



BIGCAT Update

Chris Phillips | BIGCAT Project Lead

8th November 2021

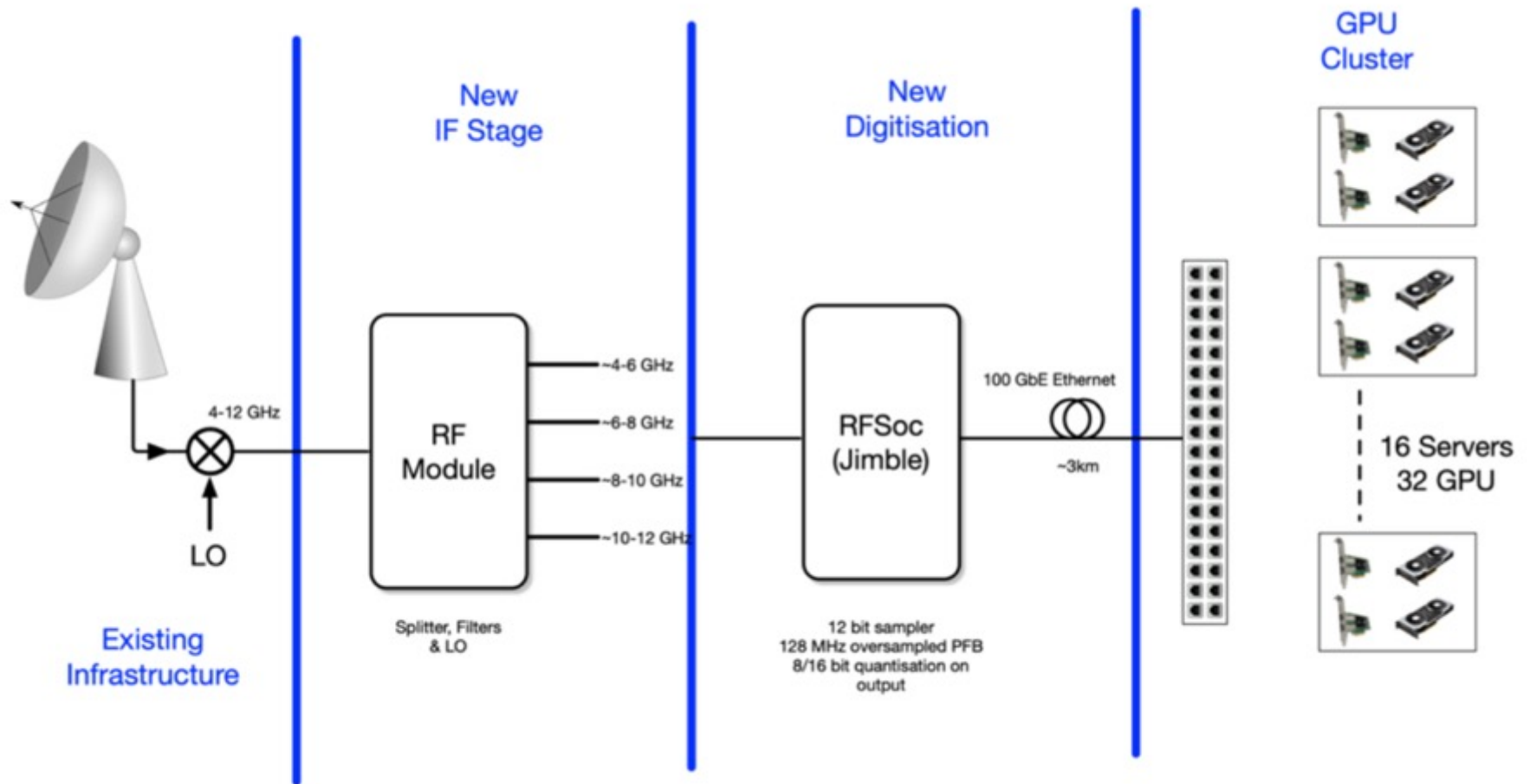
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ATCA-BIGCAT upgrade

- BIGCAT: Broadband Integrated Gpu Correlator for ATca
 - Replacement of CABB digitisers and correlator with a hybrid FPGA+GPU backend
- LIEF funded, lead by WSU (PI Ray Norris)
- Key aspects of BIGCAT:
 - Double instantaneous bandwidth to ~8 GHz
 - Flexible spectral resolution
 - Improved reliability
 - More flexibility:
 - Many more options wrt. frequency resolution and integration times
 - Ability to change quickly between different observing modes
 - More adaptable to automated observing (e.g. rapid ToO follow-up)
 - Retain standard CABB features (e.g. mosaicing)

