### THE AT CATALOGUE

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#### 1 OVERVIEW

The AT catalogue consists of two separate databases: AT\_CAT and SOURCE\_CAT. SOURCE\_CAT is a conventional astronomical catalogue composed of a small number of well-known and frequently-accessed astronomical catalogues. AT\_CAT is a record of all observations made (or to be made) on the AT. together with specific information on calibrator sources.

#### 2 DEFINITIONS

# 2.1 RDB Jargon

These catalogues use RDB, so that searches and other manipulation of data may be done fairly painlessly. Most of this document may be understood with no knowledge of RDB, given the following informal definitions:

- A RELATION is what might colloquially be called a table.
   I.e. a star catalogue might be stored as one relation.
- 2. A FIELD corresponds to one column in that table.
- 3. A RECORD corresponds to one line in that table.
- 4. Unlike a table, RDB databases can effectively be n-dimensional, by using several orthogonal but intersecting relations. This is the origin of otherwise obscure comments about creating extra relations for several frequencies. etc.

## 2.2 Source Names

The following is copied from a separate AT note (in which the choice is justified) in preparation:

Each source observed on the AT will be assigned a 15-character name in the format:

### hhmmss.stddmmss

where hhmmss.s represents the J2000 RA, and ddmmss represents the J2000 dec. These names, which are consistent with the latest IAU recommendation, will be used within all AT software to define the

source uniquely. Other names, such as 3C273,  $0rion-\lambda$ , 2134+004, will be stored as aliases. In all software that communicates with the observer, the observer may refer to a source by its alias rather than the  $\lambda T$  name.

### 2.3

#### 3 OBSERVATION STRUCTURE

The following is also taken from the separate AT note in preparation:

Figure 1 shows the structure of an observation. The largest unit is a <u>PROJECT</u>, which is defined by a proposal. This project may include many sources, frequencies, or array configurations. The project consists of a number of <u>OBSERVATIONS</u>. Each of these may typically be 12 hours long, and represents a continuous period of time assigned to a project on one array. Each observation consists of a number of <u>SCANS</u>, which may typically be 10 mins long. Each scan represents a continuous observation of one source. The scan in turn is divided into <u>INTEGRATIONS</u> and <u>CYCLES</u>, each of which have a typical duration of a few seconds.

The following assumptions implicit in these definitions should be noted.

- 1. During an OBSERVATION (typically 12 hours), the project number and array configuration are unchanged.
- During a SCAN (typically 10 mins), the source, bandwidth, and number of channels per baseline are unchanged, although the frequency may change.
- During an INTEGRATION (typically 10 s), all observing parameters are held constant.

### 4 THE AT SOURCE CATALOGUE

### 4.1 Introduction

The AT source catalogue (database pathname SOURCE\_CAT) is a compilation of astronomical source catalogues. It does not attempt to include all known source catalogues, but merely those which are thought to be sufficiently useful to be accessible at all times through the database. It is designed to fulfil three functions:

1. When preparing an observation for the AT, an observer should be able to specify a source name (e.g. Orion-KL, IRAS00085-321) and the AT scheduler will be able to recover the coordinates from the database.

- An astronomer wishing to select sources from a catalogue (e.g selecting survey sources according to some criteria) may do so relatively easily using the database system.
- 3. It also enables rapid access to any of the constituent catalogues for reference purposes. In the case of the IRAS catalogue, this includes material (e.g. source identifications) which are not available on the microfiche version.

The relations and their fields will not be described here in detail, as they simply correspond more or less to the tables in the catalogues. The only exception is the ALIAS relation, which is made seperate from the other parts of the catalogues for reasons associated with Rdb philosophy.

## 4.2 The Constituent Relations

The following relations are included in the database:

IRAS The IRAS point source catalogue BOLOGNA The Bologna redshift survey MOLONGLO The Molonglo catalogue PARKES The Parkes catalogue

PN Planetary nebulae in the Galaxy and Magellanic clouds
ALIAS A list of source identifications matched to AT-type names

# 5 THE AT OBSERVATION CATALOGUE

### 5.1 Summary

The AT observation catalogue (database pathname AT\_CAT) is a compilation of observations which have been made, or are to be made, on the  $\lambda T$ . It contains the following relations:

SOURCE A list of sources and their J2000 positions

ALIAS Other names for the sources

CALIBRATOR Information on known calibrator sources

OBSERVATION A record of which observations have, or are to

be, made,

SCAN Details of each scan in the observation

### 5.2 The SOURCE Relation

This simply gives the source position. If there is more than one entry for a source, software will always choose the most recent entry

unless explicitly instructed otherwise.

