

ATUC ASKAP Update

Aidan Hotan | ASKAP project scientist April 2020

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ASKAP observatory update

- Operations continue with contingency plans for COVID-19
 - Fast-track pilot survey observations underway
- RACS quality control leads to greater understanding of beams
 - Predictable flux correction required across FoV, primarily for outer beams
- Full steam ahead on analysis of existing pilot data
 - Data processing team working to meet requirements and upload to CASDA
- Operational support and project clashes have stretched resources
 - ASKAP-X on hold, interim operations coordination structure in place





Full ASKAP + Parkes SMC cube

- Test data in 16x zoom mode
- Imaged in CASA on OzStar
 - Pipeline developed by Nick Pingel
- Operational support for joint imaging will be challenging
 - Seeking additional resources
 - Investigating alternative software
 - Pawsey hardware refresh
 - YandaSoft portability options



Fast-track pilot survey storage space

- COVID-19 increases risk of observatory closure
- Pawsey more likely to remain operational than MRO
- Need pilot data on disk even if we can't keep up with processing
- Pawsey have agreed to provide tape storage as a data buffer
- Commissioning underway now





Fast-track pilot survey observing strategy

- Remaining observations require about one month of time
 - Assuming all goes smoothly, could be more given historical efficiency
- Lack of configuration agility means one survey at a time
 - LST constraints create gaps that we fill with CRAFT FRB searching
- Observing requires draining data to tape as fast as it arrives
 - Testing needed to determine whether this is possible
- Processing on /askapbuffer can interfere with data ingest
 - Halt on I/O intensive activities while observing
 - Some processing and development work can continue on /group



EMU pilot survey

- Discovery of Odd Radio Circles
 - No optical or IR counterparts
 - Roughly 1 arcminute in diameter
 - Not SNRs, spiral galaxies
 - Most likely not artefacts
- Leading theory is spherical shock waves from unidentified transient events

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WALLABY pilot survey

- Hydra field awaiting quality control & release on CASDA
- SoFiA source detection software identifying many new galaxies in the cluster
- Extensive quality analysis report provided to CASS
- Two more fields observed



Right Ascension (J2000)



Radio loud stars in the VAST pilot survey



Image by Joshua Pritchard



Fast imaging pilot search on GW field (GW170814bv)

- Observed 5 times for GW follow-up pilot
 - Observations at 2, 9, 33, 86, 233 days after event
 - Several auxiliary science possibilities
- 15-minute imaging (no clean) after model subtraction
 - Identified 6 'fast' variable sources
 - Likely associated with interstellar scintillation



POSSUM pilot survey

- Stokes I image from pilot field SBID 10168, and RM distribution
 - RM catalogue clipped to keep only reliable sources
 - 16 RMs per square degree, factor of 10 improvement



Pilot survey review and consolidation

- Work is needed to make ASKAP survey-ready
 - Pilots have produced good data, but highlighted operational issues
 - Some planned improvements need array time for testing and development
 - Consolidation required once current observations finish
- We seek feedback from pilot survey data analysis
 - Collect and prioritise improvements needed for full surveys
- Priority shifts back to engineering before the next pilot surveys
 - Still possible to conduct time-critical observations and tests

Rapid ASKAP Continuum Survey status report

- Last ATUC reported good progress and imminent data release
 - Quality control has since found issues that have delayed release
 - We are learning a lot about ASKAP as part of this process!
- Astrometry checks pass with precision and accuracy < 1"
- Photometry revealed some issues we need to solve
 - Flux densities systematically high by 5-10%
 - Source spectra steeper than expected

RACS source flux investigation

- ASKAP mosaics assume circular gaussian beams, FWHM from holography
- New holography measurements were made of the RACS footprint
- These confirm significant but systematic departure from circularity in outer beams
 - Can be corrected in the image plane until position-dependent beam models are made



RACS data quality and PSF variability

- PSF variability is present, both within and between tiles
 - Caused by early observations being taken away from the meridian, and some differential flagging
- Methods developed to characterise PSF changes over each tile
 - A selection of badly affected fields have been re-observed
- These recent data show great improvement in telescope performance since early last year





Image by David McConnell

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SWAG-X: ASKAP eFEDS project

- MoU between AAL and the eROSITA-DE collaboration
 - ASKAP is observing the GAMA-09 field, alongside eROSITA performance verification
 - Continuum test at 888 MHz covering the full area complete (release pending)
 - One 888 MHz spectral epoch observed and awaiting processing (behind pilots)
- Widespread interest in GAMA-09 from ASKAP SSTs
 - Obtain a wide range of data products as an observatory project
- Provide a public resource for multi-wavelength comparison





SWAG-X: ASKAP eFEDS project

Image by Vanessa Moss eROSITA footprint in green





Suspension of ASKAP-X

- ASKAP is a new instrument and fully operated by CASS
 - Routine operations frequently cross into engineering domains
 - No operations structure existed to coordinate activities
- The ASKAP-X project was created to implement fixes and features
 - Operational support was always out of scope, but taking most of people's time
- With the need to complete pilots, we have suspended ASKAP-X
 - Focus on operational support during this difficult time
 - Develop a viable operations model for surveys in future
 - Manage "funded" features like CRAFT coherent search mode as projects



Creating a workable operations model

- We need an operations structure crossing the observatory site, science, development, engineering and computing domains
 - All highly distributed and often de-centralised teams
- ASKAP-X SAFe procedures weren't always working in our situation
 - Overheads too high when teams are fractional
 - Skills isolated to individuals, making team focus difficult to achieve
- Need to keep the cross-team coordination we developed
 - Team representatives, structured test time, release planning, etc.



The path to beginning full surveys

- Priority is currently on finishing the existing pilot survey plans
- The first pilot surveys have been very instructive
 - We are planning a second round that will focus on commensality
- Pilot survey experience will inform a Combined Survey Strategy
 - This should optimise the science/time ratio for the full survey projects
 - Must clearly identify resource requirements for each survey



ASKAP survey project review

- The combined survey strategy will be the observatory's recommended approach to maximise efficiency
 - It will be developed in consultation with the science teams
- An external review of the strategy will be arranged
- As a fall-back, the review panel will rank updated science cases
 - This can be used as a tie-breaker when survey strategies conflict



Conclusions

- Pilot surveys are progressing and producing exciting results
- RACS is teaching us a lot about the telescope
 - It should still be the world's benchmark survey at 1 GHz
- Operating ASKAP requires new ways of thinking for CASS
 - We are navigating this and iterating on an optimal operations model
- Science teams encouraged to discuss commensality opportunities

