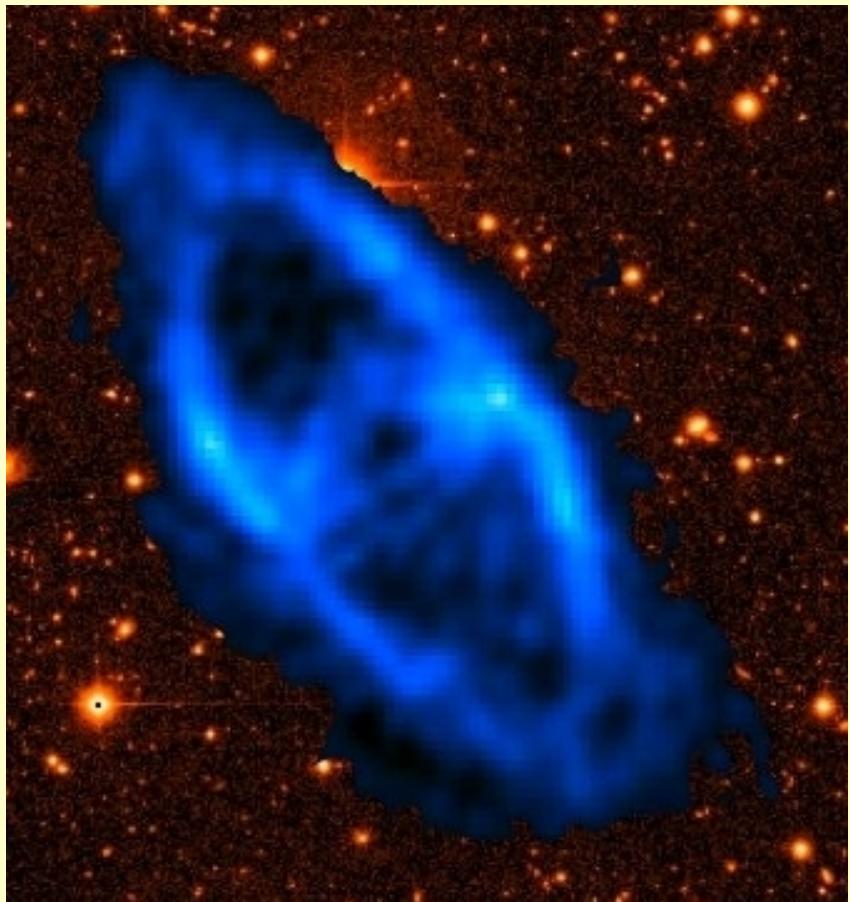


## TiRiFiC



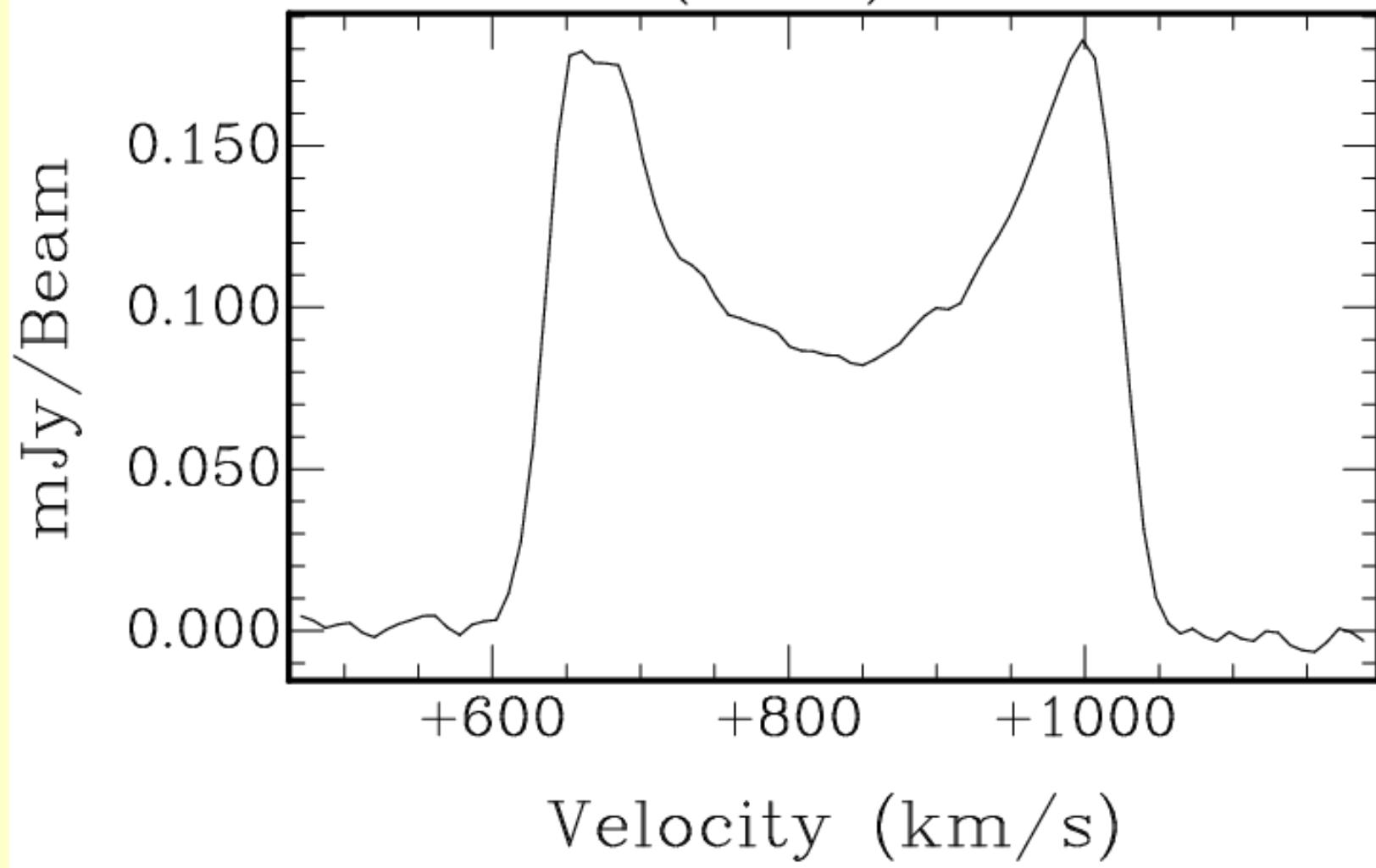
**Gyula I. G. Józsa**

# The task for HI WFSs

- HI in large galaxy surveys today                          -> interpret 1D spectra

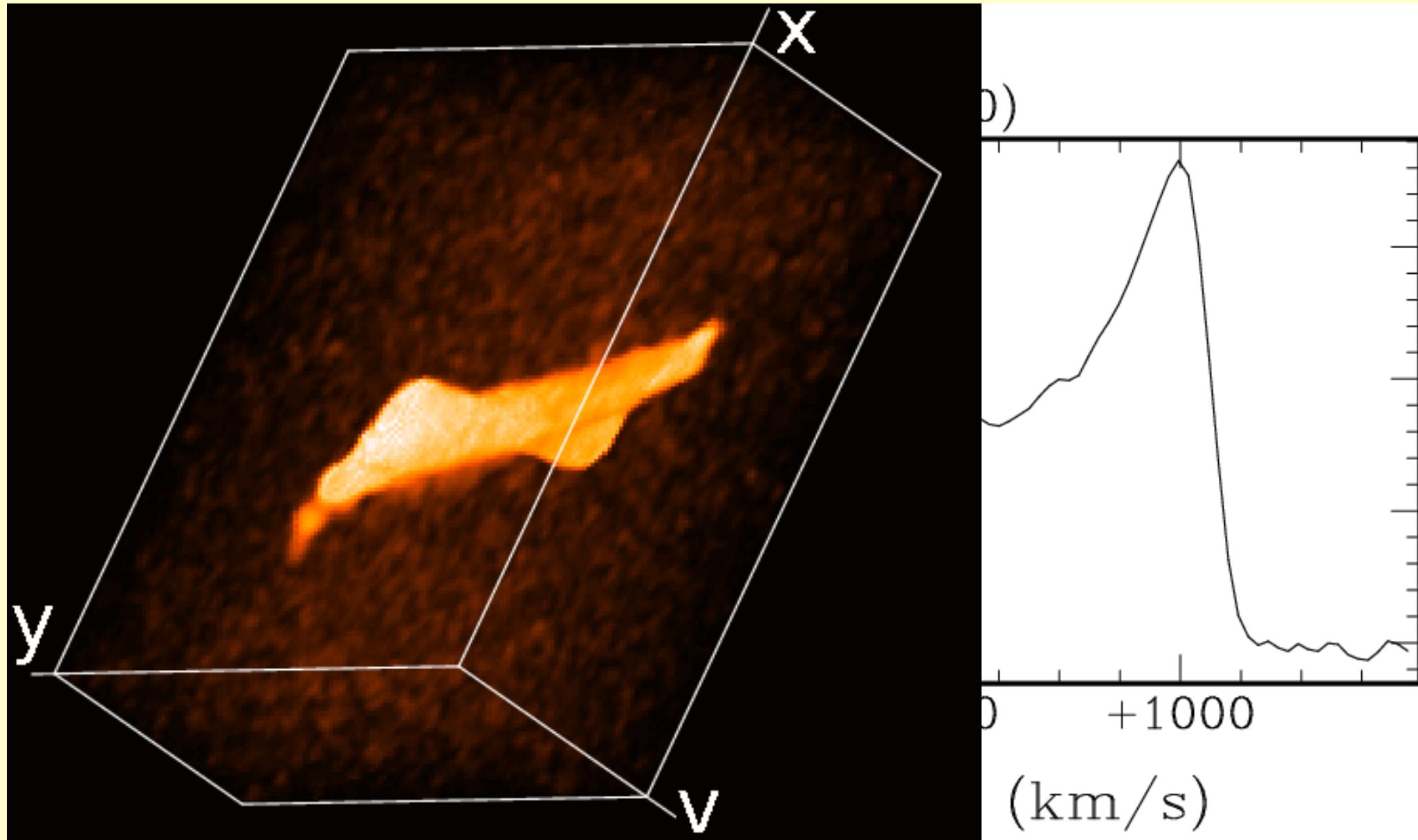
Ra:  $11^{\text{h}} 58^{\text{m}} 32.05^{\text{s}}$  (J2000)