DATE	Date of entry	TEXT 18
NAME	hhmmss.s-ddmmss	TEXT 15
RA	J2000 radians	G FLOAT
DEC	J2000 radians	G_FLOAT

### 5.3 The ALIAS Relation

This is used to translate from a well-known name (e.g. 3C273) to the AT name. Note that one source may have many ALIAS records.

DATE	Date of entry	TEXT :	18
NAME	hhmmss.s-ddmmss	TEXT :	15
ALIAS	other name	TEXT :	15

## 5.4 The CALIBRATOR Relation

This contains the following fields:

DATE	Date of entry	TEXT	18
ACTIVE_FROM	Active from date	TEXT	18
ACTIVE_TO	Active to date	TEXT	18
NAHE	Name (hhmmss.s+ddmmss)	TEXT	15
BAND	Freq. band (L.S. etc)	TEXT	1
1	Value of Stokes param I	REAL	
Q	Value of Stokes param Q	REAL	
ប	Value of Stokes param U	REAL	
V	Value of Stokes param V	REAL	
UVHIN	Min UV at which useful	REAL	
UVMAX	Max UV at which useful	REAL	
POSN	Position calibrator?	TEXT	1
FLUX	Flux calibrator?	TEXT	1
POLN	Polarisation cal?	TEXT	1

Note that each calibrator source has one record for each  $\Delta T$  frequency band.

# 5.5 The Observation Relation

This gives the data for each observation (typically 12 hours) and has the following fields:

DATE	Date of entry	TEXT 18
OBS_NO	Observation serial no.	SIGNED WORD
STATUS	See below	TEXT 8
START	Starting date/time	TEXT 18

STOP	Stopping date/time	TEXT 18
ARRAY_ID	ID of array/config	TEXT 8
PROJECT	ID of project	TEXT 8
NOTES		SEG.STRING

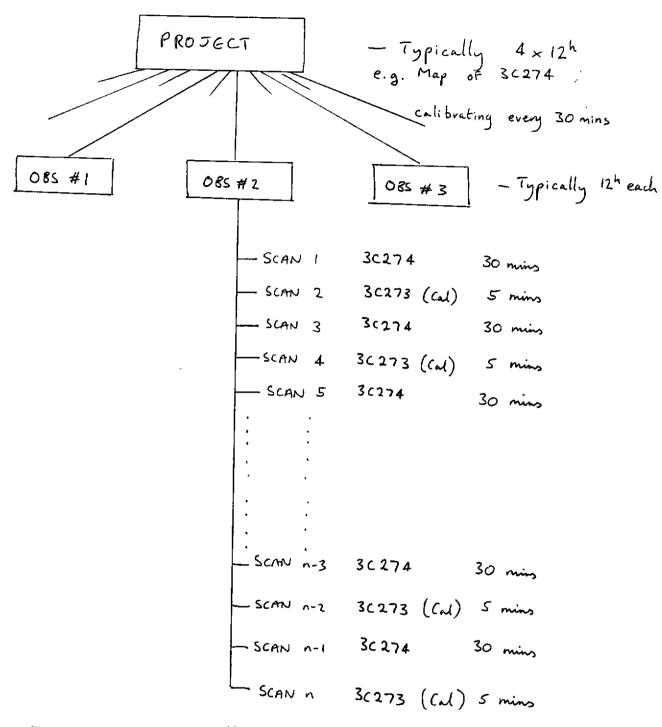
# 5.6 The Scan Relation

This gives the data for each scan (typically 10 minutes) and has the following fields:

DATE OBS_NO STATUS START STOP NAME BW NCHAN FREQ1 FREQ2 FREQ3 FREQ4 VELOC	Date of entry Observation serial no. See below Starting date/time Stopping date/time Source name Bandwidth/baseline Channels/baseline Centre freq for IF1 Centre freq for IF1 Centre freq for IF1 Centre freq for IF1 Doppler Velocity	TEXT 10 TEXT 18 TEXT 18 TEXT 15 REAL SIGNED WORD REAL REAL REAL REAL REAL REAL
VELOC_CODE	Doppler Velocity Optical, LSR, etc.	REAL TEXT 8

# Notes:

- The field STATUS might contain keywords such as 'PROPOSED'. 'SCHEDULED', 'COMPLETED', 'ABANDONED'
- 2. If thought necessary, the fields BW and NCHAN could also be expanded to enable separate values for each frequency band. In this case, a separate BANDS relation for each scan would probably be justified.



NB Frequency and bandwidth might also be changed from Scan to Scan E.g. when observing different moses transitions.

Fig. 1: Structure of an Observation