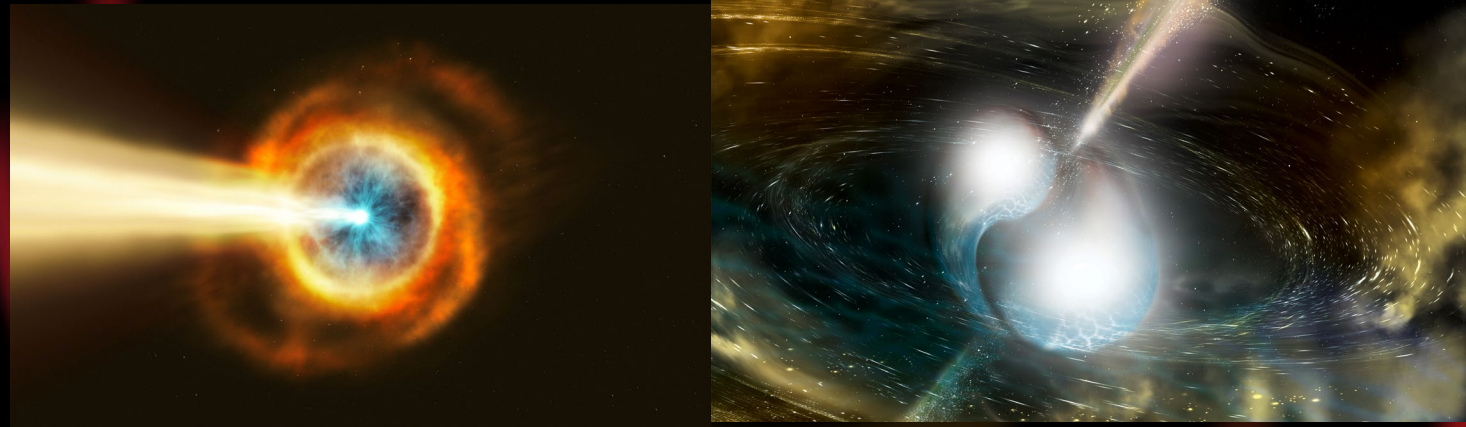




International
Centre for
Radio
Astronomy
Research

Rapid-response triggering on transients with ATCA



Gemma Anderson

ICRAR- Curtin University

ATCA Science Day, 8 April 2025