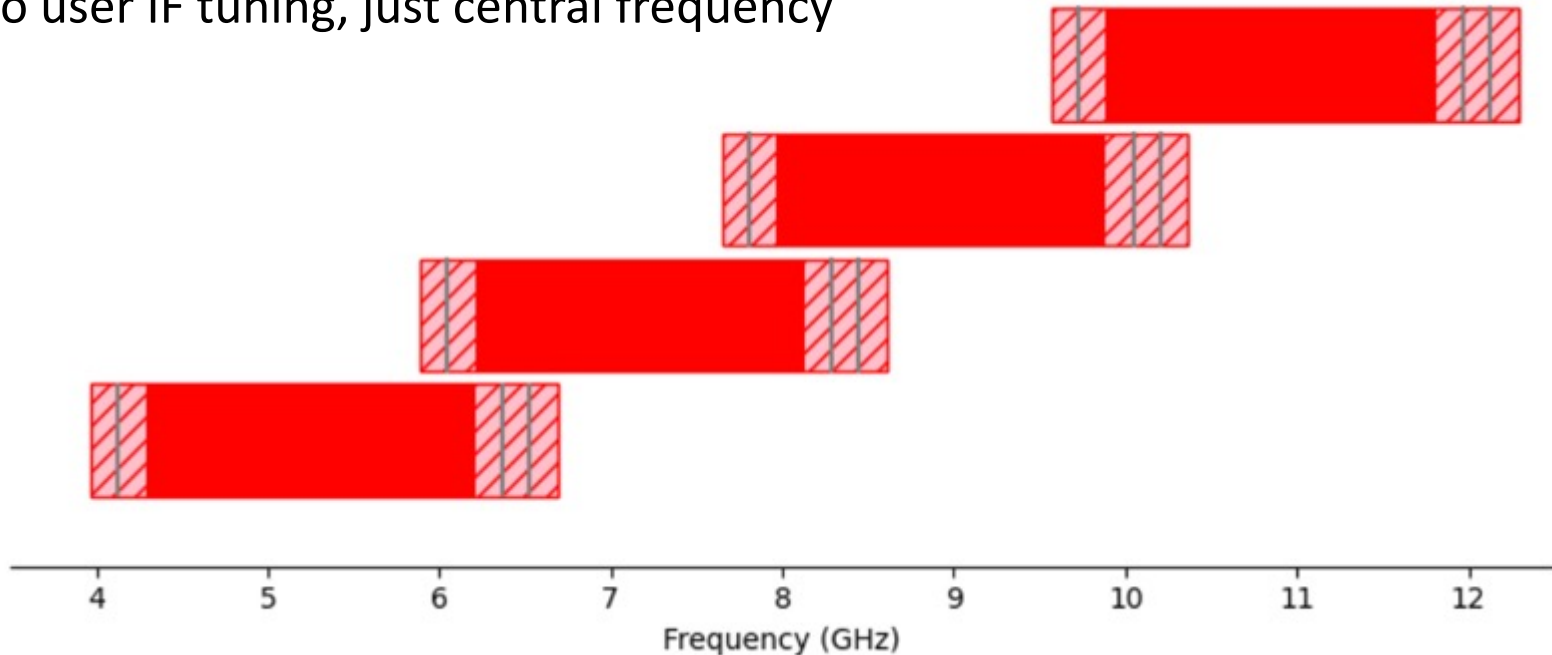
ATCA-BIGCAT signal chain



https://www.atnf.csiro.au/research/bigcat/BIGCAT_Overview.pdf

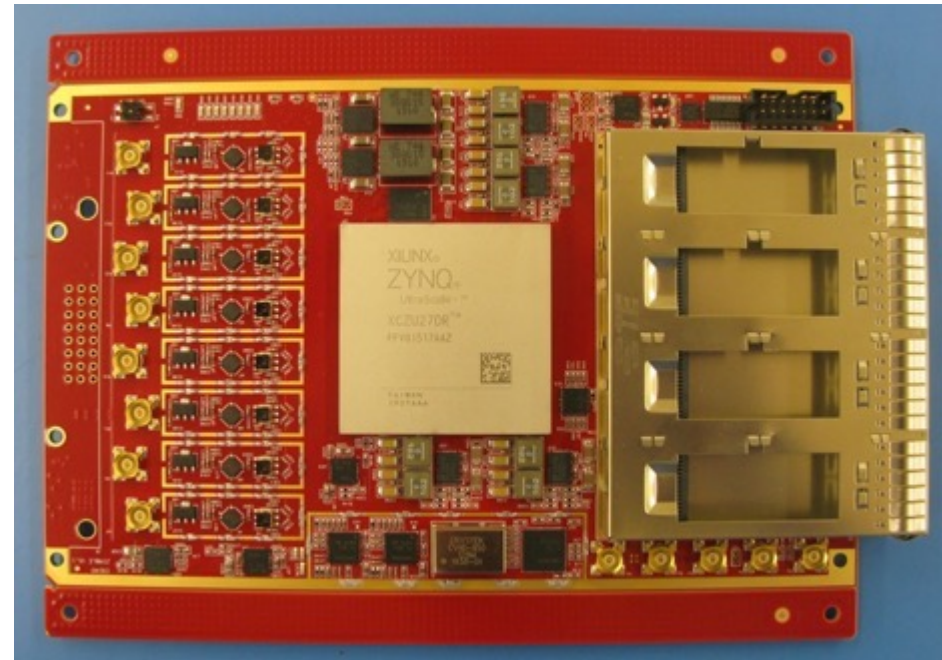
BIGCAT Frequency Bands

- 4 x 1.92 GHz IFs stitched together to form a contiguous 7.68 GHz bandwidth
- Limited tunability in 160 MHz steps
- Single "8 GHz" IF chunk for higher frequencies
- No user IF tuning, just central frequency



BIGCAT Digitisers

- 2 GHz IFs sampled using Xilinx RFSoc ADC
 - 8 samplers and FPGA fabric on single chip
 - 12 bits/sample
- Built into CASS “Jimble board”
 - 4x 2 GHz per board
 - 2x 100 GbE output
 - 128 MHz PFB, oversampled
 - 8/16 bit streamed to GPU
- Jimble common to CryoPAF project
 - Plan to upgrade for Parkes UWB systems using same firmware as BIGCAT



GPU Backend

- GPU backend 16 servers, each with 2 GPU and 2x100 GbE
 - RTX 3000 series current plan (High end gaming card)
 - 58 Gbps/GPU (without reference antenna)
- Each 128 MHz coarse frequency channel processed independently
 - Same frequencies from all telescopes (ATCA + reference) to single GPU
 - No GPU cross connect planned
 - 2 coarse channels/GPU (except LS)
- Plan to allow bespoke “user” processing on GPU



BIGCAT Science Requirements

- Prepared by Elizabeth Mahony and distributed to LIEF members, ATUC, ATCA users
 - Continuum (~ 1 MHz, Mosaicing supported, on-the-fly)
 - Spectral line (flexible frequency setup)
 - VLBI, voltage beams, multiple beams
 - Pulsar binning (up to 1024 bins, 128 kHz spectral resolution)
 - Fast dump visibilities (up to 60 Gbit/s)
 - Satellites & near field asteroids

https://www.atnf.csiro.au/research/bigcat/bigcat_techreq.pdf

Spectral Flexibility

- Up to 18.5 kHz over entire 1-3 GHz band (4 km/s)
