

# ASTERICS

## Connecting multi-messenger astrophysics

Giuseppe Cimò

*Joint Institute for VLBI - ERIC (JIVE)*

*Netherlands Institute for Radio Astronomy (ASTRON)*

# what is ASTERICs?

## Astronomy ESFRI & Research Infrastructure Cluster

A major collaboration (23 partner institutions) in astronomy, astrophysics and astroparticle physics.

It is funded by EC Horizon 2020 framework at € 15M for 4 years (2015-2019)



### Scope of ASTERICs:

To help solve the **Big Data** challenges of European astronomy  
To provide direct interactive access to the best European astronomy data in an international framework

*Cross-cutting synergies and common challenges*

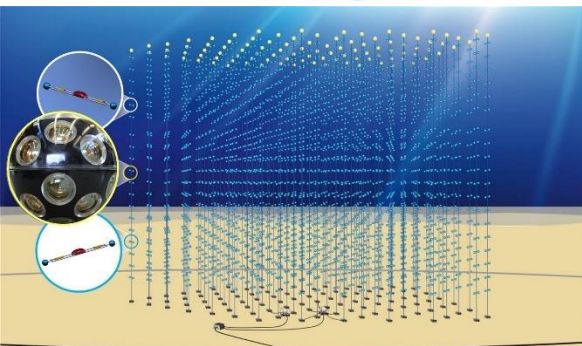


# concept and approach

- Supporting the European Strategy Forum on Research Infrastructures (ESFRI)

European Strategy Forum  
on Research Infrastructures





## A multi-km<sup>3</sup> neutrino telescope

Exploring our galaxy for high energy neutrino sources  
KM3Net2 on timescale of 2020



### SKA-LOW, Australia

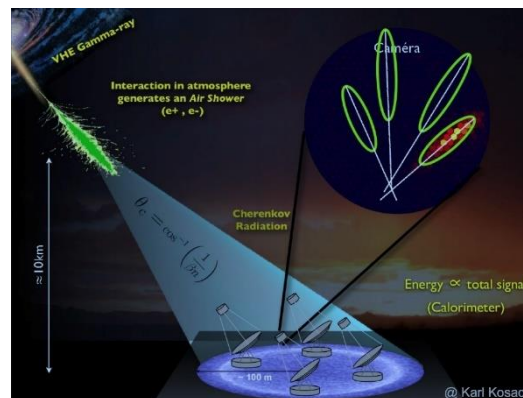
Phase 1: 130,000 dipoles over 80 km  
Phase 2: 500,000 dipoles over 250 km

### SKA-MID, South Africa

Phase 1: 200 dishes over 150 km  
Phase 2: 2500 dishes over 3500 km

Phase 1 (2018-2023)  
Phase 2 (2025-2033)

**Challenges everything...**



Very high energy  $\gamma$ -ray observatory  
Two arrays of 100 (N) and 20 (S) telescopes  
Event re-construction  
Complex metadata  
Streaming and processing challenges  
Precursors: MAGIC and HESS



## General purpose optical/infrared telescope

- high redshift galaxies
- star formation
- exoplanets
- protoplanetary systems

### E-ELT



# concept and approach

- Supporting the European Strategy Forum on Research Infrastructures (ESFRI)
- Aspiring ESFRI projects + pathfinders
- Other world-class research infrastructures
  - e.g. LOFAR, Euclid, LSST, Virgo

European Strategy Forum  
on Research Infrastructures



ESFRI

Together, the ESFRI projects – and their pathfinders – open new windows on the universe, significantly extending our observational capabilities across the electromagnetic spectrum, in addition to neutrino detectors and gravitational waves





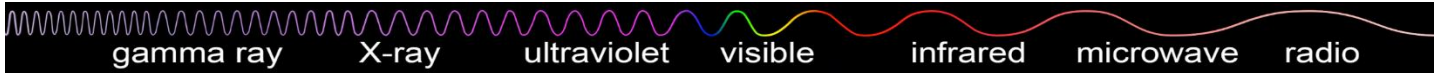
# addressing common challenges in astronomy and astroparticle physics

- ***supporting*** and ***accelerating***  
the implementation of a new  
generation of observatories
- ***enhancing performance***
- helping scientists to access  
data

*ESFRIs interoperating as an  
integrated multi- $\lambda$ , multi-  
messenger facility*



# multi- $\lambda$ , multi-messenger

- messengers: **photons,  $\nu$ , gravitational waves, VHE $\gamma$**
- multi- $\lambda$ : 
- transient source astronomy

## *To make it happen...*

- Cooperation, Interoperability, Open Data
- Scalability – processing and analysis
- Big Data, Data mining, Data Access
- Streaming and timing

