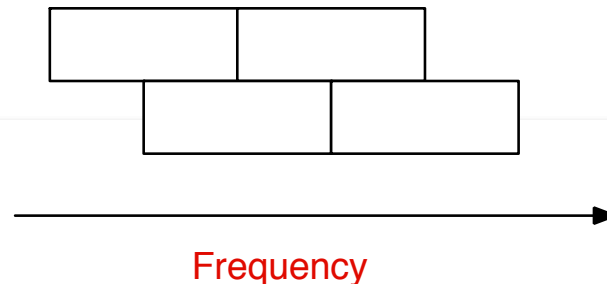
Zoom modes: high spectral resolution



- Up to 16 “zoomed in” channels per each 2 GHz window (user selected)
- Positioned in steps equal to half of the wide-band spectral resolution (i.e. 0.5 MHz or 32 MHz)
- Each zoom window has 2048 spectral channels

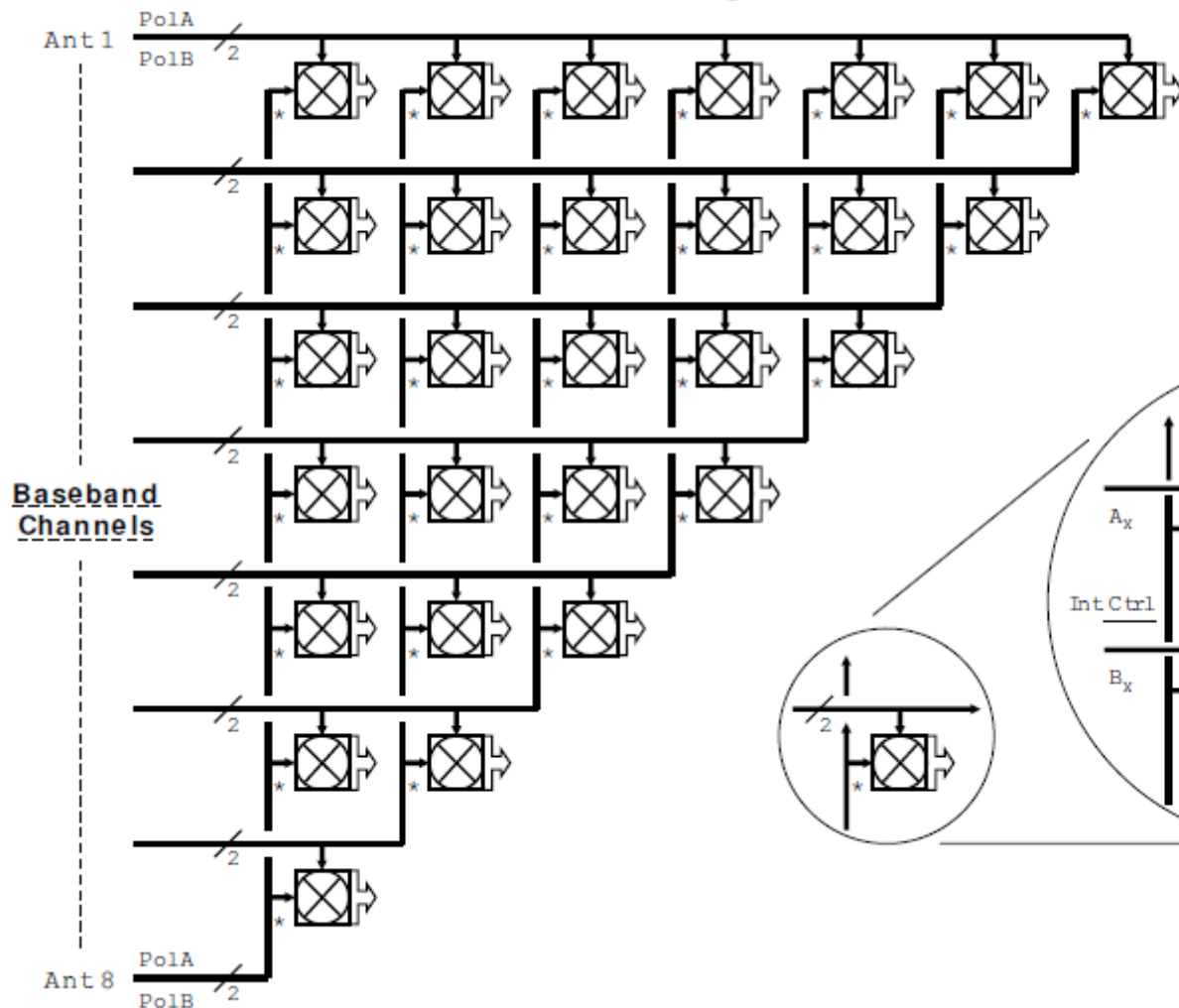


Stitching zoom windows:



CABB Correlator - correlation engines

Correlation Cell Array

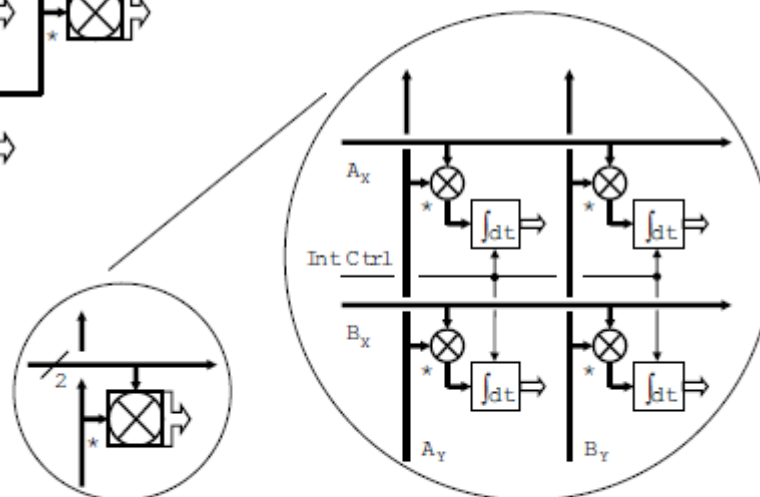


Integrated Visibilities

- $6 \times (6-1)/2 = 15$

baselines

- Full Stokes parameters



Correlation Cell

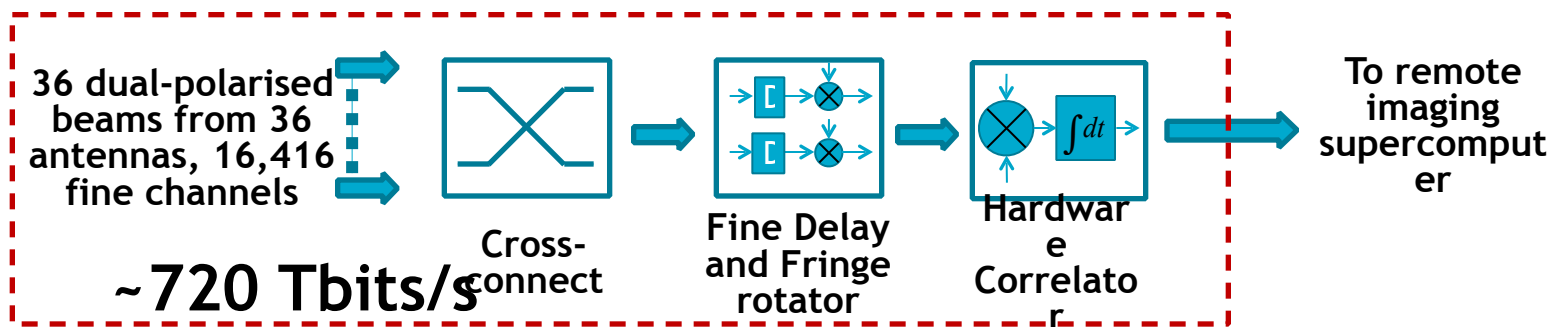
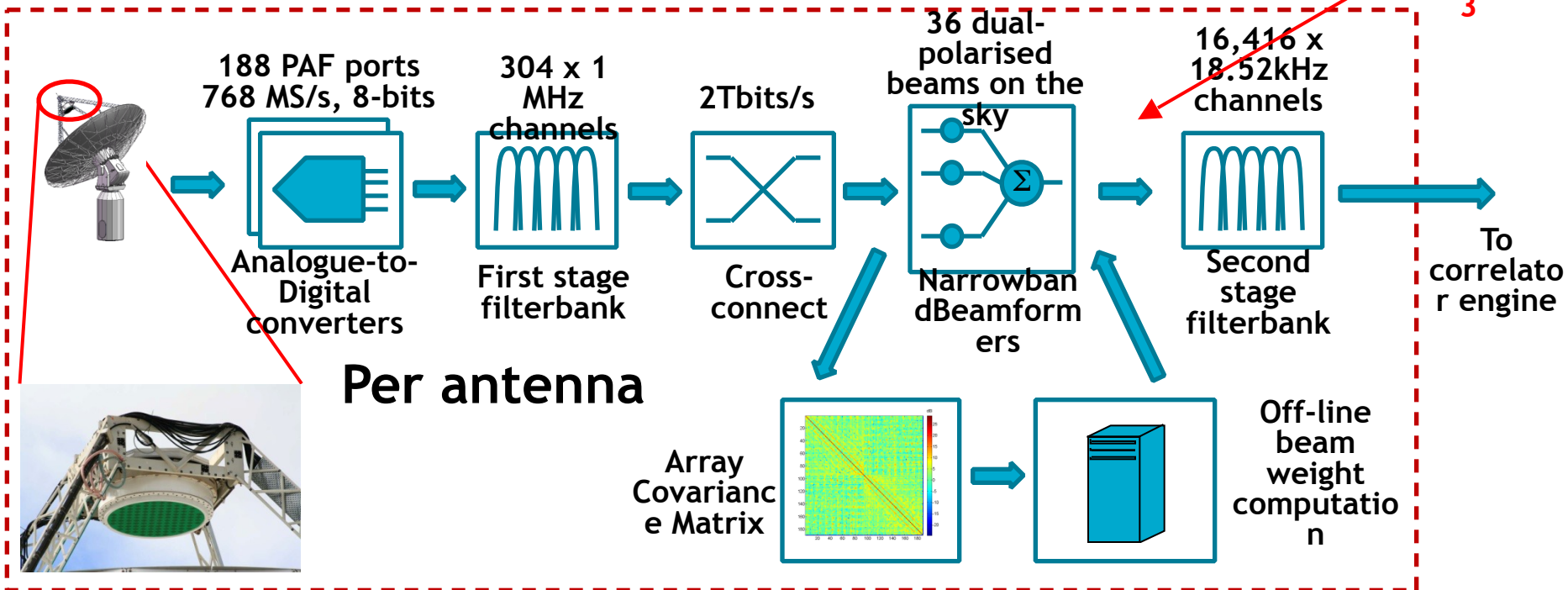
Australian Square Kilometre Array Pathfinder