Dec:  $+43^{\circ} 56' 37.00''$  (J2000)



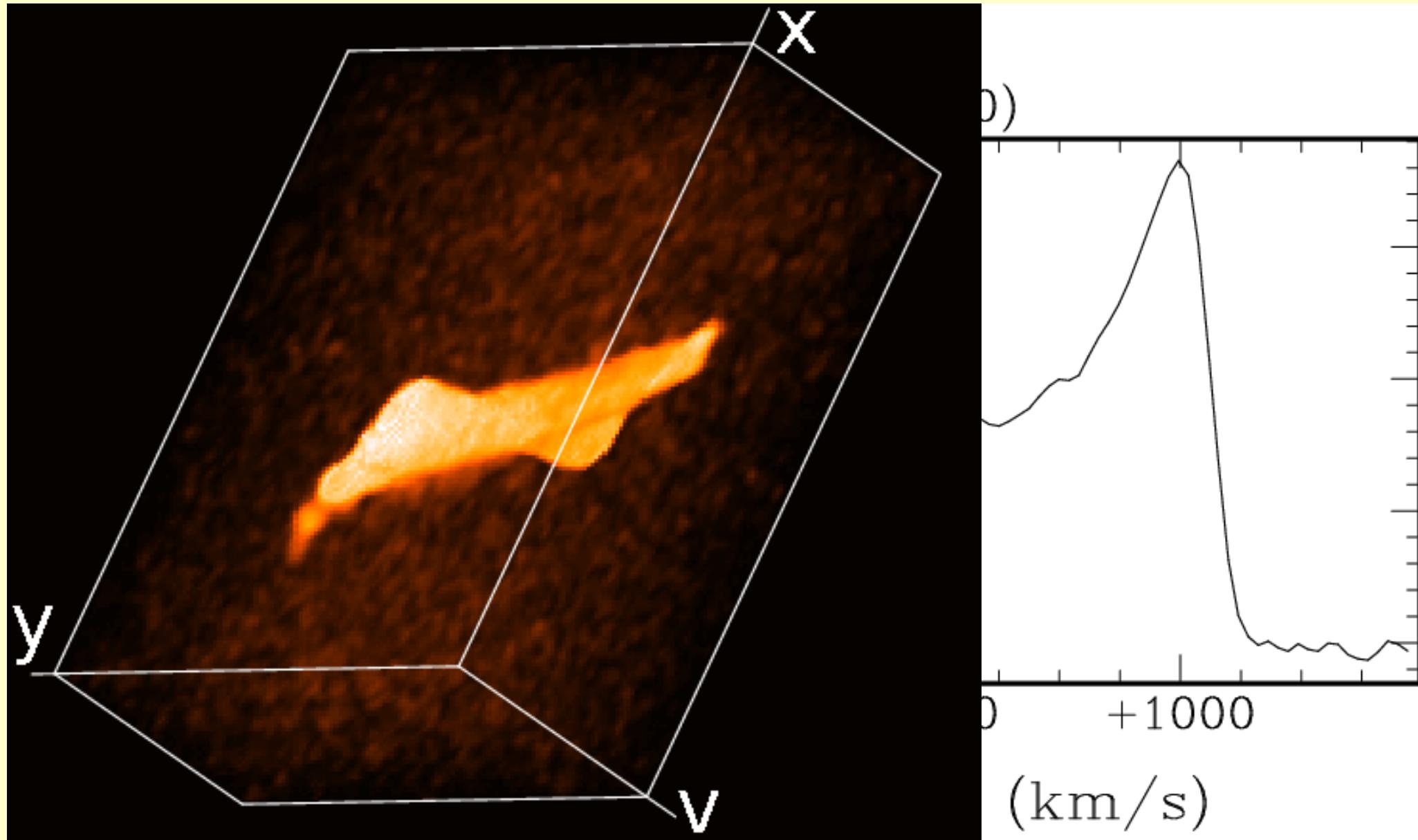
# The task for HI WFSs

- HI in large galaxy surveys tomorrow  
-> interpret 3D data cubes



# Exercising

- HI in galaxy surveys ~~tomorrow~~ **today** -> interpret 3D data cubes



Spatially not well/marginally resolved/shallow:

**Current surveys:** up to  $n \cdot 10^2$  galaxies

(most recent: ATLAS<sup>3D</sup> HI survey, Serra et al. 2011, arXiv 1111.4241)

**WALLABY, WNSHS** ( $n \cdot 10^4$  galaxies)

- Automated parametrisation necessary
- Rotation curves, rotation amplitude
- HI surface density profiles, "aperture photometry"
- Orientation of (inner and outer) disks
- physical status (interacting?)
- morphology (symmetry, clumpiness)
- Environment ("aperture photometry")

Spatially well resolved (medium/deep):

**Current surveys** (e.g. **HALOGAS**, Heald et al. 2011):  $n \cdot 10^{1-2}$  galaxies

**WALLABY, WNSHS** ( $n \cdot 10^3$  galaxies)

...

**MHONGOOSE** (30 galaxies)

- (Semi-)automated fitting necessary
- Higher-order kinematics and morphology
- Anomalous components (tracing accretion)

Spatially not well/marginally resolved/shallow:

**Current surveys:** up to  $n \cdot 10^2$  galaxies

(most recent: ATLAS<sup>3D</sup> HI survey, Serra et al. 2011, arXiv 1111.4241)

**WALLABY, WNSHS** ( $n \cdot 10^4$  galaxies)

- Automated parametrisation necessary ←
- Rotation curves, rotation amplitude ←
- HI surface density profiles, “aperture photometry” ←
- Orientation of (inner and outer) disks ←
- physical status (interacting?)
- morphology (symmetry, clumpiness)
- Environment (“aperture photometry”)

Spatially well resolved (medium/deep):

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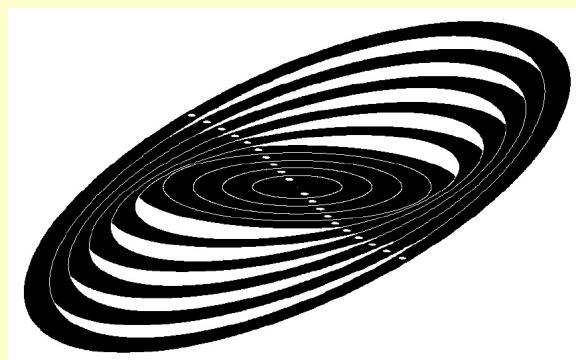
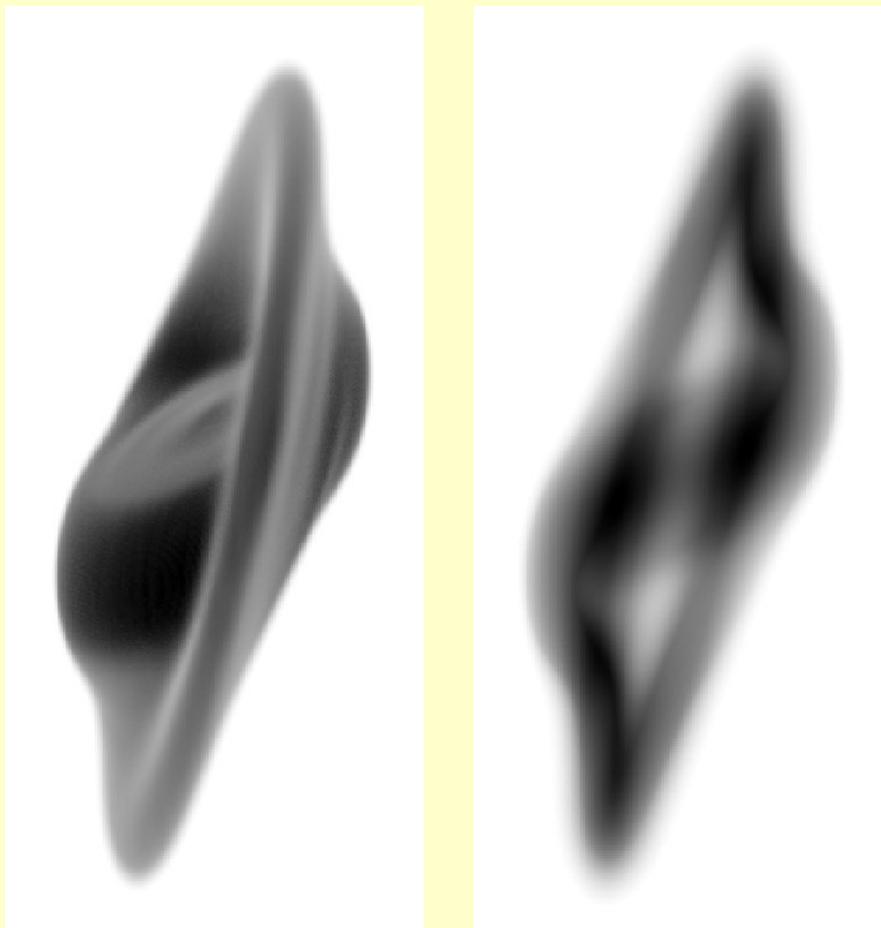
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...