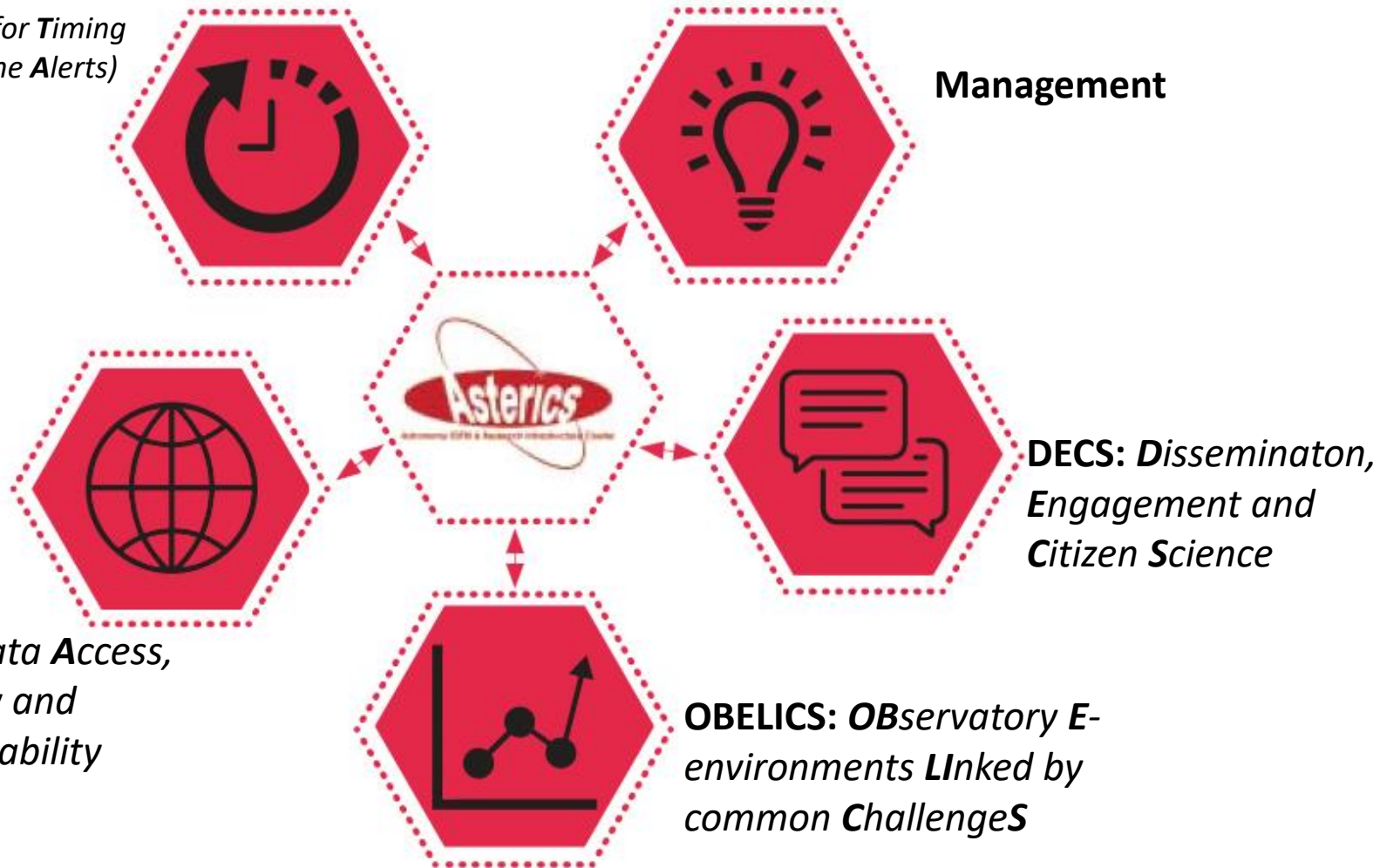
# connections & openness

- connecting infrastructures:  
enhancing individual capabilities -  
necessary for science!
  - *ICT: high speed data  
transport/timing*
- Embracing **Open Science, Open  
Access, Open Data**
  - improve knowledge circulation
  - many challenges, many  
opportunities
- Engage with society at large
  - Astro community, education, public





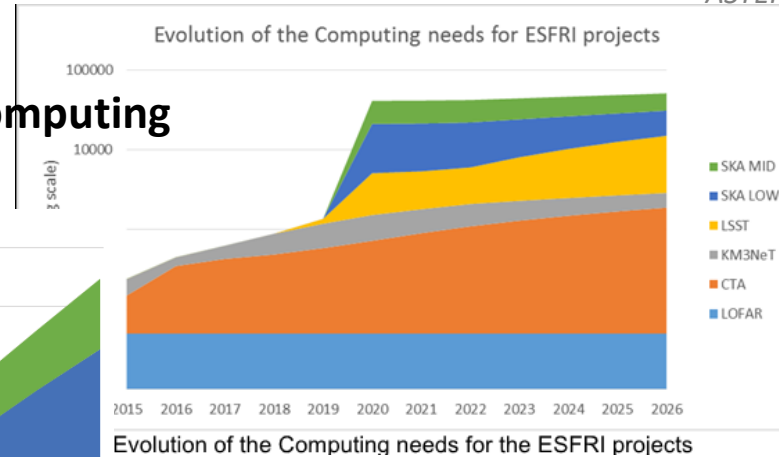
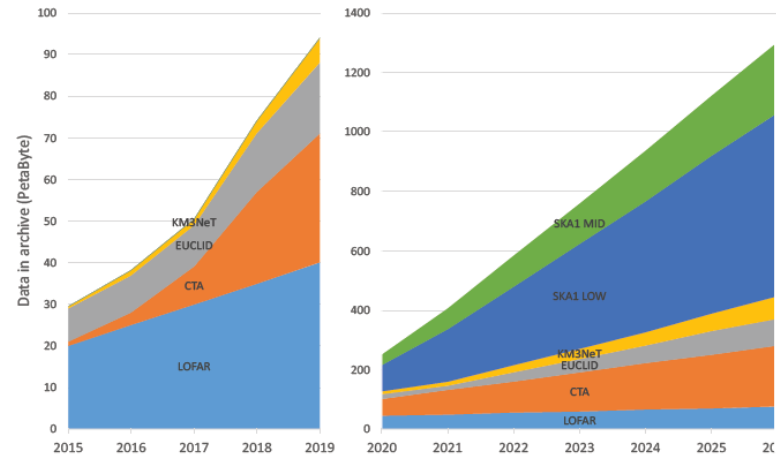
**CLEOPATRA:** *Connecting Locations of  
ESFRI Observatories and Partners in  
Astronomy for Timing  
and Real time Alerts)*



