



Murriyang updates

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Australia's National Science Agency





Murriyang update and the wide-bandwidth receivers

Jane Kaczmarek

ATUC requests:

- 7.6: Details on the UWMH project itself, the properties of the receivers and the status and the timeline for the project will be presented at the next ATUC meeting.



Updates on & around site



New signage following
Murriyang naming guidelines



On-site accommodation trial
for ATNF staff & affiliates



Maintenance Projects

- Successful refurbishment of azimuth gearboxes & roller bearings in June 2024
 - Once-every-20-years job!
- Zenith gearbox refurb & modification planned for 2025APR semester





Scheduling Constraints

- Remainder of 2024 OCT semester to be released soon
 - Subject to CryoPAF installation timeline
- Spacecraft tracking commitments continue to have a fluid schedule
 - ~2-week support anticipated early 2025
- Month-long gearbox refurbishment has been shifted to the 2025 APR semester





The Ultra Wide-bandwidth Mid & High (UWM-H)

Mid

4160 – 15680 MHz

High

15680 – 27072 MHz

- Two distinct receivers
- Will share the UWL platform & warm electronics
 - Removes all foreseeable receiver changes



Proposed Ultra Wide-bandwidth System

- Jimbles for the entire Ultra Wide package
 - UWL adjusted frequency of 704 - 4288 MHz
 - Oversampled filterbank for all receivers
- Digitised calibration signal
 - Can be completely removed from sky signal
 - Gain & phase calibration without pulsed signal
- GPU backend
- ST data available as digital or analogue
 - Non-commensal with astronomy modes

Low

704 – 4288 MHz

Mid

4160 – 15680 MHz

High

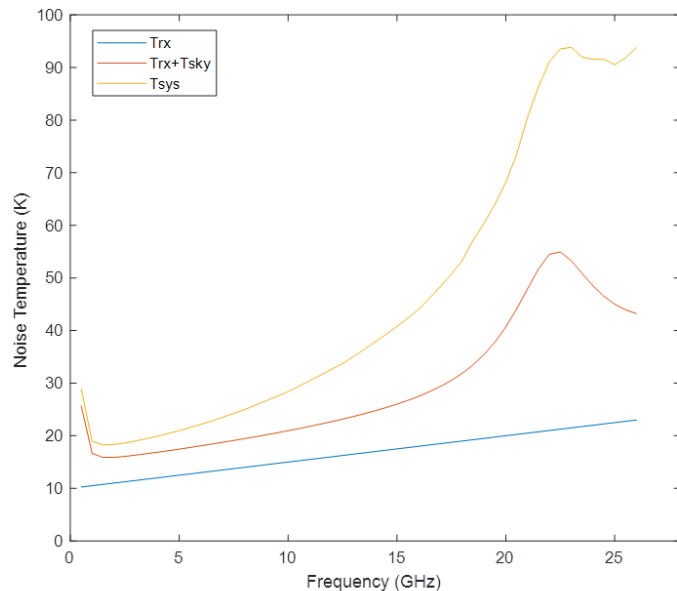
15680 – 27072 MHz

ST

< 32300 MHz



Expected System Performance



Credit: Alex Dunning

- Spillover becomes noticeable at higher frequencies
- Improved (or equivalent) performance to existing fleet:

Receiver	Frequency Range (GHz)	Average Tsys (K)
METH6	6.3 - 6.8	58
MARS	8 – 9 GHz	~25
13MM	20 – 23.5	80

Anticipated Timeline

- Project is planned to run for three years
 - Potential to start Jan 2025
 - Installation late 2027
- Will require UWL to be offline for 8-10 weeks
- Timeline is subject to SKA tender (see talk by Mark)
 - Start time will be known by the end of the year





CryoPAF update

Simon Johnston



CryoPAF General Specifications

Frequency range of 700-1950 MHz in two bands: 700-1100 MHz and 1100-1950 MHz

Processed bandwidth of 600 MHz, with potential to expand to 900 MHz.

$T_{\text{sys}} < 20\text{K}$, $S_{\text{sys}} \sim 25 \text{ Jy}$

Phased array feed with 98 dual linear polarized elements

Maximum of 72 beams (8 for pulsar timing and VLBI). FoV approx 2 sq deg.

