



Spectral line science with ATCA

Max Voronkov | ATCA Science day – 08 April 2025

Australia's National Science Agency



Lots of spectral lines beyond H I

(Quasi)thermal



Harder with ATCA used as
an interferometer

Masers



qualitative

Estimate of physical parameters (e.g. T, density), abundance

Astro-chemistry and tracers of shocked gas and interactions

Kinematics (including kinematic distance), morphology

Magnetic fields & Zeeman splitting (OH in particular)

Recombination lines

Ionisation, electron temperature and density, metallicity

Kinematics of HII regions and their interaction

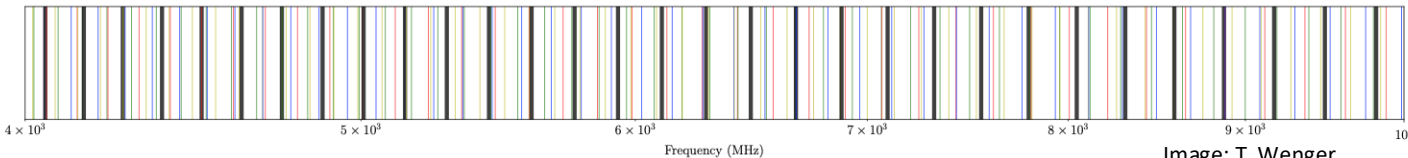
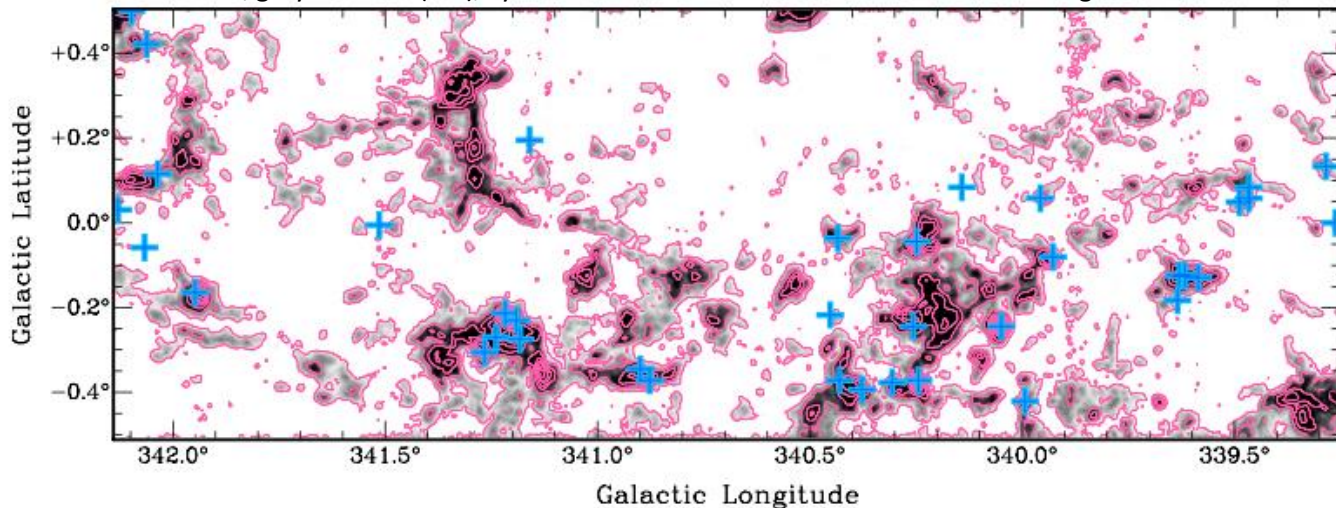


Image: T. Wenger

StarFISH: ATCA as Mopra++

Star Formation In the Southern Hemisphere (PI: S.Breen)

Contours/greyscale: CS (1-0), symbols – class I methanol masers at 44 GHz. Image credit: S.Breen