(1) GRB / flare star /
XRB

RAPID-RESPONSE RADIO TELESCOPES

(2) Swift Burst Alert
Telescope

γ -rays

(3) Position transmitted

Radio Afterglow
1-20 GHz

Radio FRB
80-300 MHz

MWA

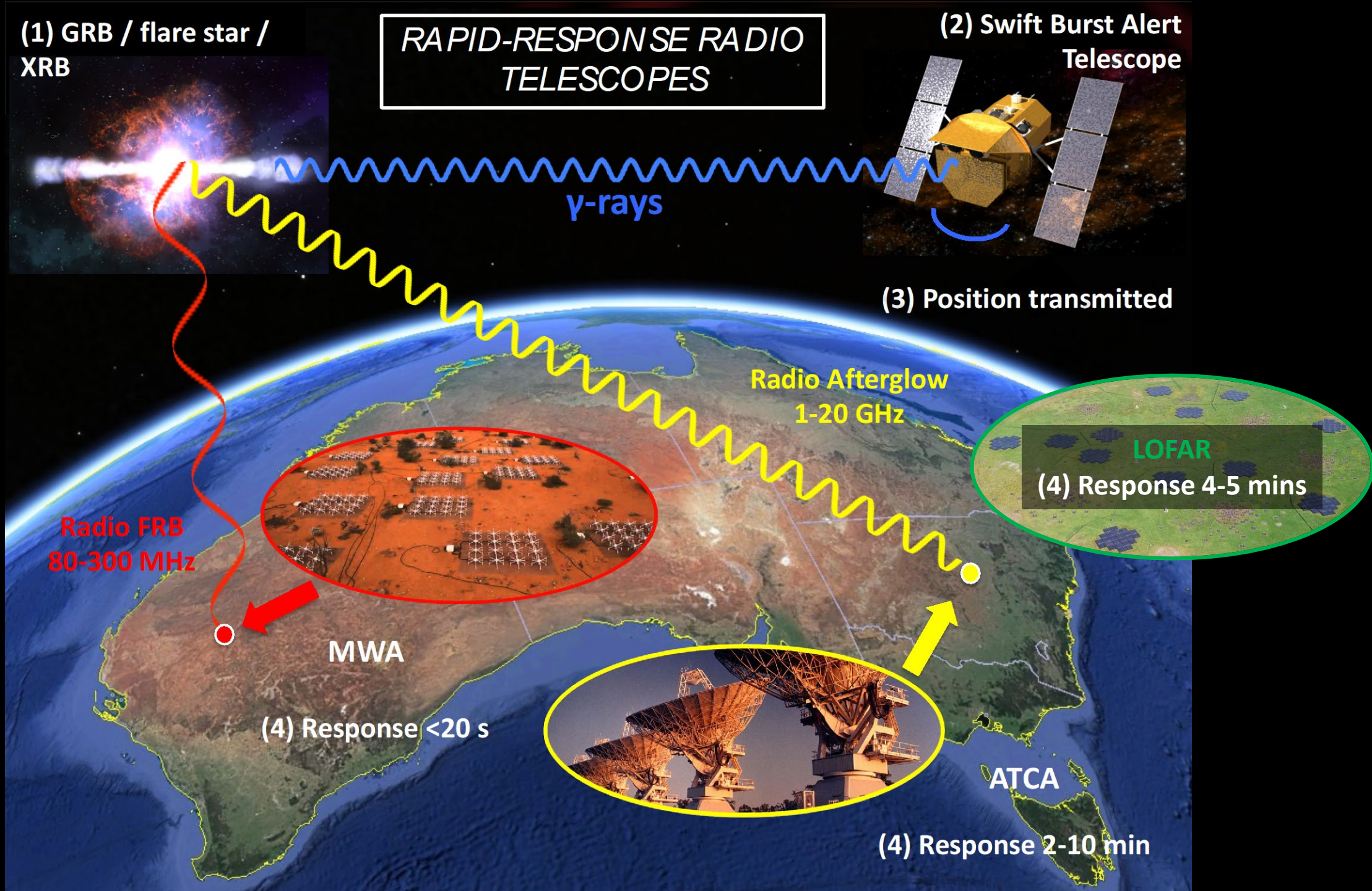
(4) Response <20 s