# Challenges

- **Big Data**
  - **Data Access**
  - **Multi-Messenger**
  - **Open Science**
- Parallel software programming techniques, to adopt big-data software frameworks, to benefit from new processor architectures and e-science infrastructures.
  - Software re-use and co-development of technology
  - Adapt and optimise extremely large database systems
  - Virtual Observatory and VOEvents
  - Cooperation with the ESFRI pathfinders, computing centres, e-infrastructure providers and industry
  - Public engagement and Citizen Science

## Evolution of storage needs and computing needs for the ESFRI Projects



## Innovative application of the *docker* container

Jupyter notebook of the CASA 3C391 VLA continuum tutorial running inside a web browser on a tablet

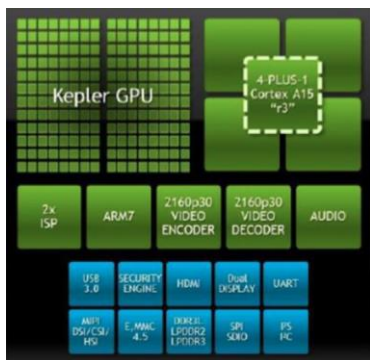


**Lossless compression algorithm** produced for astro(particle) experiments. Now under discussion and tests in more ESFRI scenarios.

Libraries for **fast, vectorised, array reductions and moment calculations** written in C++ with Python binding, potentially have **an impact outside of astronomy** since they implement fairly generic operations

# Big Data

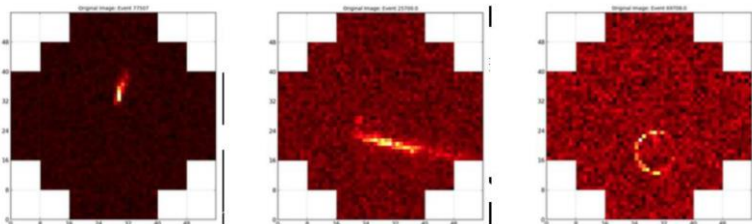
**Low power architectures benchmarking** for building innovative computing, storage and data handling platforms.



Optimisation and parallelization of algorithms for data reduction and analysis in order to exploit the low-power computing facilities

- Development of **high performance computing solution**
- Notebook interface (VLBI in the Cloud) can be of use for **all radio interferometric data analysis**
- General-purpose library of **A&A** and **workflow management** systems
- Contribution to define specifications in an **open data format** within the gamma-ray community
- **Data streaming and architecture** of the data processing unit in SKA

Exploitation of LPC architectures for **map/reduce applications**.



# Data Access

## Virtual Observatory

- The VO can be seen as a kind of club of **data services** that all follow the same rules.
- International Virtual Observatory Alliance  
    IVOA standards
- Interface between domain-specific & generic infrastructure
- New and archive data

<b>DADI technology Forum</b>	17-18 Sept 2015, Strasbourg
<b>ASTERICS European Schools</b> open call to European Astronomers (PhD, Post-doc) 50 participants	Dec 2015
<b>ESFRI Forum and training event</b> network, share lessons learnt, discuss requirements, training	Dec 2015
<b>Data provider Forum</b> open to all European Data providers	Nov 2016, Heidelberg



# Virtual Observatory

Visualisation with VO tools  
 IceCube-40 String data

- 12,877 candidate neutrino events
- Event list managed using the Topcat VO tool

<http://www.star.bris.ac.uk/~mbt/topcat/>

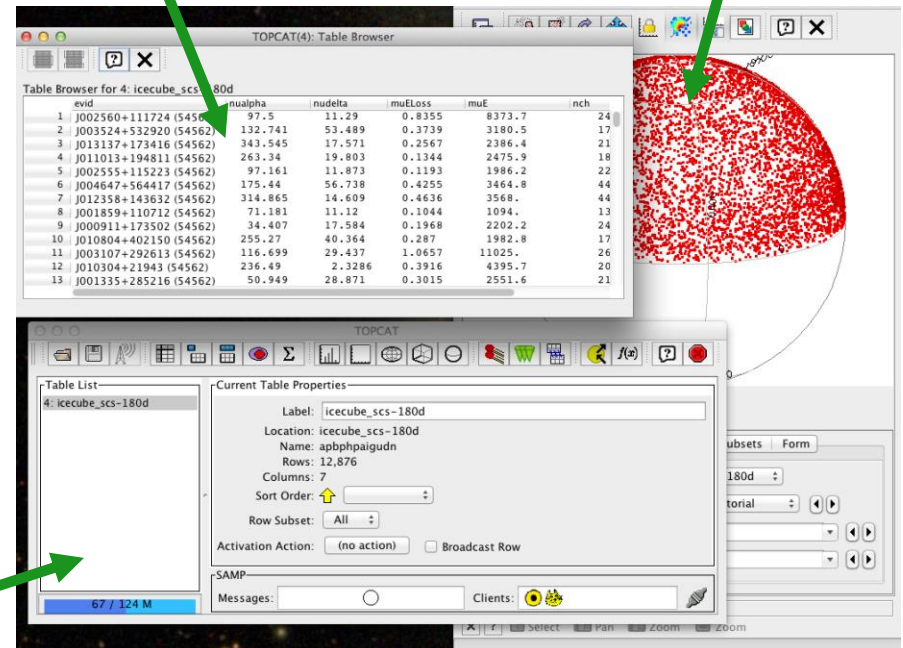
- Visualisation through the Aladin VO portal