**MHONGOOSE** (30 galaxies)

- (Semi-)automated fitting necessary ←
- Higher-order kinematics and morphology ←
- Anomalous components (tracing accretion) ←

# The tilted-ring model



Tilted-Ring-Model  
(Rogstad et al. 1974):

parametrise rings at different radii by

- two orientation parameters (inclination, position angle)
- central position
- surface brightness (thickness)
- rotation velocity

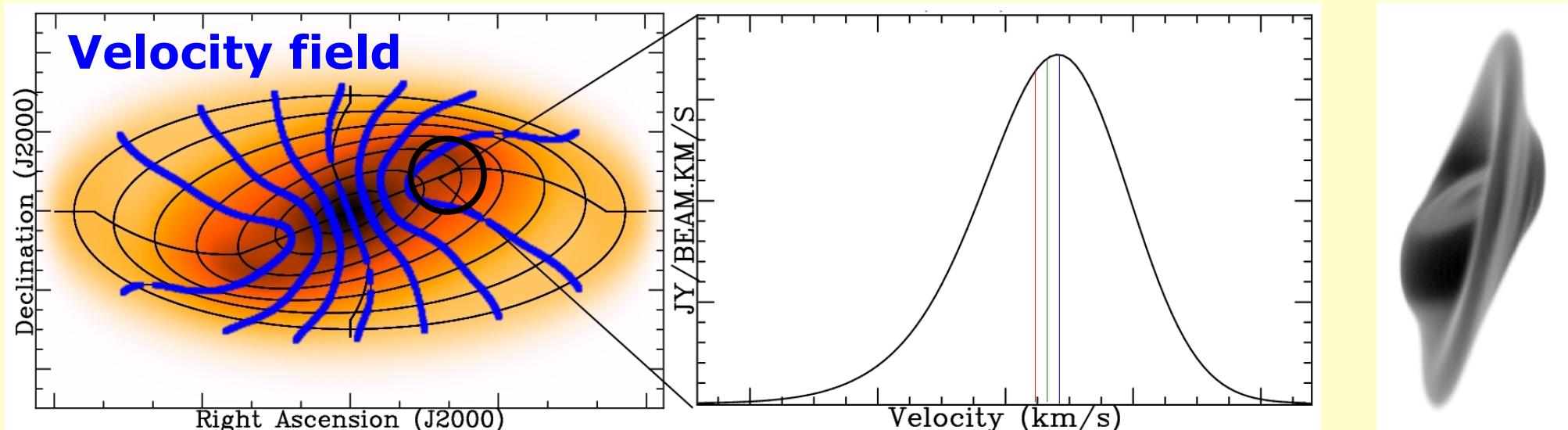
García-Ruiz 2001

Method 1: Fit to a “velocity-field” (Begemann 1987, Schoenmakers 2001, ..., Spekkens 2007)

- Velocity field is impossible to be constructed if the line-of sight crosses the disk twice, does not reproduce any thickness

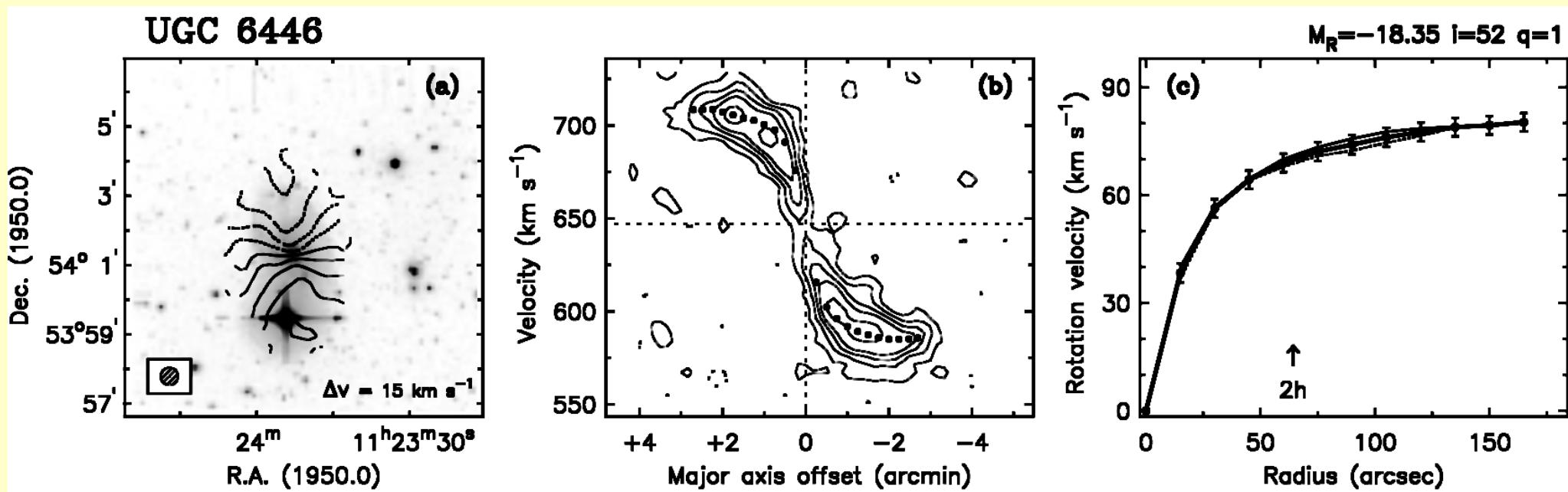
Method 2: direct fit to the data cube (“Galmod”, (Corbelli & Schneider 1997, Józsa et al. 2007, Wiegert et al. 2006))

- Slow



# Tilted-ring modelling: velocity field **ASTRON**

- Works very well for data cubes with **high spatial resolution** (de Blok et al. 2008)
- Fast
- **Beam smearing** stabilises solutions at high resolution
- For lower resolution (compared to the extent of the disk) **beam smearing** becomes an issue.
- filtering of anomalous components

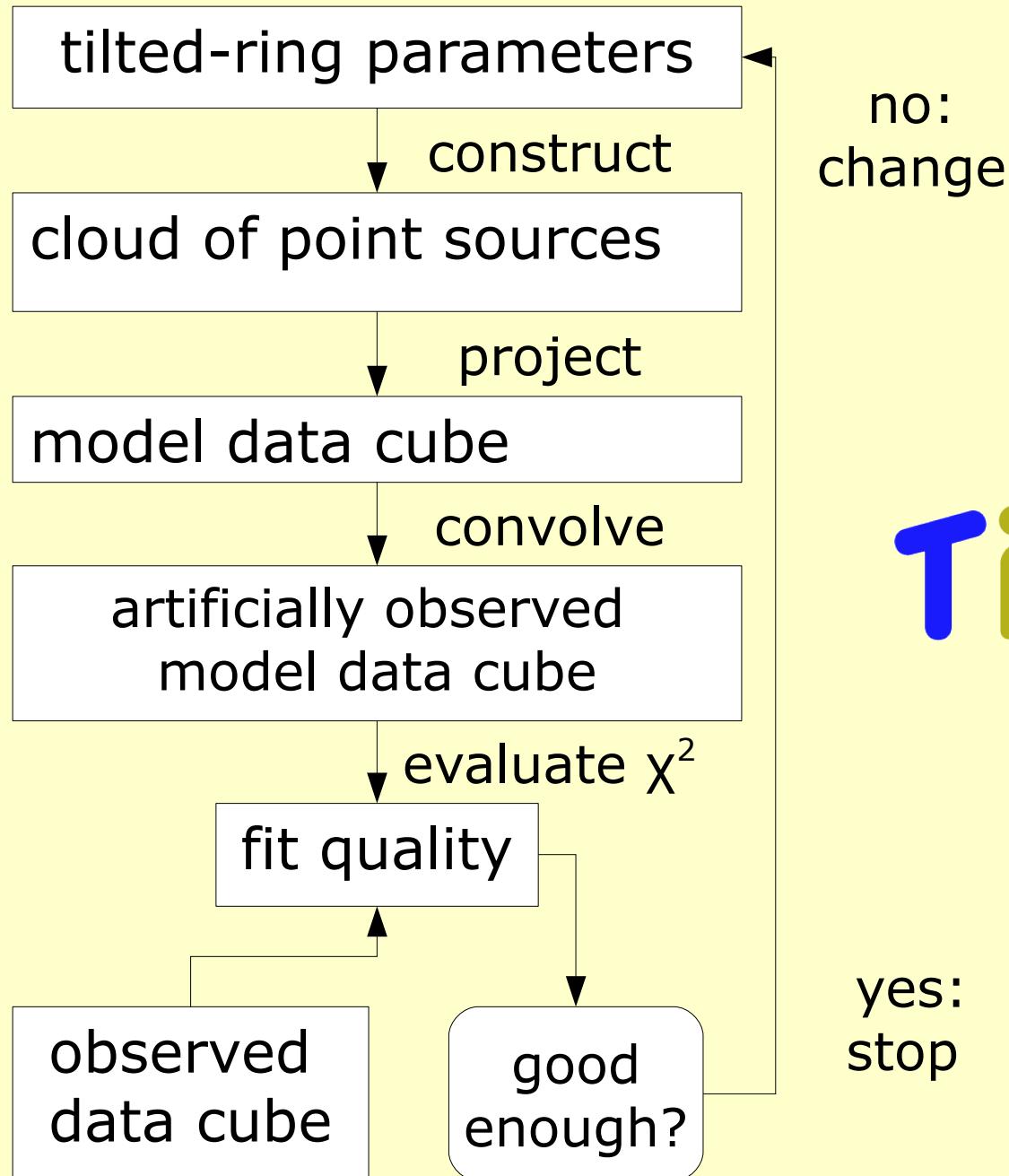


Swaters 1999

- WALLABY/WNSHS parametrisation group (**Spekkens**, Oh, Elson, de Blok, Westmeier, Józsa)
- Current focus: automated tilted-ring modelling
- Strategy:
  - Use velocity field method for moderately inclined and well resolved galaxies  
→ **DiskFit/velfit** (Spekkens 2007, Spekkens et al.)  
<http://www.physics.rutgers.edu/~spekkens/velfit/>
  - Use direct fit method for highly inclined or poorly resolved galaxies  
→ **TiRiFiC** (Józsa et al. 2007)  
<http://www.astron.nl/~jozsa/tirific/> (v2 in progress)