- Has beamformer before the correlator (needs channelisation)



ASKAP (BETA) digital back-end

Data throughput reduced by a factor of 3

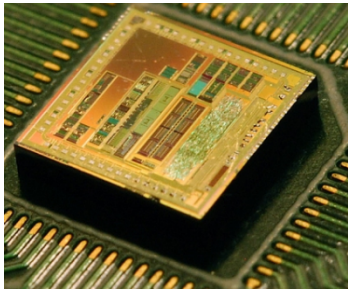


Calculation Engines: so many choices...

Hard-wired logic

Stored (programmed) logic

Application-Specific
Integrated Circuit



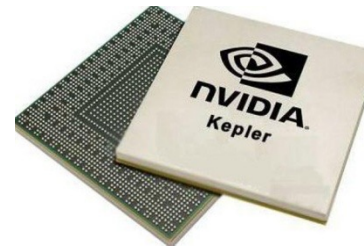
- EVLA
- ALMA

Field Programmable
Gate Array



- CABB
- ASKAP

Graphics
Processing Unit



- MWA
- MeerKAT

Central Processing
Unit/ Digital Signal
Processor



- DiFX



- More flexible
- Higher power/
computation
- Lower initial development

- Less flexible
- Lower power/
computation
- Higher initial development

Further Reading...

- Radio Astronomy:
 - H. C. Ko, “Coherence Theory in Radio-Astronomical Measurements,” *IEEE Trans. Antennas & Propagation*, pp. 10-20, Vol. AP-15, No. 1, Jan. 1967.
 - G. B. Taylor, G. L. Carilli and R. A. Perley, *Synthesis Imaging in Radio Astronomy II*, Astron. Soc. Pac. Conf. Series, vol. 180, 2008.
- CABB
 - W. E. Wilson, et. al. “The Australia Telescope Compact Array Broadband Backend (CABB): Description & First Results,” *Mon. Not. R. Astron. Soc.*, Feb. 2011
- ASKAP
 - D. R. DeBoer, et.al, “Australian SKA Pathfinder: A High-Dynamic Range Wide-Field of View Survey Telescope,” *Proc. IEEE*, 2009.
- Filter Banks
 - R. E. Crochiere and L. R. Rabiner *Multirate Digital Signal Processing*, Prentice Hall, 1983.
 - F. J. Harris, *Multirate Signal Processing for Communication Systems*, Prentice Hall, 2008.
 - P. P. Vaidyanathan, *Multirate Systems And Filter Banks*, Prentice Hall, 1992.
- Beamforming
 - B. D. Van Veen and K. M. Buckley, “Beamforming: A Versatile Approach to Spatial Filtering,” *IEEE ASSP Magazine*, April 1988

Thank you

Astronomy and Space Science

Max Voronkov

ASKAP Software Scientist

t +61 2 9372 4427

e maxim.voronkov@csiro.au

w www.narrabri.atnf.csiro.au/people/vor010

Astronomy and Space Science

www.csiro.au