<http://aladin.u-strasbg.fr/aladin.gml>

database  
 loaded into  
 Topcat

view  
 table

sky plot

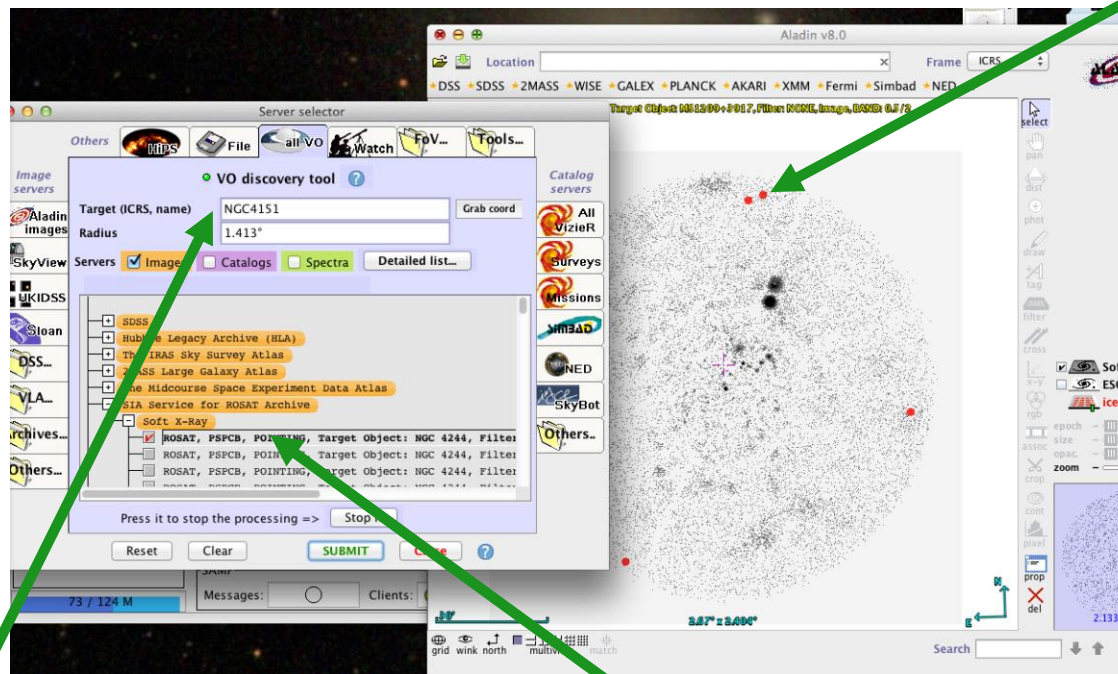


The screenshot shows the Topcat Table Browser interface. The main window displays a table with the following columns: 'evId', 'alpha', 'nDelta', 'muLoss', 'muE', and 'nch'. The table contains 13 rows of data. A green arrow points from the 'view table' text to the table. Another green arrow points from the 'sky plot' text to a circular plot on the right side of the interface, which displays a distribution of red points representing the neutrino events on a celestial sphere.

evId	alpha	nDelta	muLoss	muE	nch	
1	J002560+111724 (S4562)	97.5	11.29	0.8355	8373.7	24
2	J003524+532920 (S4562)	132.741	53.489	0.3739	3180.5	17
3	J013137+173416 (S4562)	343.545	17.571	0.2567	2386.4	21
4	J011013+194811 (S4562)	263.34	19.803	0.1344	2475.9	18
5	J002555+115223 (S4562)	97.161	11.873	0.1193	1986.2	22
6	J004647+564417 (S4562)	175.44	56.738	0.4255	3464.8	44
7	J012358+143632 (S4562)	314.865	14.609	0.4636	3568.	44
8	J001859+110712 (S4562)	71.181	11.12	0.1044	1094.	13
9	J000911+173502 (S4562)	34.407	17.584	0.1968	2202.2	24
10	J010804+402150 (S4562)	255.27	40.364	0.287	1982.8	17
11	J003107+292613 (S4562)	116.699	29.437	1.0657	11025.	26
12	J010304+21943 (S4562)	236.49	2.3286	0.3916	4395.7	20
13	J001335+285216 (S4562)	50.949	28.871	0.3015	2551.6	21

# Virtual Observatory

show neutrino events  
overlaid on X-ray image



In Aladin, search for  
images of NGC 4151

pick this ROSAT image



# ASTERICS connections: gravitational waves

ASTERICS fostered  
use of VO for grav  
wave EM follow-up

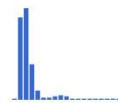


## Skymap Viewer

A sky atlas for understanding LIGO-Virgo skymaps. Help [here](#), or watch a [video about Skymap Viewer](#). Plenty simulated skymaps [here](#). If you do not see the big dark sky map, look below and widen your browser. Zoom with the + and - at the right of the sky.

### LIGO-Virgo Skymaps

This is skymap  
GW150914:LALI.  
50% area = 149.0 sq deg  
90% area = 616.4 sq deg



South ..... North

Show Weighted Galaxies (or table).