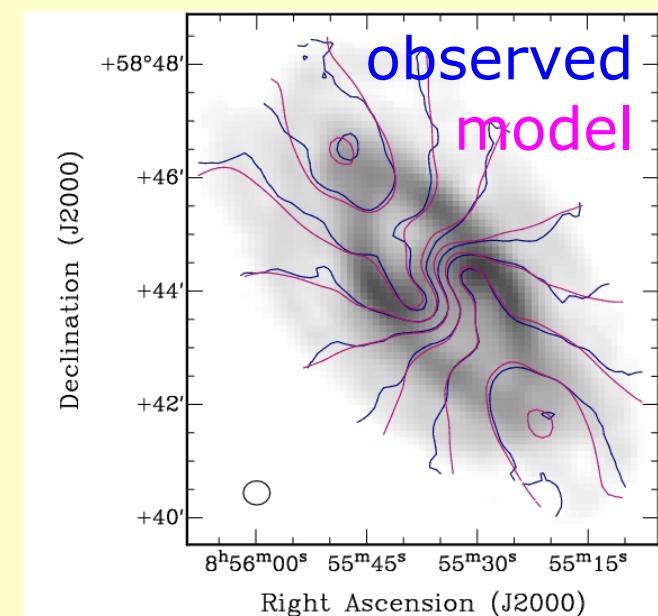
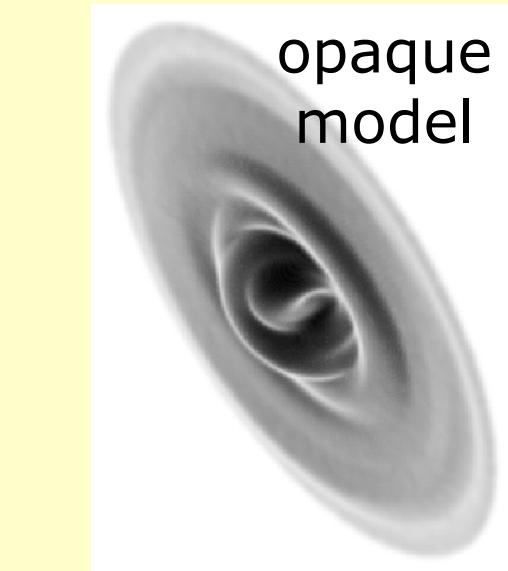
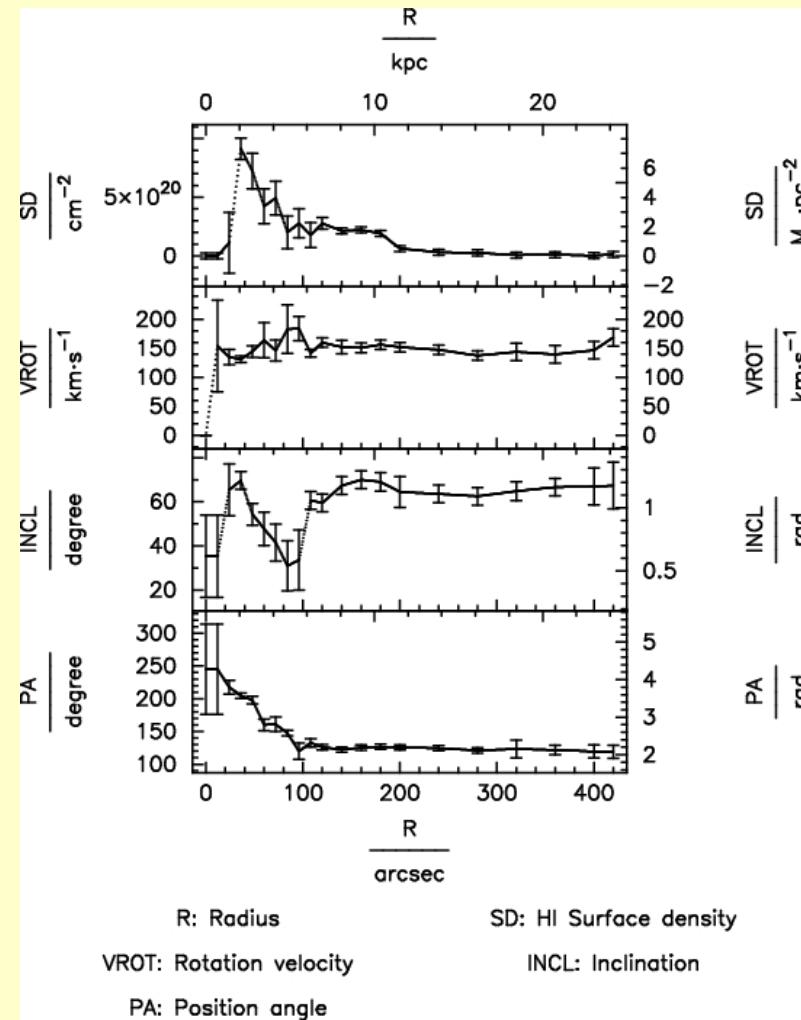
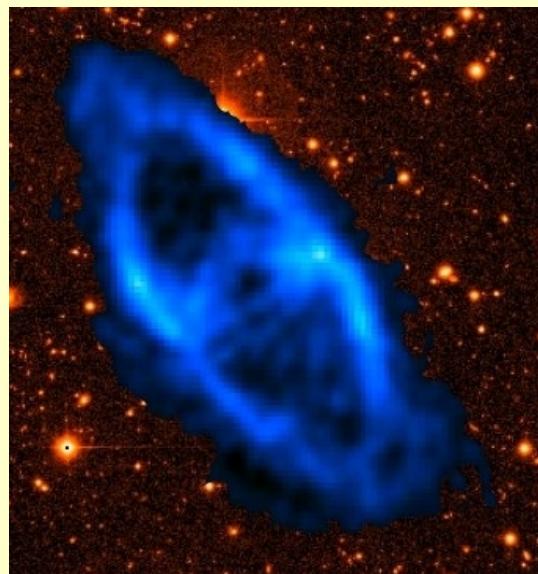
# TiRiFiC basic functionality

ASTRON



# NGC 2685

ASTRON



- Modelling of some (even complicated) galaxies possible

- Enable complex tilted-ring modelling for well-resolved galaxies
- Speed-up
- Publication (online description, download)
- User-interface
- Establish fitting pipeline for moderately and marginally resolved galaxies
  - Script TiRiFiC to enable automated quality control
  - Find optimal  $\chi^2$ -minimisation scheme
  - Regularise parametrisation
  - Alternative parametrisation schemes to reduce number of parameters