Transition	Frequency, GHz	Transition	Frequency, GHz
SiO (1-0) $\nu=3$	42.51934	H51 α (RRL)	48.15360
SiO (1-0) $\nu=2$	42.82048	C ³⁴ S (1-0)	48.20694
H53 α (RRL)	42.95197	CH ₃ OH 1 ₀ -0 ₀ A ⁺	48.37246
SiO (1-0) $\nu=1$	43.12203	CH ₃ OH 1 ₀ -0 ₀ E	48.37689
SiO (1-0) $\nu=0$	43.42376	OCS (4-3)	48.65160
CH₃OH 7₀-6₁A⁺	44.06941	CS (1-0)	48.99095

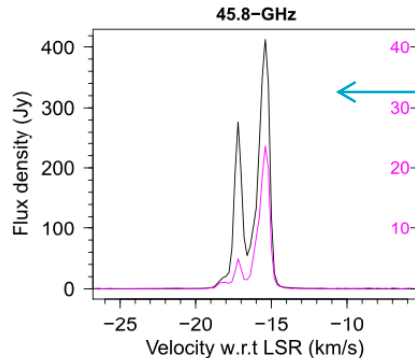
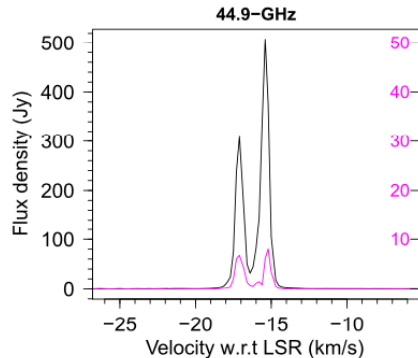
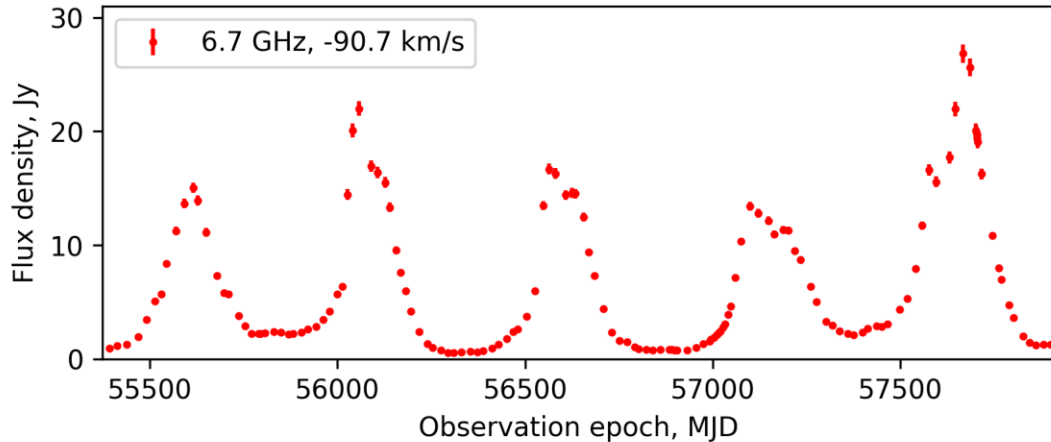
ATCA is used as
6 single dishes

Periodic variability and bursts in masers

About 30 (quasi)periodic masers known (largely class II methanol)

Example from
my own
monitoring work →
with ATCA

G331.13-0.24



Flare in G358.931-0.030
(Breen et al., 2019, ApJ, 876,
25)

Lots of unusual
transitions!

Benefits of the new system

- More user-friendly / (hopefully) less bugs
- Better spectral resolution options, especially for 12mm
- Potentially different spectral resolution for masers and thermal lines and RRLs
- More flexible positioning of zooms, better default resolution (i.e. no zooms needed for high frequency)
- Main benefits are with the RF upgrade and 8 GHz
 - More stuff observable simultaneously
 - Can finally get 12.2 GHz methanol line (surveys, VLBI?, monitoring together with 6.7 GHz)

Wishlist

- Simultaneous observations at 36 and 44 GHz
- Mopra 3mm receivers at ATCA antennas



Thank you

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