### Time and Place

Universal time

2015-09-14T09:50:45

E Longitude  Latitude

Sun =  and  = Moon

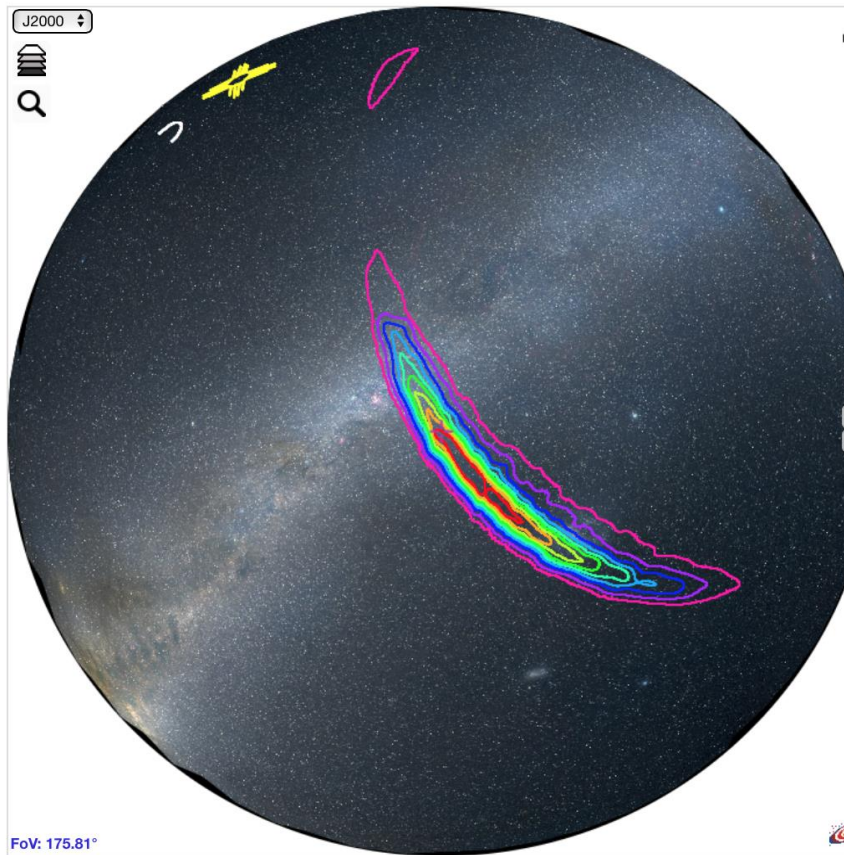
### Catalog Sources

Click the Layers icon  to switch on catalogs.

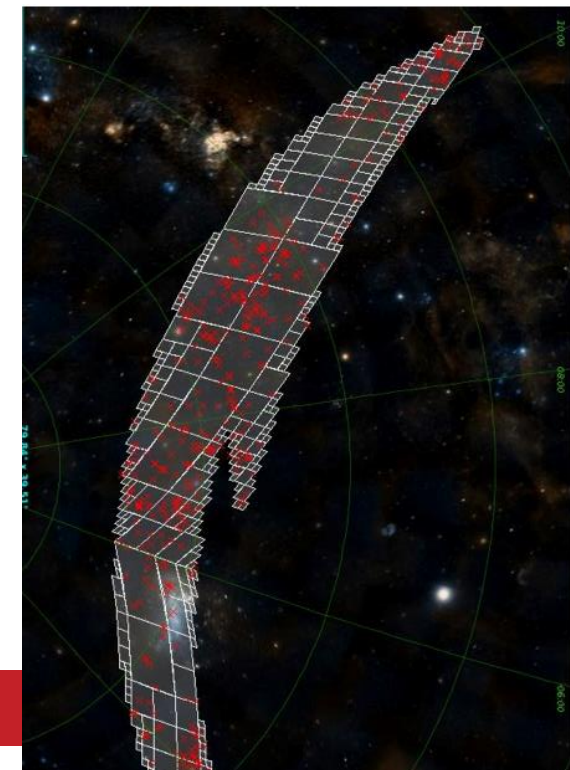
If you click on the sources on the sky, information will appear here with links to Simbad and NED.

### Zoomable Multiwavelength Sky

Zoom in on the sky with the mouse or the +/- icons on the right of the sky. To change the image

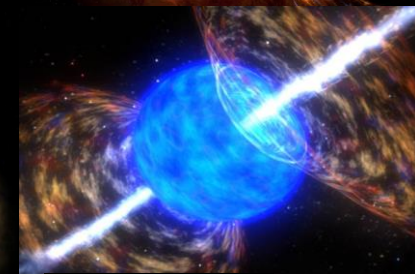
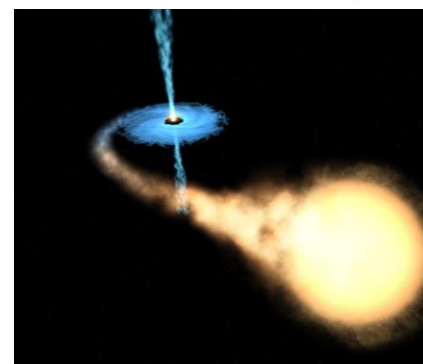
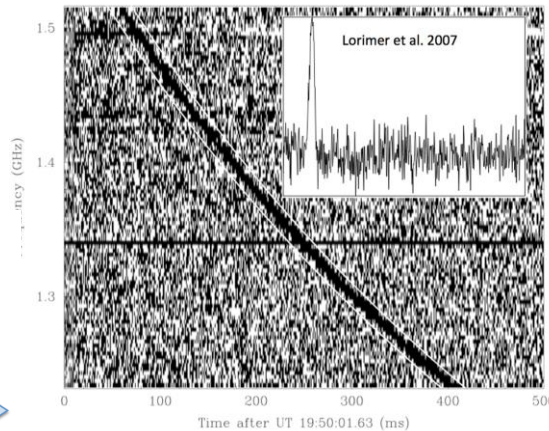
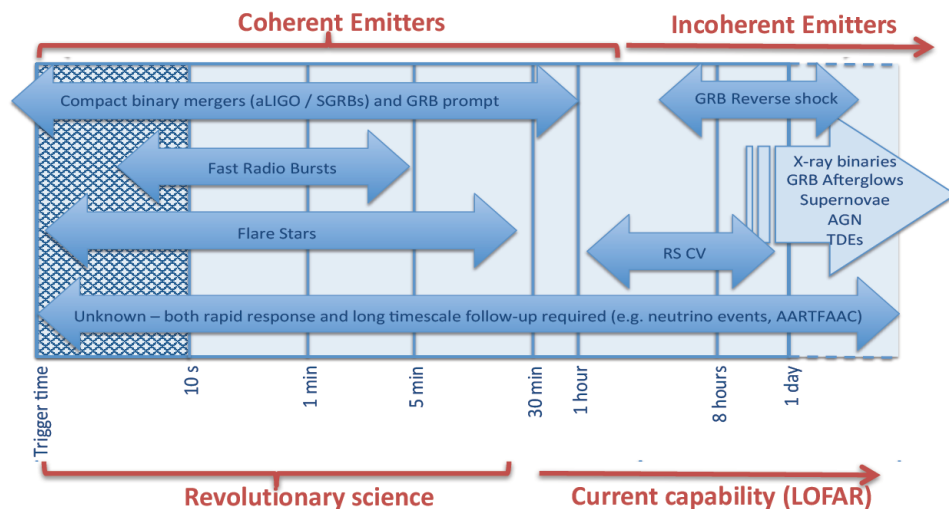