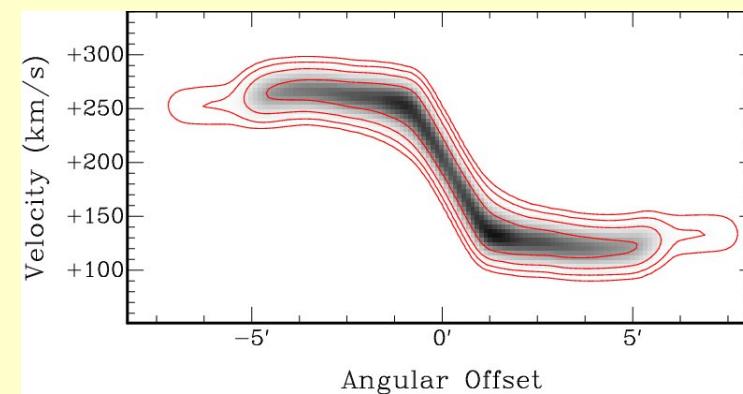
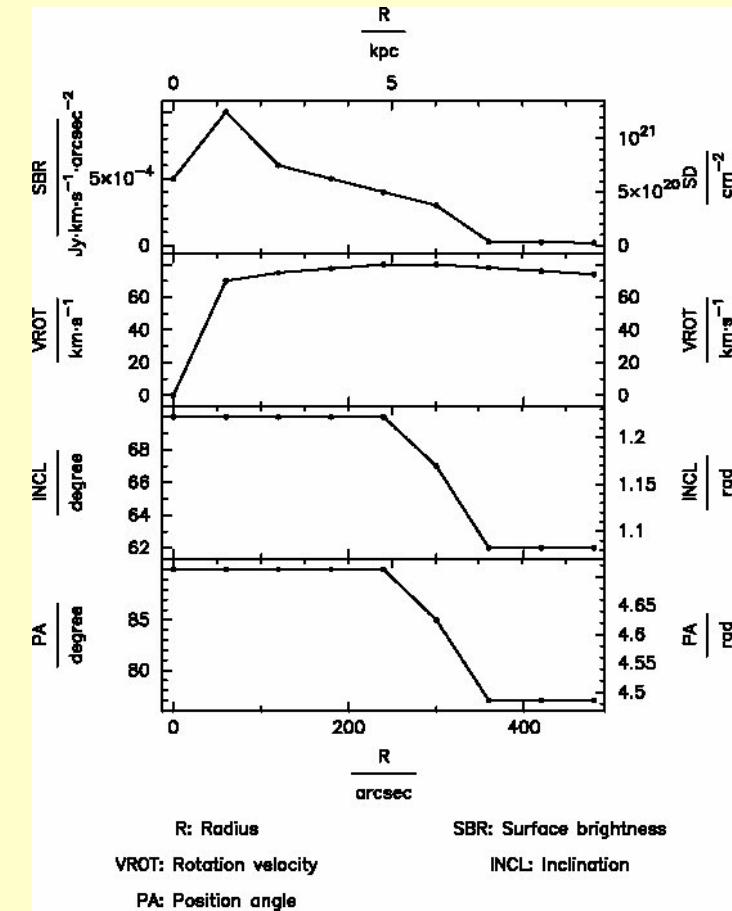
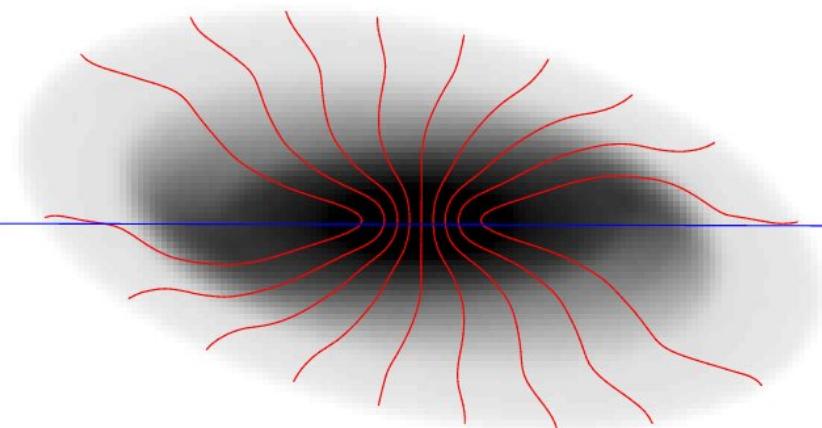
# TiRiFiC parameters

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## Basic TRM (galmod)

for each radius

- Surface brightness SBR ( $\text{Jy km s}^{-1} \text{arcsec}^{-2}$ )
- Rotation velocity VROT ( $\text{km s}^{-1}$ )
- Position angle PA ( $^{\circ}$ )
- Inclination INCL ( $^{\circ}$ )
- Scale height Z0 ( $\text{arcsec}$ )
- Dispersion LTYPE (Gaussian, sech<sup>2</sup>, constant)
- Ring centre RA SDIS ( $\text{km s}^{-1}$ )
- Ring centre Dec XPOS ( $^{\circ}$ )
- Systemic velocity YPOS ( $^{\circ}$ )
- Systemic velocity VSYS ( $\text{km s}^{-1}$ )

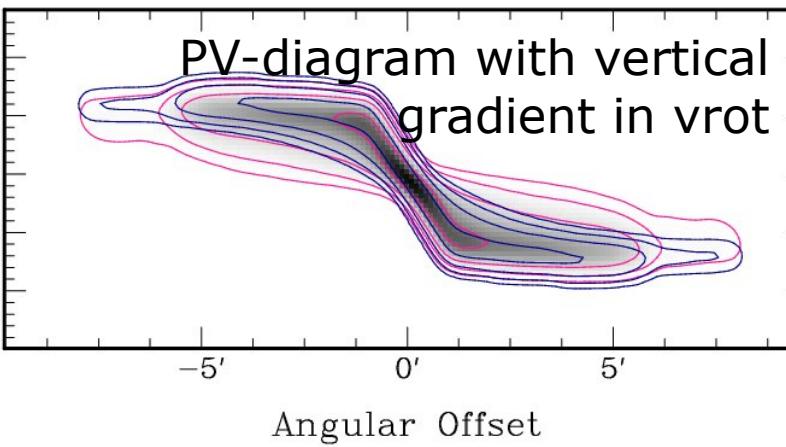
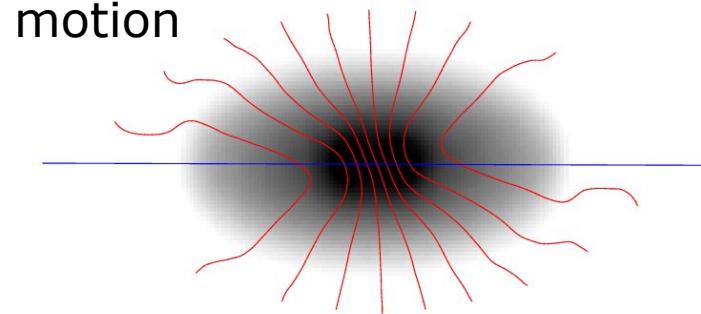


## Global symmetric motion and gradients

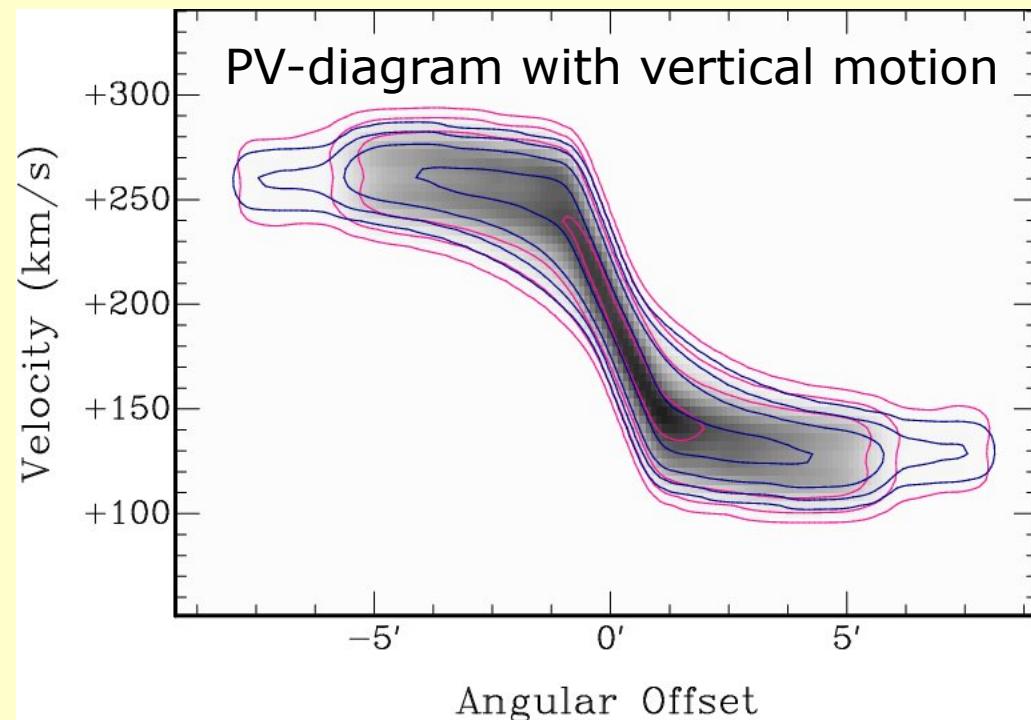
- Radial motion VRAD ( $\text{km s}^{-1}$ )
- Vertical motion VVER ( $\text{km s}^{-1}$ )
- Vertical gradients DVRO (VR<sub>OT</sub>,  $\text{km s}^{-1} \text{arcsec}^{-1}$ )  
DVRA (VRAD,  $\text{km s}^{-1} \text{arcsec}^{-1}$ )  
DVVE (VVER,  $\text{km s}^{-1} \text{arcsec}^{-1}$ )

**Example: flat disk,  $i = 60^\circ$ , thick disk**

Velocity field with radial motion

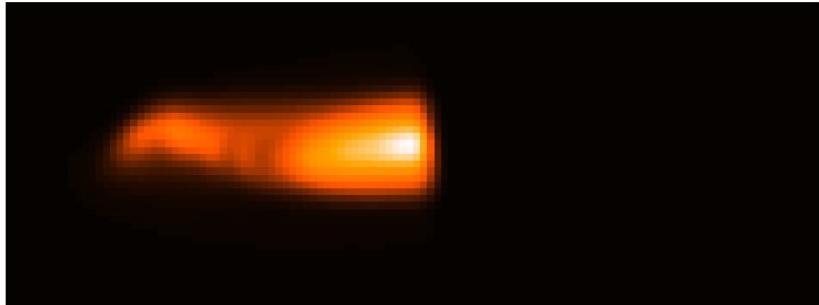


PV-diagram with vertical motion

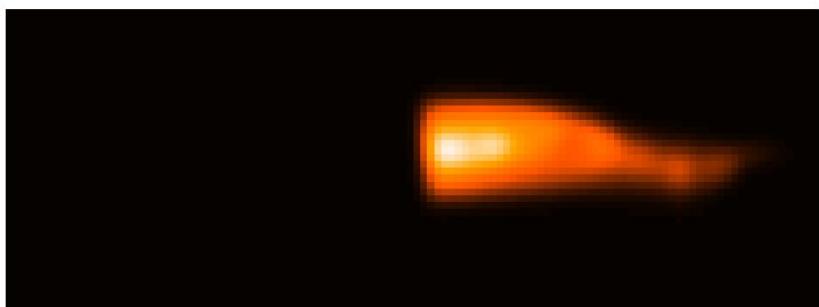


# Asymmetric galaxies

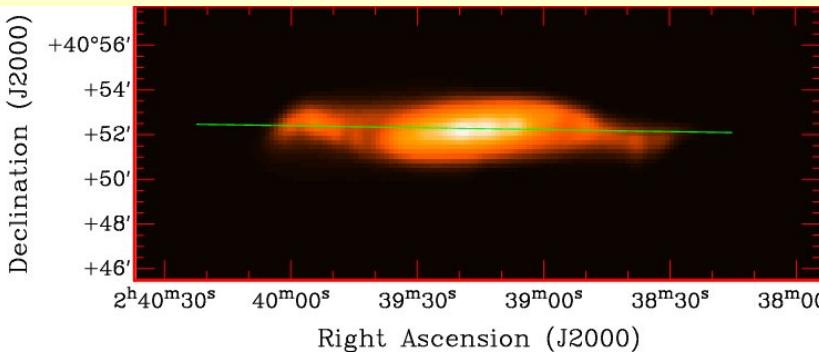
- asymmetric disks, modelled as two half disks