LOFAR

(4) Response 4-5 mins

ATCA

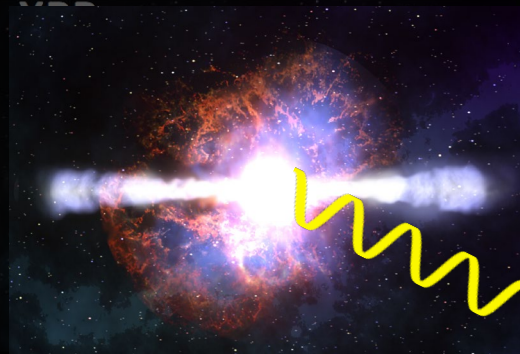
(4) Response 2-10 min



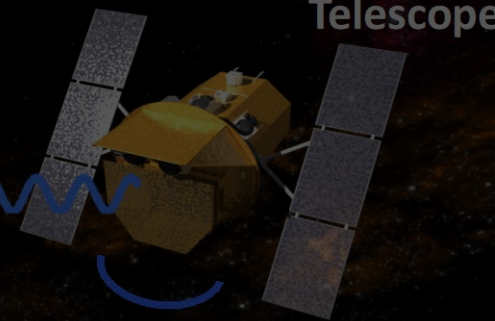
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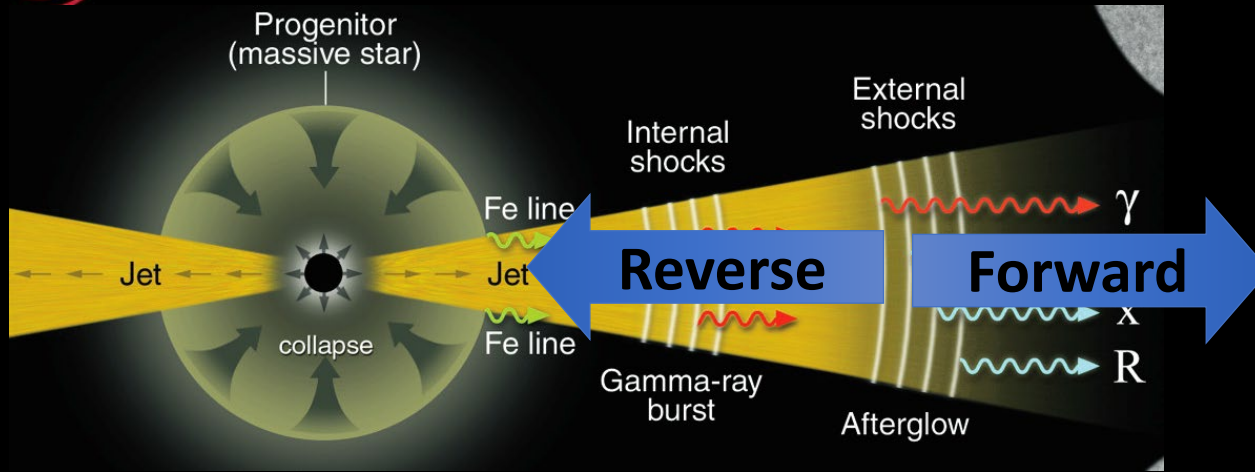