FoV: 175.81°





# Multi-Messenger Challenges

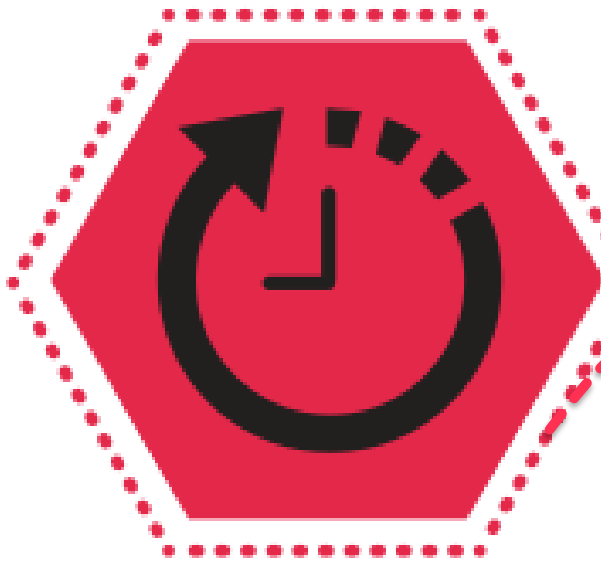


- Develop standards for generation, dissemination, distribution and reaction to transient events (based on **VOEvents**)
- Demonstration:  
LOFAR, EVN, follow up a GW event
- Investigate scientific synergies for automated followup observations

# multi-messenger timing and synchronisation

Connecting real facilities now as path to connected  
future facilities

- Building on success of e-VLBI
  - EXPReS, NEXPReS
- ...here comes the White Rabbit