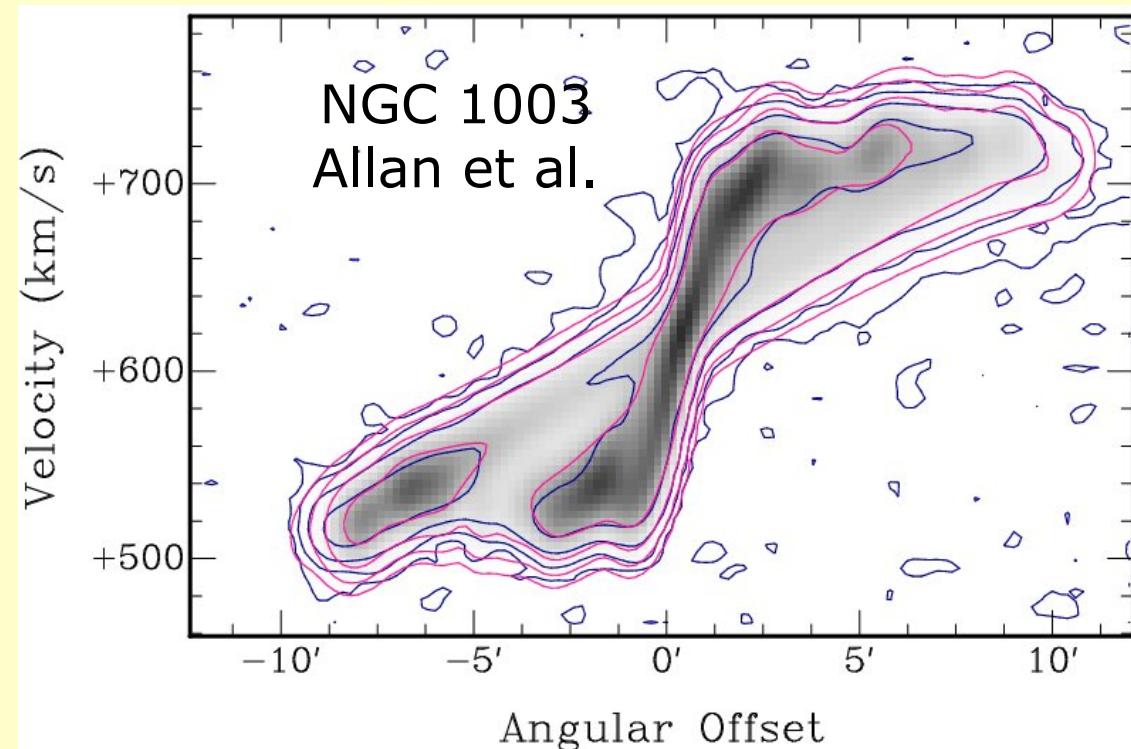
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Heald et al. 2011



- Enable complex tilted-ring modelling for well-resolved galaxies
  - Basic tilted-ring model (galmod)
  - Radial, vertical motion, vertical gradients in velocity
  - Higher-order warps
  - Global shifts along projected axes
  - Harmonics in surface brightness
  - Bar-like and spiral distortions in surface brightness
  - Harmonic terms in velocity
  - Multiple disks and partial disks

**Use with great care!**

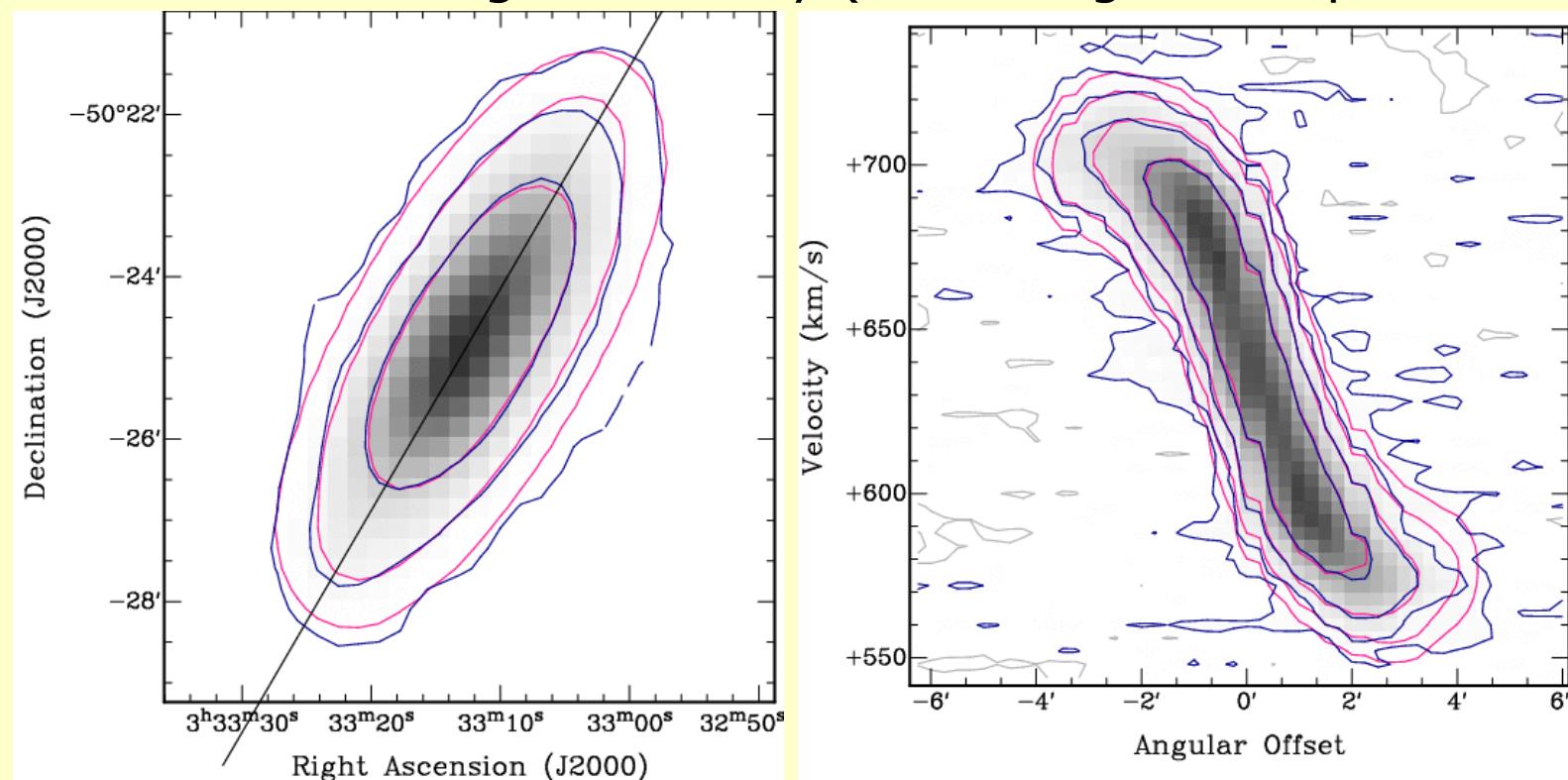
**Btw....**

**TiRiFiC can be used without automated fitting (faster than alternatives)**

- Enable complex tilted-ring modelling for well-resolved galaxies ✓
- Speed-up:
  - Optimisation of I/O
  - Internally adjusting methods to demand
  - Partial model generation (re-calculate only what's needed)
  - Crude parallelisation (multi-threading with OMP, thanks to Hyun-Jin Bae, Yonsei Univ.)