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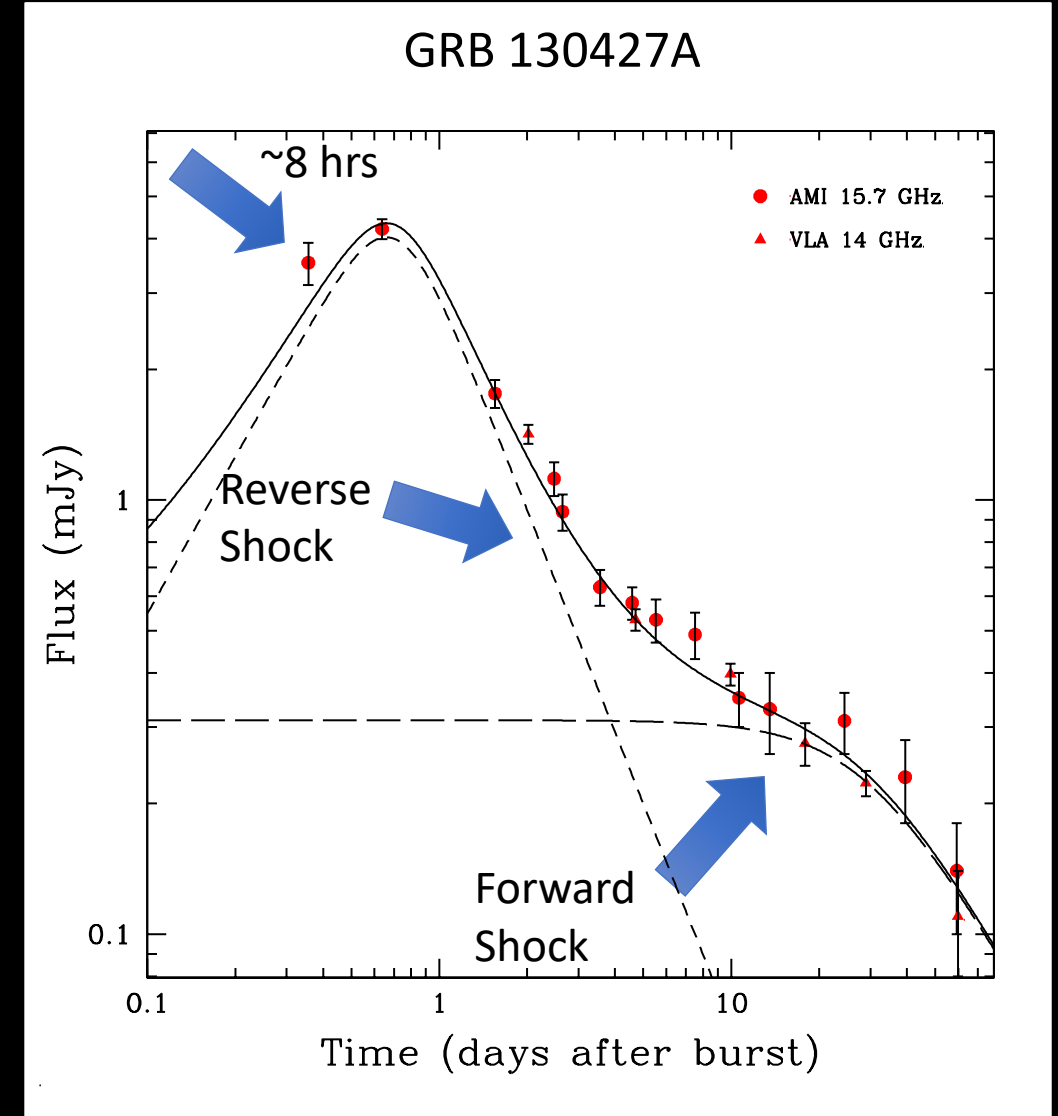
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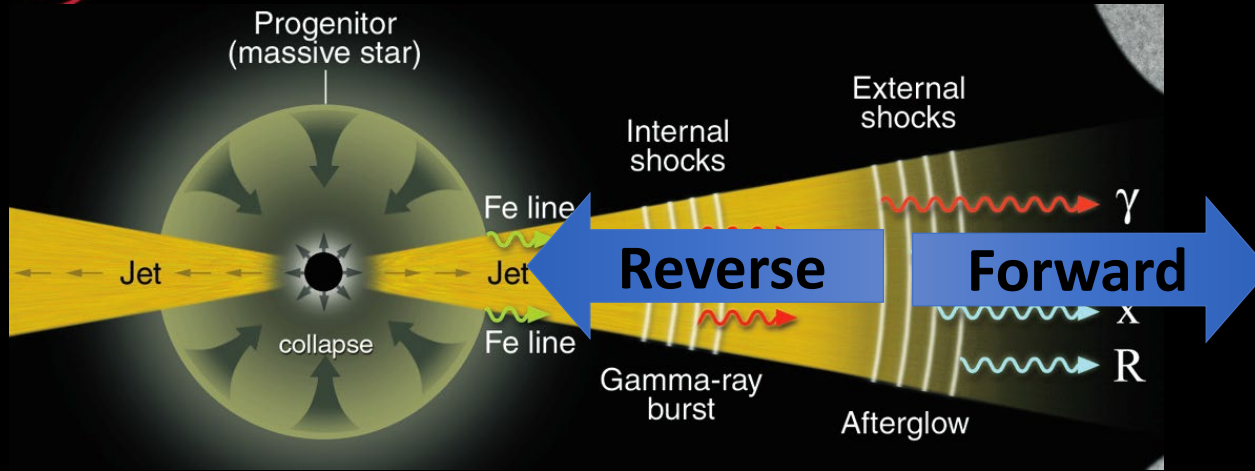
Science Cases for rapid-response