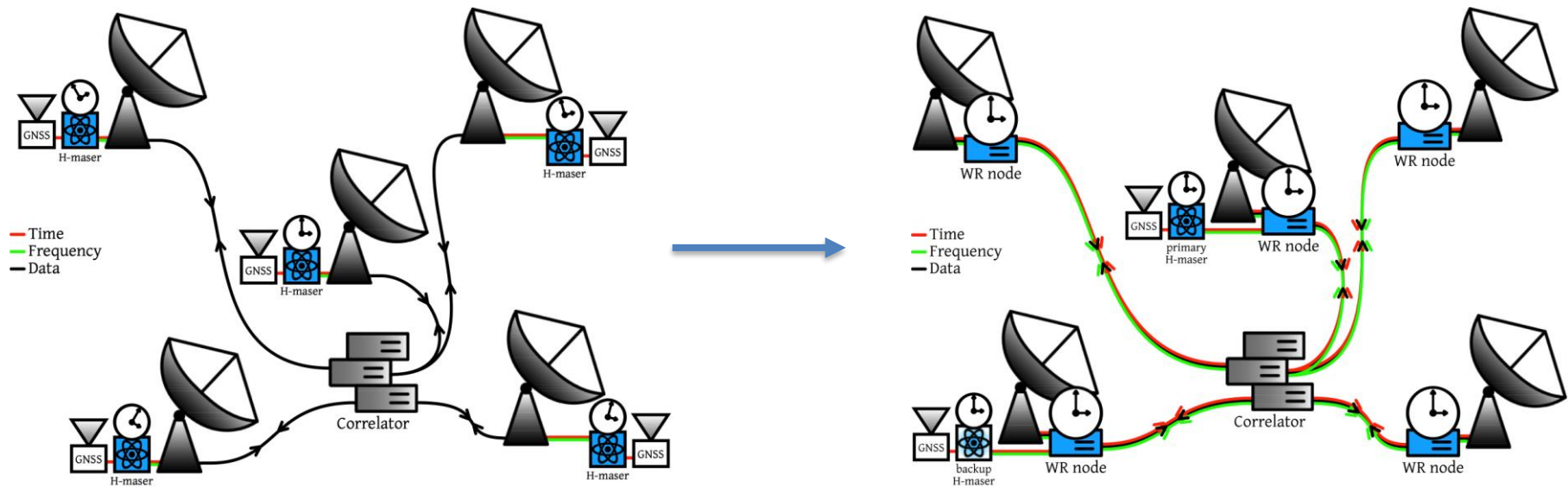




## White Rabbit Ethernet (CERN, based on IEEE Precision Time Protocol)

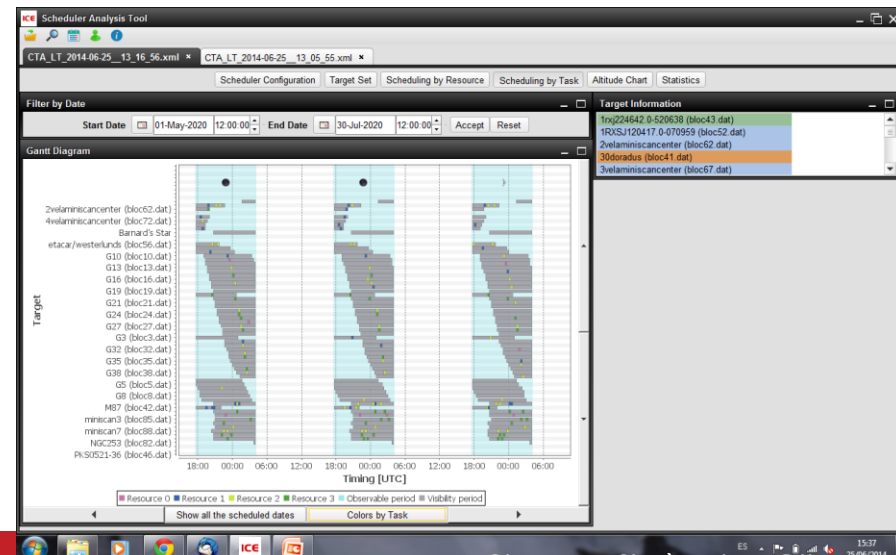
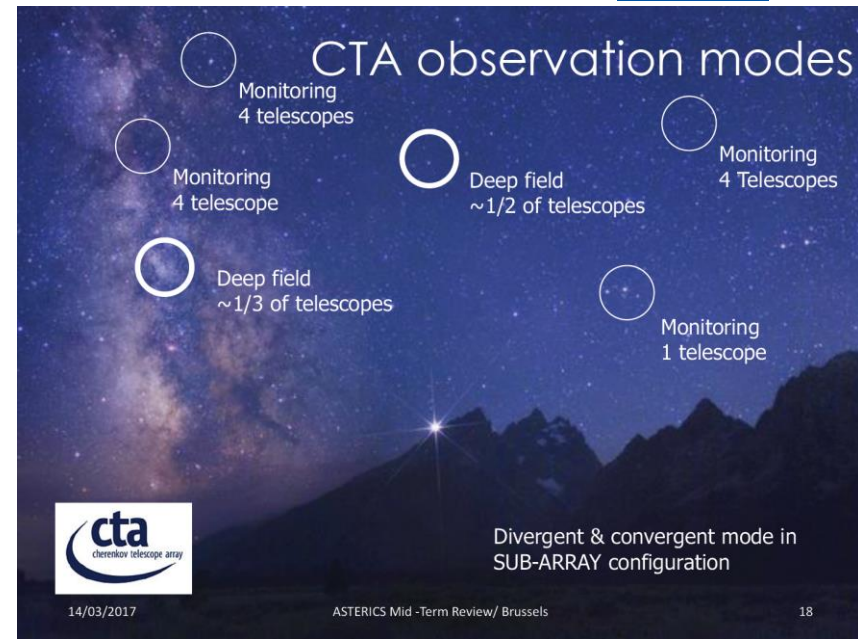
- Time, frequency, and 1 Gb/s data in one
- 1 PPS, 10 – 125 MHz
- Designed for 1 ns timing over distances <10 km (LHC, CERN)
- Commercially available

- ASTERICS made real impact in WR community
  - Improvements are fed back → open source
- Generated serious outside interest
- Fed into tender for new photonic equipment for SURFnet
- WR timing in design of SKA-low
  - Frequency stability now good enough for frequency transfer
- Commercial applications under development



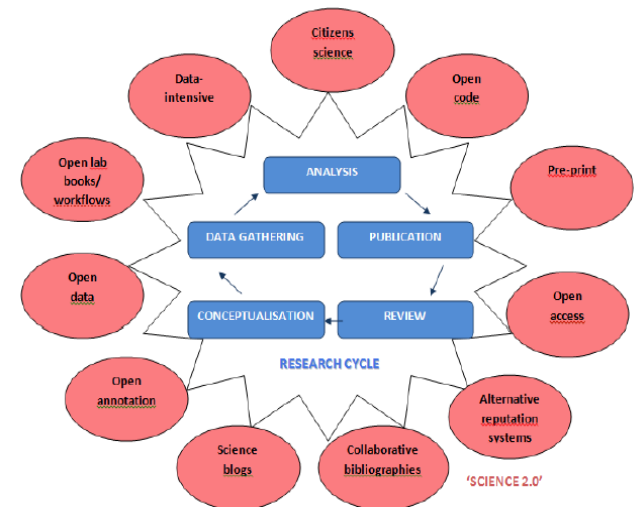
## Scheduling of large astronomical infrastructures

- Complex, many-element detector arrays
- Artificial Intelligence (AI) techniques to optimize procedures
  - Metaheuristic Optimization: Genetic Algorithms, Multi-objective Evolutionary Algorithms
  - Constraint-Based Reasoning: constraint propagation
- Maximize science return of SKA and CTA
  - But ensure solutions are applicable for other instruments
- Incorporate multi-frequency, multi-messenger astrophysics
- Provide a framework to coordinate and schedule multiple facilities.