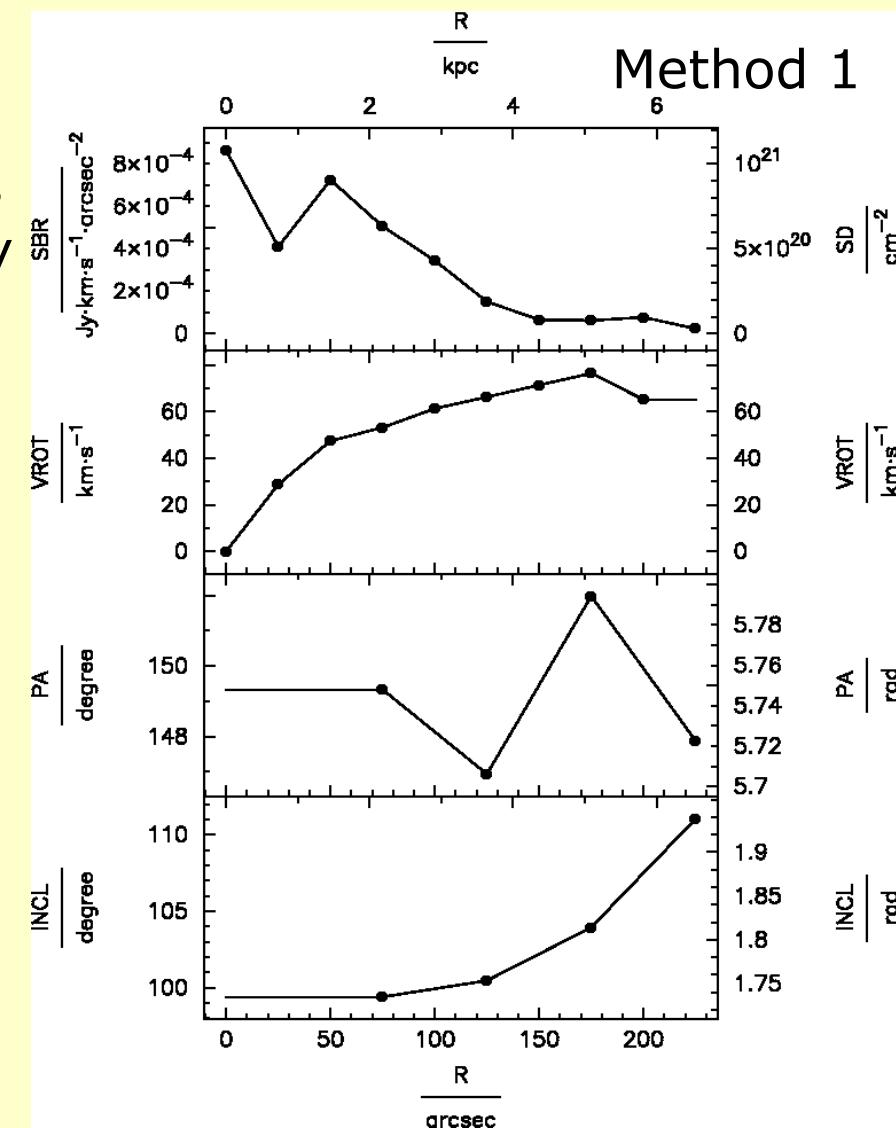
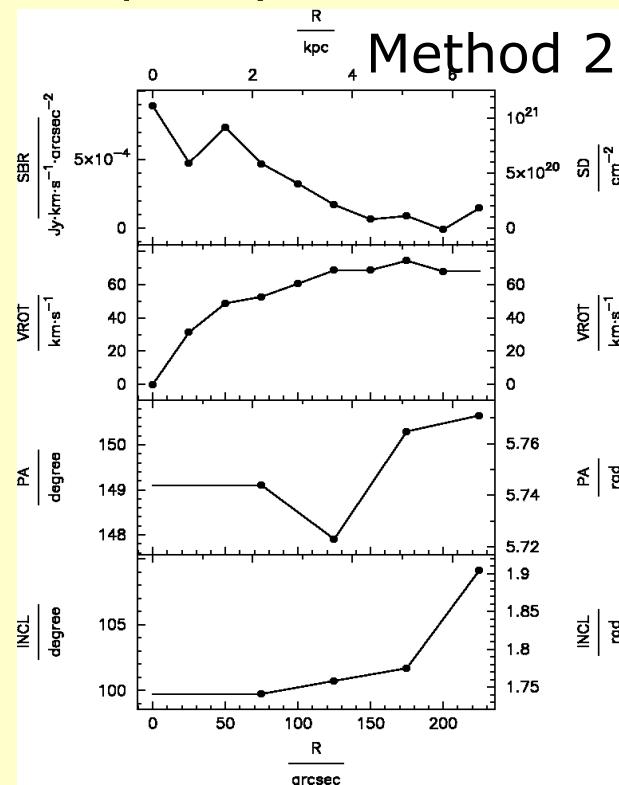
IC 1959 (ATCA, LVHIS, Koribalski et al.), minimisation method 1

- 500 kB data cube
- corresponding to  $D_{\text{HI}} = 7.5 \text{ arcmin}$  @ HPBW = 60 arcsec
- 34 parameters, simple TRM
- 8 Intel Xeon 2.4 GHz cores
- 1 fitting run (1000 models) 2 minutes (7 minutes 1 core)
- 5 re-iterations -> 10 minutes/galaxy (35 minutes 1 core)
- 2 8-core computers -> 250-300 galaxies/day (660 using 1 core per gal.)
- **BUT** no quality control, no errors



IC 1959 (ATCA, LVHIS, Koribalski et al.), minimisation method 2

- 500 kB data cube
- corresponding to  $D_{\text{HI}} = 7.5 \text{ arcmin}$  @ HPBW = 60 arcsec
- 34 parameters, simple TRM
- 1 Intel Xeon 2.4 GHz core
- 2 fitting runs (2500 models) 50 minutes
- 2 8-core computers -> 450 galaxies/day
- **BUT** no quality control, no errors



- Enable complex tilted-ring modelling for well-resolved galaxies ✓
- Speed-up ✓
- Publication (online description, download)
- User-interface
- Establish fitting pipeline for moderately and marginally resolved galaxies
  - Script TiRiFiC to enable automated quality control
  - Find optimal  $\chi^2$ -minimisation scheme
  - Regularise parametrisation
  - Alternative parametrisation schemes to reduce number of parameters

- Enable complex tilted-ring modelling for well-resolved galaxies ✓
- Speed-up ✓
- Publication (online description, download):
  - 90% finished



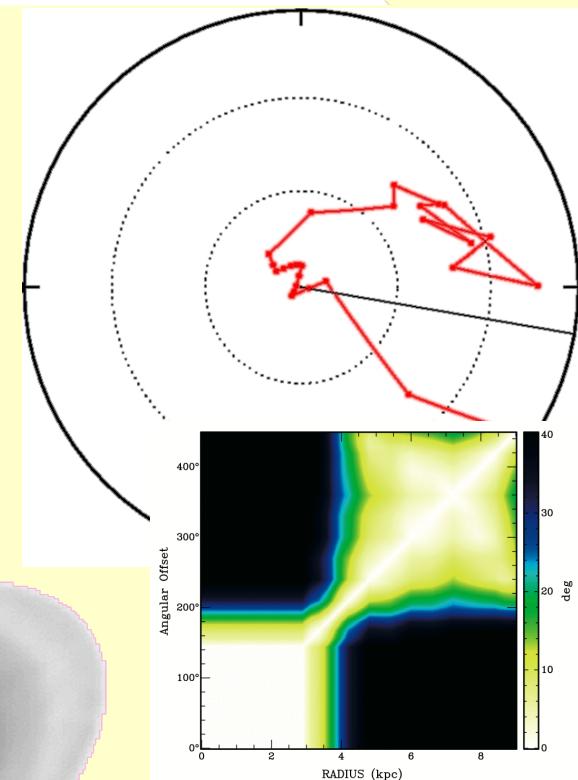
**<http://www.astron.nl/~jozsa/tirific>**  
(currently: contact info)

- Enable complex tilted-ring modelling for well-resolved galaxies ✓
- Speed-up ✓
- Publication (online description, download): not quite, but soon ✓
- User-interface ✓
- Establish fitting pipeline for moderately and marginally resolved galaxies
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# TiRiFiC user interface

ASTRON

- GIPSY (<http://www.astro.rug.nl/~gipsy/>)
  - Most complete package for HI analysis
  - Installation time consuming (esp. Mac)
  - Complex data structure
  - TiRiFiC standalone planned, but low priority
- Several output possibilities (tables, plots, data sets, verification in progress)
- GUI possible, Python prototype, but low priority