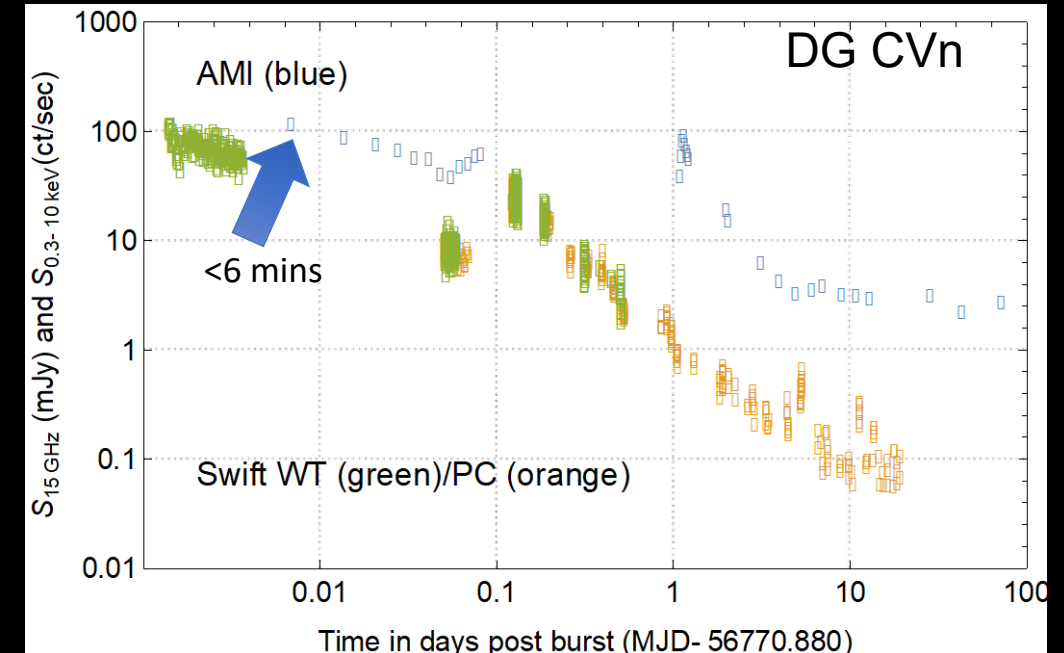
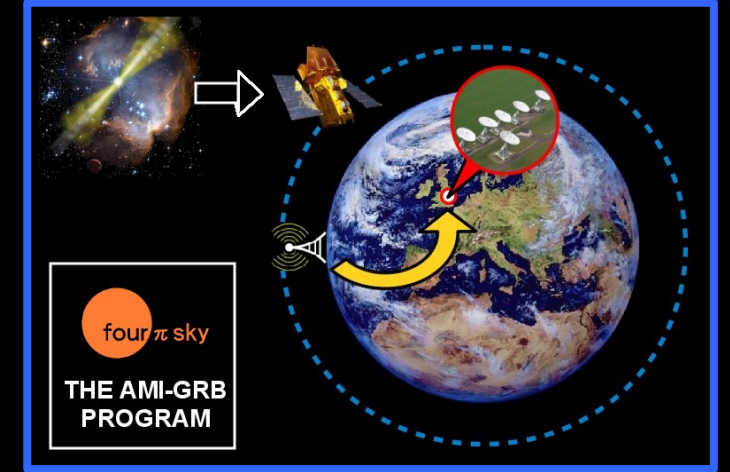
- Gamma-ray Bursts (GRBs)
 - Reverse shock – probe outflow composition
 - Central engine – energy injection, magnetar
- Flare stars
 - Gyrosynchrotron emission
 - Coherent emission mechanisms
- X-ray binaries – L_X - L_R evolution
- LSST transients
 - Supernova shock break-outs?
- Multi-messenger transients – LVK, CTA