- Up to 74 kHz over 8 GHz at > 4 GHz (5.5 km/s @ 4 GHz)
 - “Spectral window” selectable to reduce data rate
- Up to 0.01 km/s over limited bandwidth (zoom bands)
 - Up to 8 zoombands per 128 MHz
 - Maximum 262144 spectral channels over ALL spectral windows and zoom
- Specific spectral resolutions will be “pre-canned” modes, but new modes easily created then tested before being offered to astronomers
 - E.g. 4 zoom bands (per 128 MHz) of 16 MHz, each with 4096 spectral points
 - Constrained by GPU resources and visibility data rate
 - Zoom bands and spectral windows “tunable” by observer

Timeline

- RF design being worked on
 - Competition with CryoPAF and Quasar
 - RF prototypes expected to be built and tested over 2022
 - Production 2023
- Jimble hardware to be assembled by May 2022
 - Currently CryoPAF firmware being designed. Will be then ported to BIGCAT
 - Completed mid-2022
- Install 4 GHz BIGCAT system (with existing IF system) October 2022
 - Decommission CABB
 - Unclear if practical to have overlap period
- RF upgrade second half of 2023
- Catastrophic failure of CABB “unlikely”, degradation possible

BIGCAT Processing Software

- Plans to continue to support Miriad and CASA
 - RPFITS to be deprecated so new ingest code to be written
- No major changes to processing code planned
 - “Only” 2x bandwidth change

Feedback and questions:

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BIGCAT Overview: https://www.atnf.csiro.au/research/bigcat/BIGCAT_Overview.pdf

BIGCAT Requirements: https://www.atnf.csiro.au/research/bigcat/bigcat_techreq.pdf

Thank you

CSIRO Astronomy and Space Science

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