Rotation allows tracking in parallactic angle





The good, the less good and the future

Good news

- On dish since end August
- Excellent Tsys performance
- Cryogenic issues seem largely to be resolved (fingers crossed)
- Voltage mode (8 beams) largely working with some caveats
- 9 beam-formed beams created at 1300 MHz with 240 MHz bandwidth.
- Progress with TOS/Garriwang

Less good news

- Project delayed 2 years in the last 2 years
- LNA bias cards need replacing
- Pulsar search mode under testing
- Spectrometer mode under testing
- Beamforming still to be demonstrated across the entire band (RFI)
- Black belt operation/observing
- Transient buffer and real time transient detection mode only started
- SETI mode not started

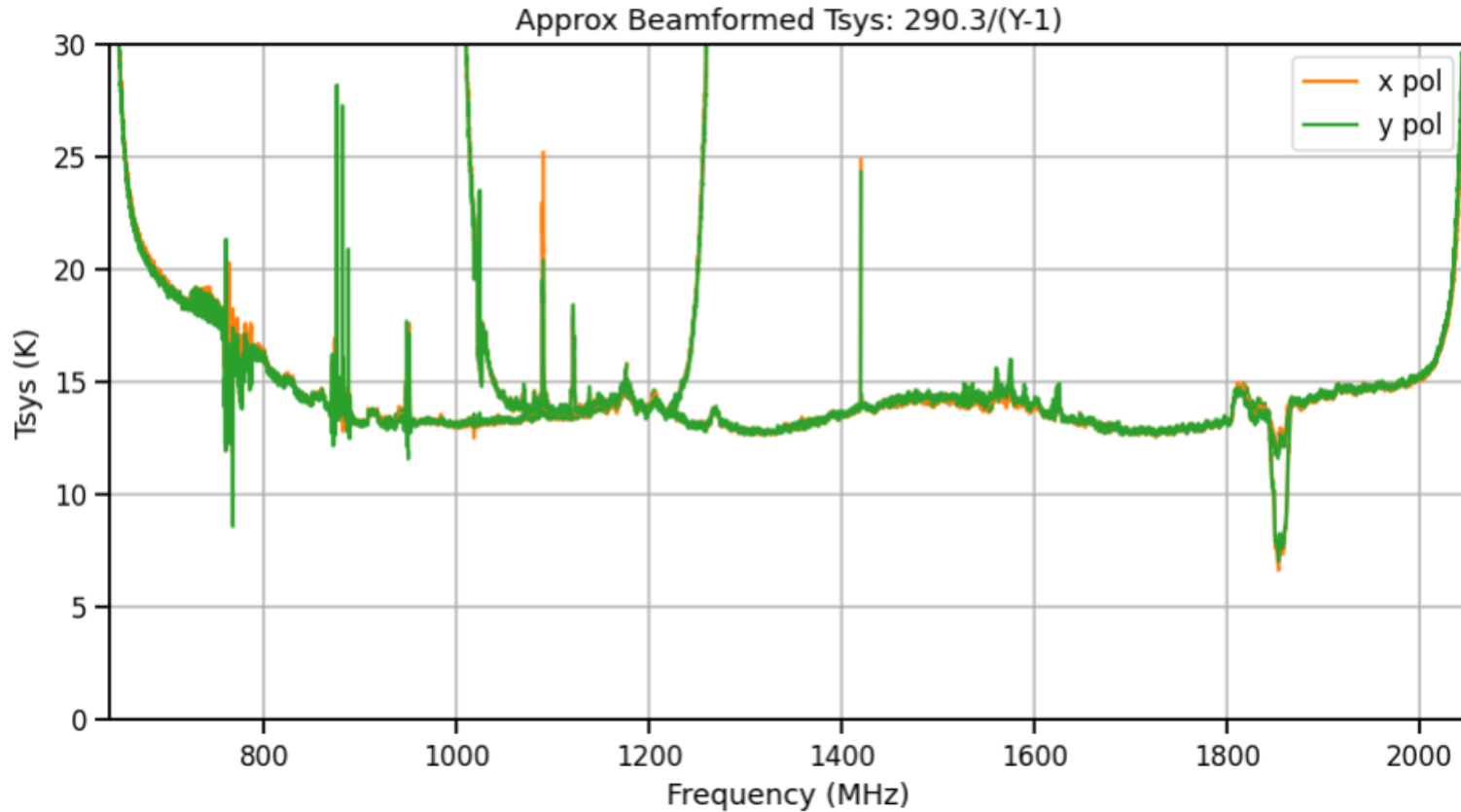


Hot box testing – mid August



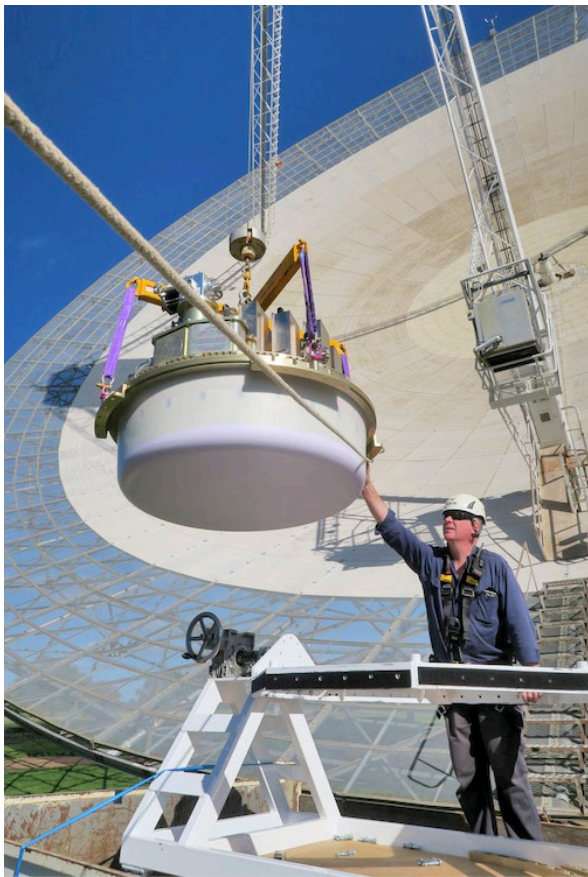


Hot box testing - Results

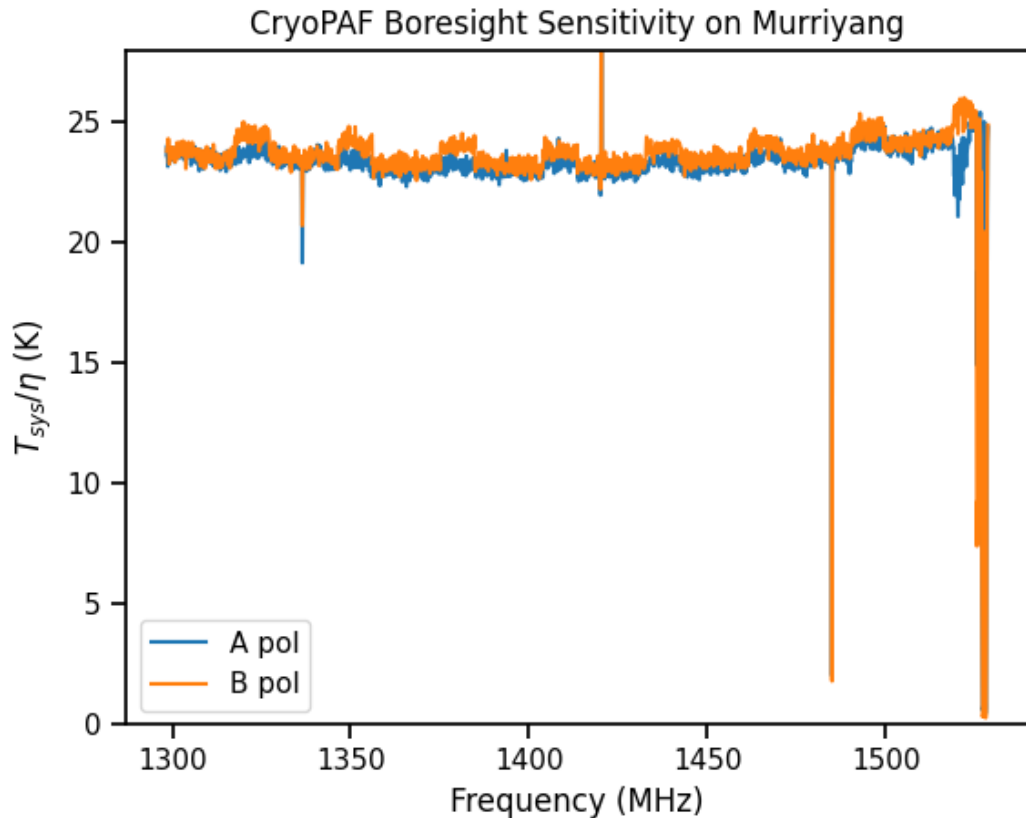