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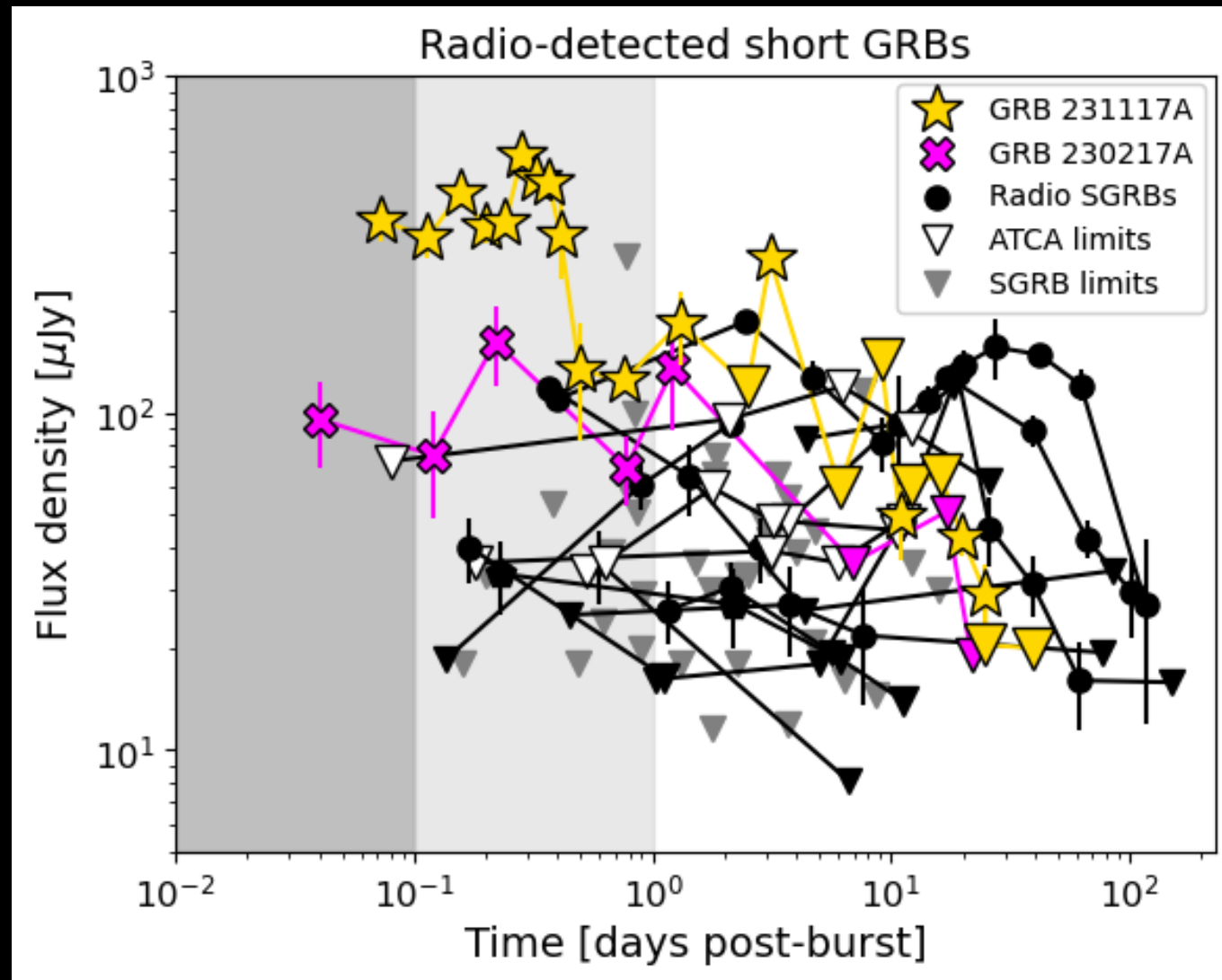