# Open Science

- Open ESFRI facilities to wider stakeholders through citizen science (Open Science, or ‘Science 2.0’)
- Audiences: scientific & technical communities, academia, private industry, other public research centres, SMEs, policy makers, general public
- Coordinated citizen science experiments to open ESFRIs & pathfinders/precursors to public
- Educational resources & efficacy metrics



**Citizen Science is not outreach!**

# Citizen Science

## Pulsar Hunters ++

- Lead: Rene Breton
- Science goal: Extend the successful Pulsar Hunters Zooniverse project to harder-to-find pulsars
- Activity: interactive data visualization of pulsar time and frequency domain data; future application to SKA

## CREDO

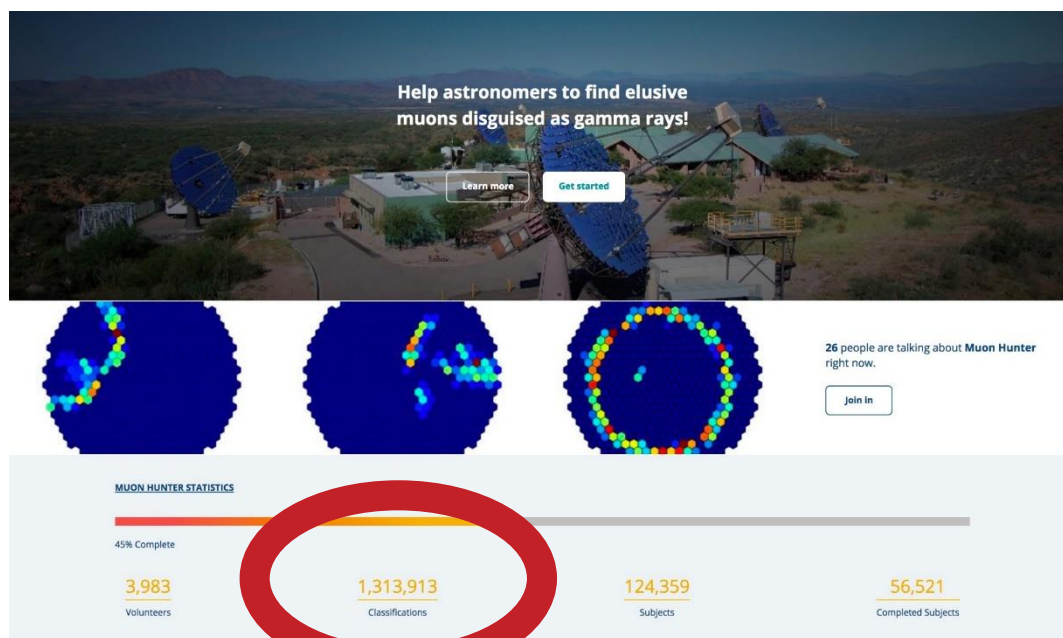
### Cosmic Ray Extremely Distributed Observatory

- Lead: Piotr Homola  
<https://credo.ifj.edu.pl/>
- Science objective: detect ultra-high-energy charged particles with a whole-Earth Cherenkov detector
- Activity: use mobile phones as charged particle detectors

# Citizen Science

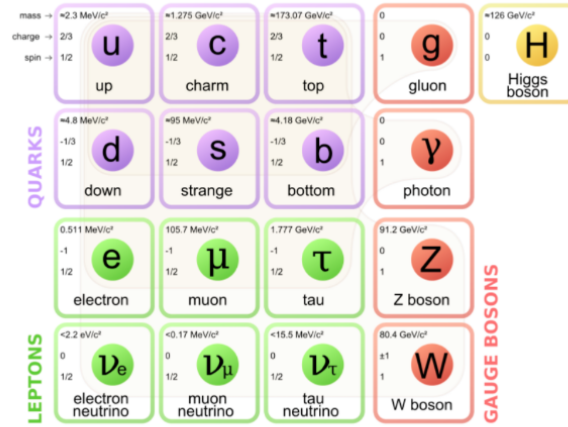
## Muon Hunters

- Lead: Lucy Forston, CTA
- Science goal: detect fainter Cherenkov events by visual classification
- Activity: classify hadron vs. photon events in the CTA telescopes, morphologically and in the time domain; apply first to simulations and to e.g. HESS



**1.3 million classifications in the first five days!**





The 'standard model' of particle physics. The electron and muon can be found in green boxes halfway down the diagram.

## So what is a muon?

A muon is a type of subatomic particle, which is very similar to an electron – for instance, they both have the same negative electric charge. The main difference between a muon and an electron is their mass. A muon is 207 times more massive than an electron! For comparison, you might have known that the mass of a proton (the nucleus of a hydrogen atom), is about 1,800 times that of an electron. However, unlike the proton, which has substructure and is composed of other particles, the muon is a fundamental particle in its own right.

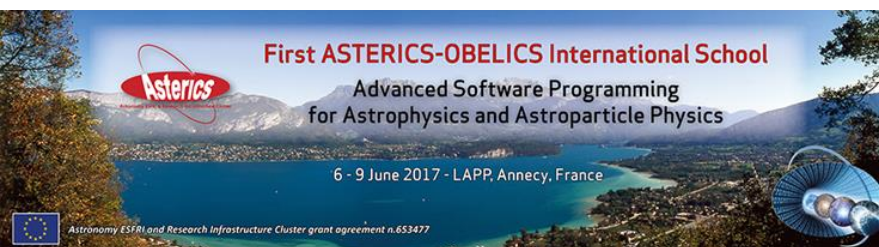
If you think the existence of the muon is strange, you're in good company. The world-famous physicist I. I. Rabi, when first told of the discovery of the muon, said in response, "Who ordered that?" There's good reason why the muon is such an unfamiliar particle: muons are radioactive; they decay with a mean lifetime of 2.2 microseconds. That's  $2.2 \times 10^{-6}$  seconds, or 2.2 millionths of a second. Muons don't stick around long enough to become part of the matter we encounter day to day.

However, there are lots and lots of muons all around us, created in interactions we don't usually think of...

# Strengths from connections

## Enabling data science

- Training and support
- Skill sets for astronomy



## Participating institutions



## Supporting organisations and networks