Installation on the Dish – end August



First Results – on dish testing

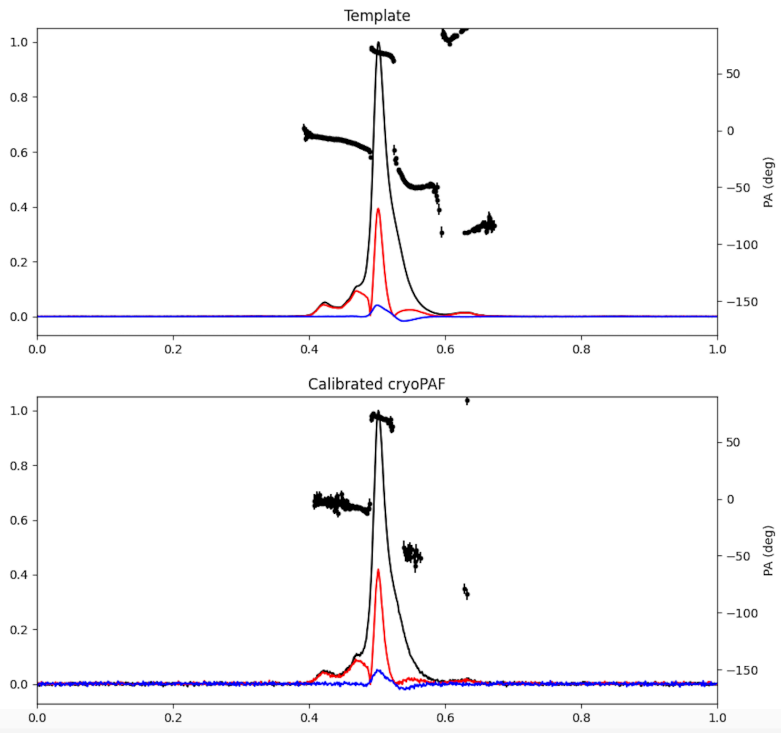


Preliminary results at 1300 MHz with 240 MHz of bandwidth.

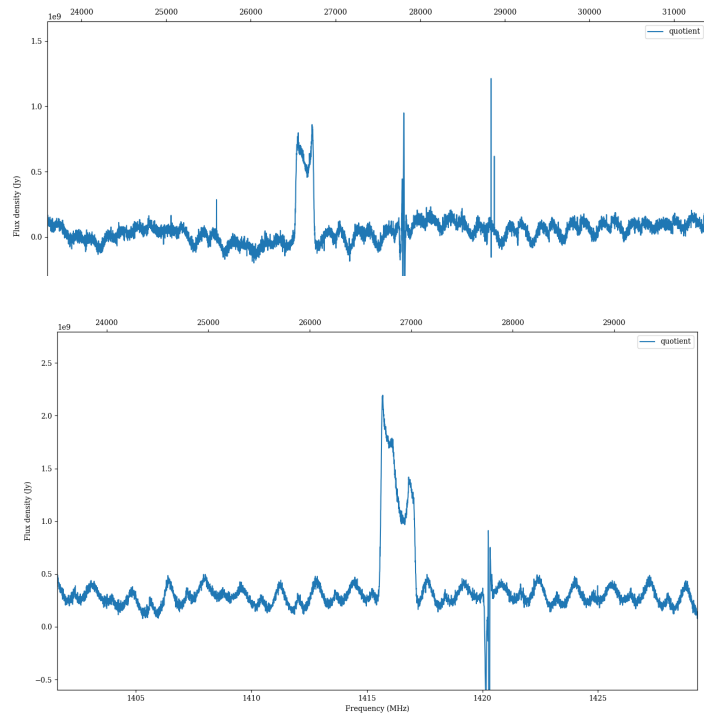
Some artifacts but overall excellent performance and a bit below the spec of 27 K.



First Results – on dish testing



Polarization of the pulsar J0437-4715. Voltage/fold mode



HI spectrum from NGCs 6744 and 4592. Voltage mode.