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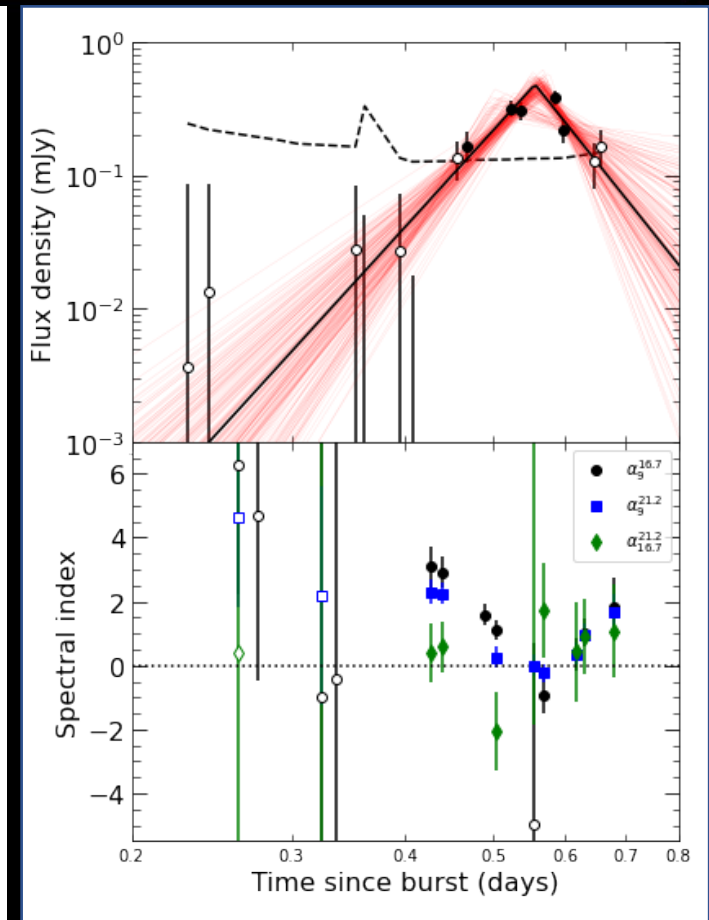
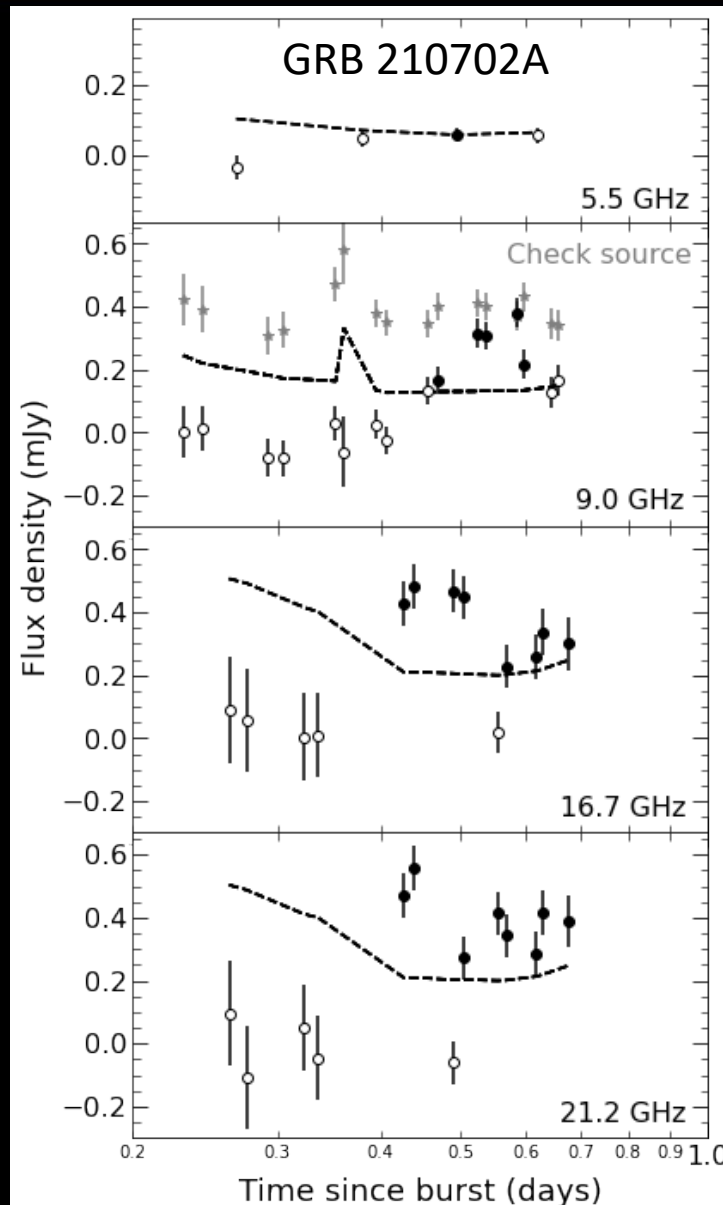
ATCA's Unique Capability

1. Only rapid-response telescope at GHz frequencies
 - Slew time 1-10 mins
 - Phase cal observation
 - Its **interruptibility**



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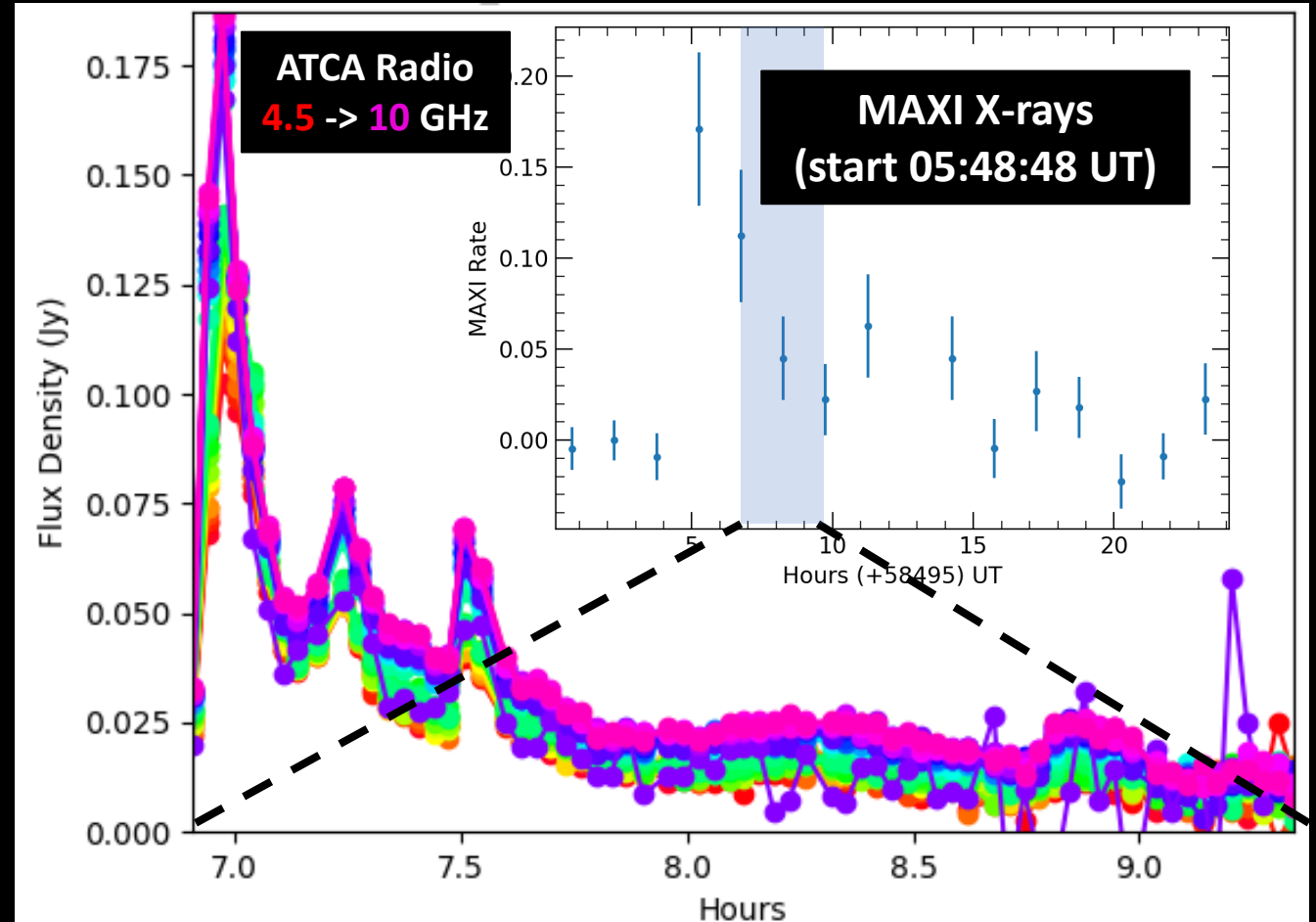
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AT Mic: ATCA trigger on MAXI-detected flare



Credit: Daniele d'Antonio



ATCA's Unique Capability

GRB 240205B

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2. Large instantaneous bandwidth:
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3. Designed for longer integrations
 - East-West array
 - Can justify up to 12 hours
 - VLA: ~3 hour triggers

ATCA in the SKA era

- Rapid-response unique at GHz frequencies until SKA-Mid
 - First steps to fully automated observing
- A dedicated transient machine
 - Transient SKA-Mid dedicated subarray?
 - Could ATCA do this?
 - SKA transient follow-up?
 - Rapid-response
 - Long term multi-frequency follow-up
 - A second North-South spar at 2 or 3 km
 - Rounder beam
 - Shorter integrations for transient snap shots