The good, the less good and the future

Future

- The CryoPAF needs to return to Marsfield to fix the WEM/LNA issue
 - Dec/Jan or Mar/Apr still to be decided
- Work continues on :
 - Quantifying system performance
 - Getting science modes operational
 - Making beam weights easier to compute / upload
 - System integration into TOS / Gariwang
 - Simplifying taking data
- Observing proposals already submitted will go forward into next semester
 - Resubmission of cover page / targets required



The pulsar search pipeline

Shi Dai/Lawrence Toomey

ATUC requests:

- 4.2: Report on the feasibility of providing pulsar searching as a service for the Murriyang community



Pulsar searching as a service for the Murriyang community

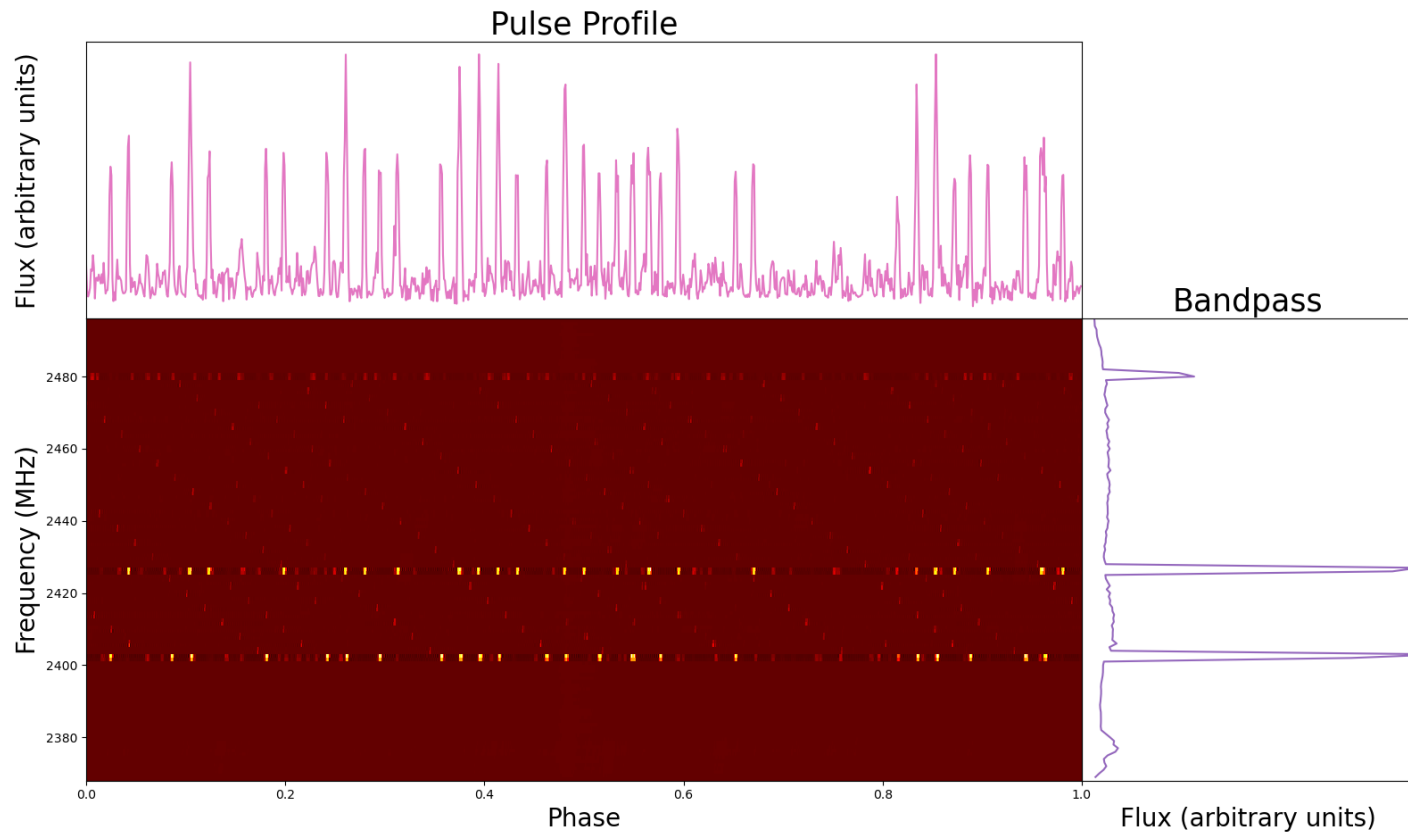
- We will be trialling an end-to-end pulsar search pipeline and a real-time FRB detection mode as part of CryoPAF commissioning – the trial has not yet started
- Two options for users:
 - The pipeline service will likely be a “one-shot” process – the user will receive a candidate list, and raw data will be deleted
 - The pipeline itself is available as a container, however the user will need to allocate their own computing/storage resources



Removing RFI from Parkes data

Tommy Marshman

Parkes Default Settings (No Spectral Kustosis)



Parkes Default Spectral Kurtosis Settings