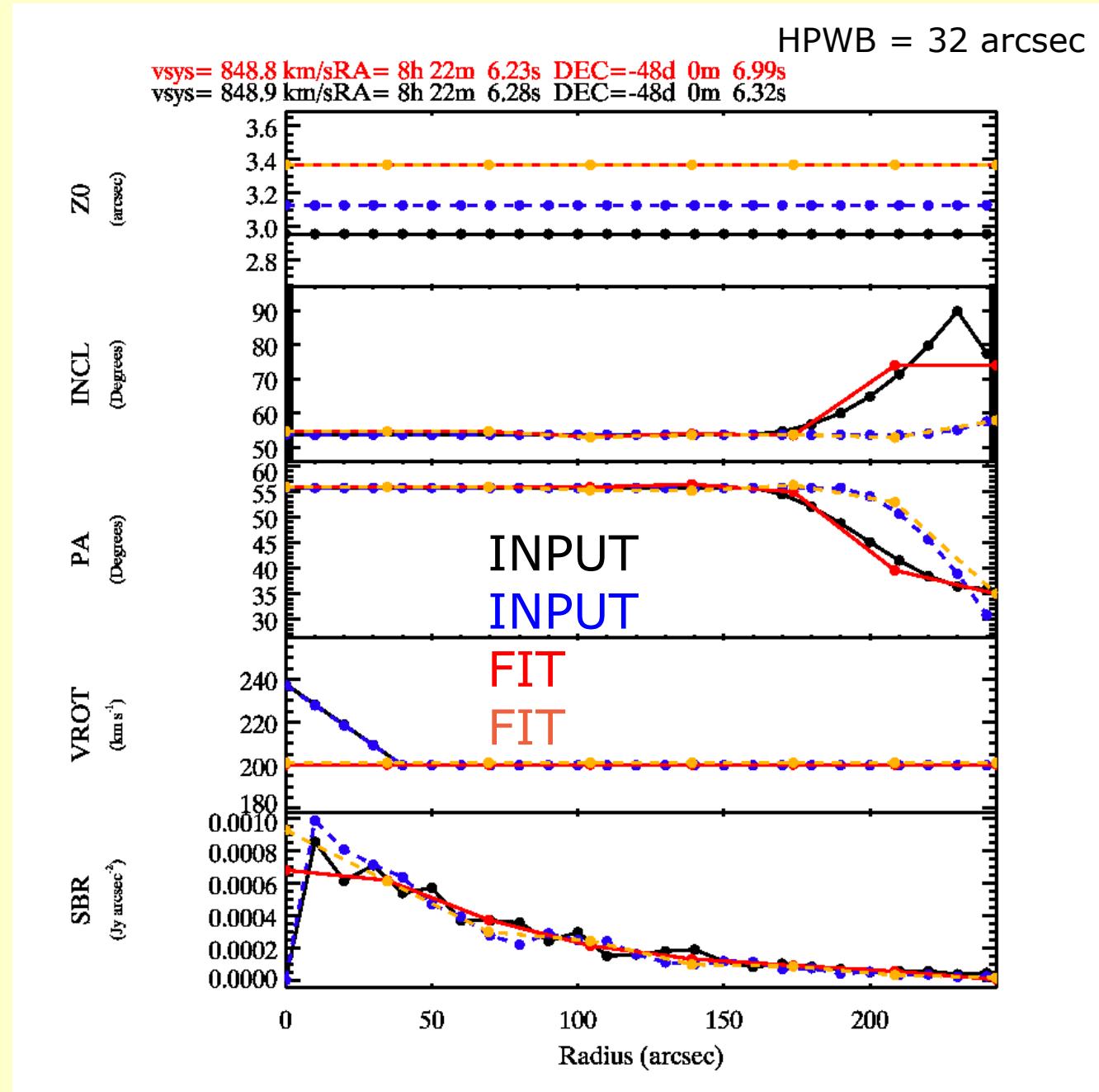


- Enable complex tilted-ring modelling for well-resolved galaxies ✓
- Speed-up ✓
- Publication (online description, download): not quite, but soon ✓
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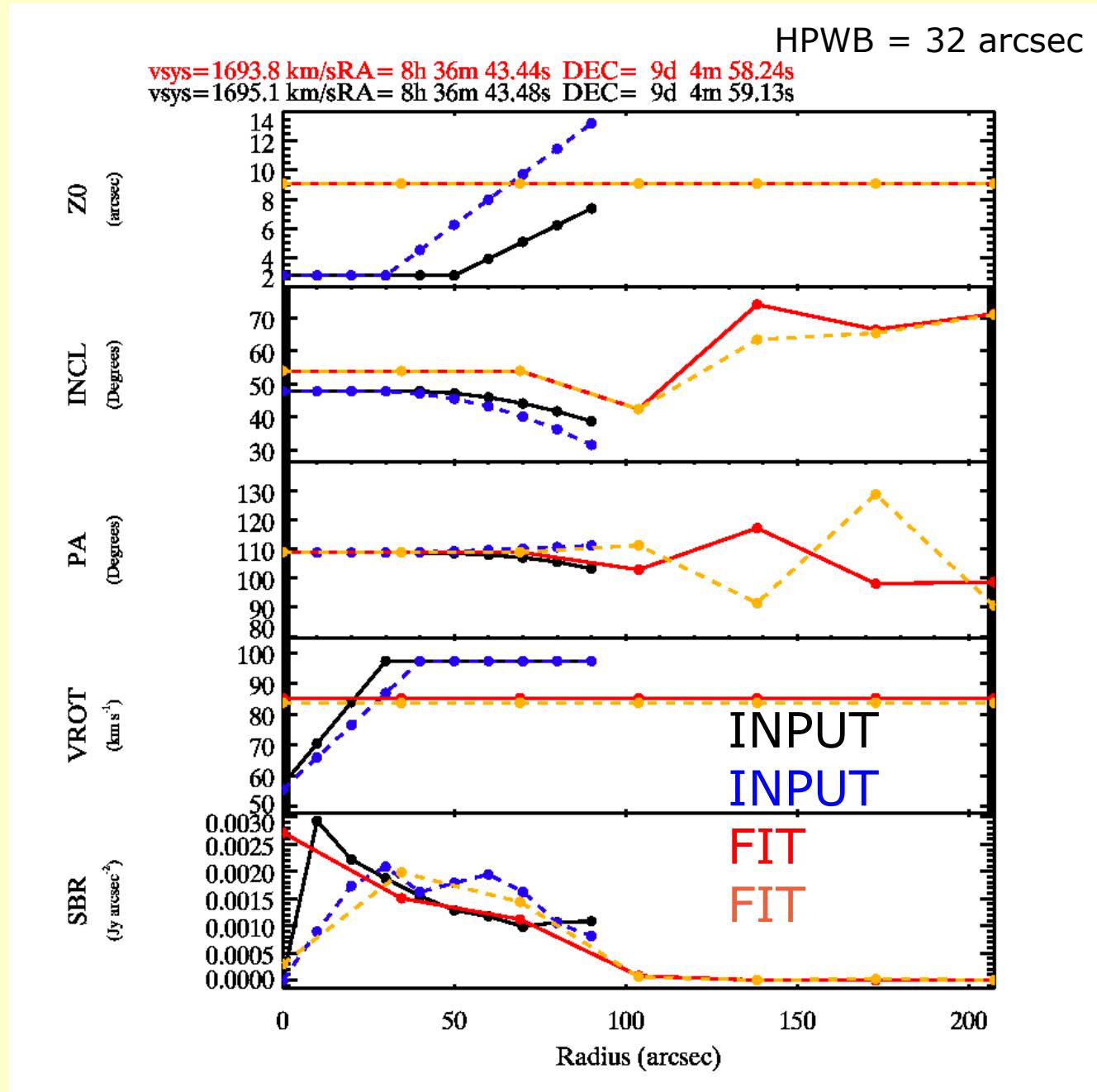
# Scripting TiRiFiC

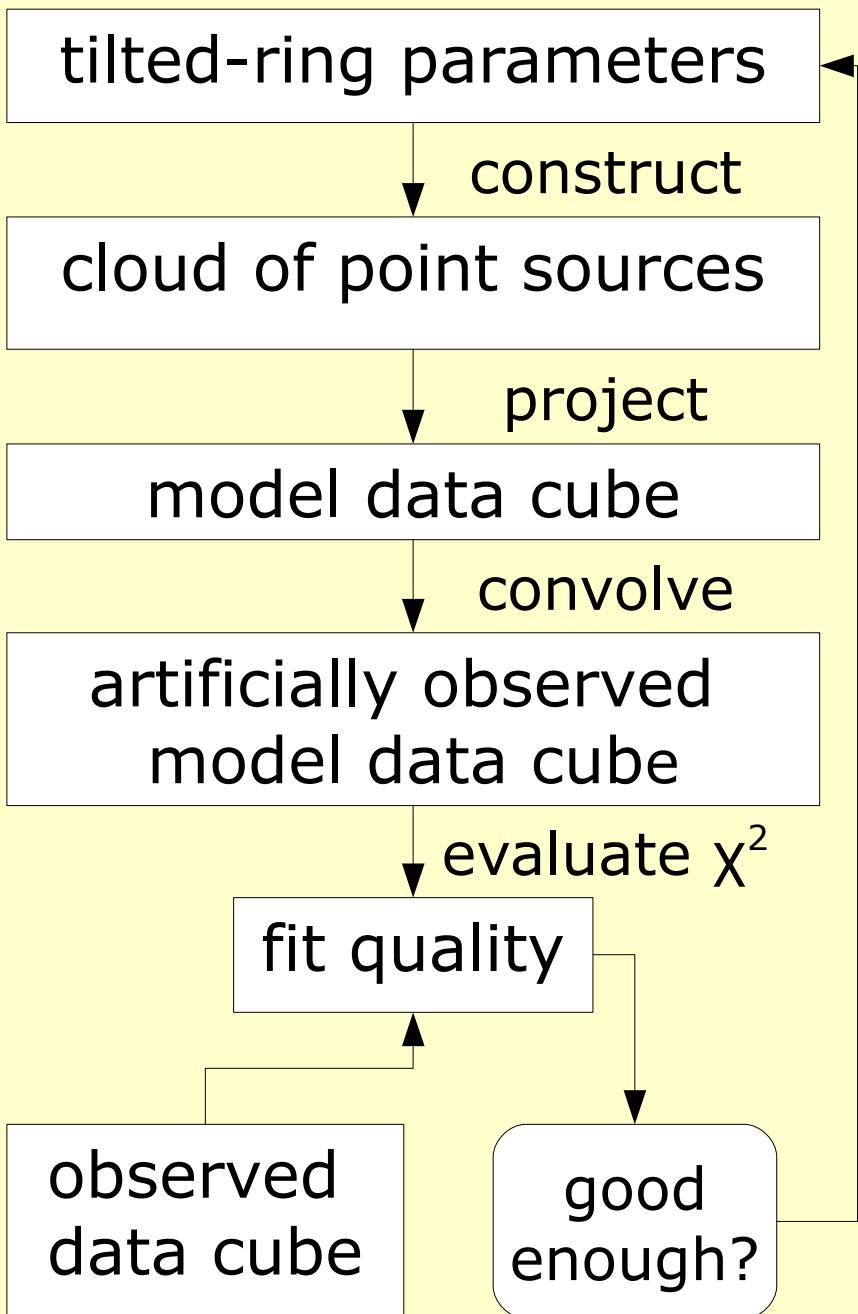
ASTRON

- Currently, TiRiFiC requires human control
- Implementation of automated wrapper in progress (by Peter Kamphuis, nearly ATNF)
- Combination of TiRiFiC with several GIPSY routines
- Final fit always with TiRiFiC
- No error-bars yet
- On average 5 runs
- ➔ 1h/galaxy (again)



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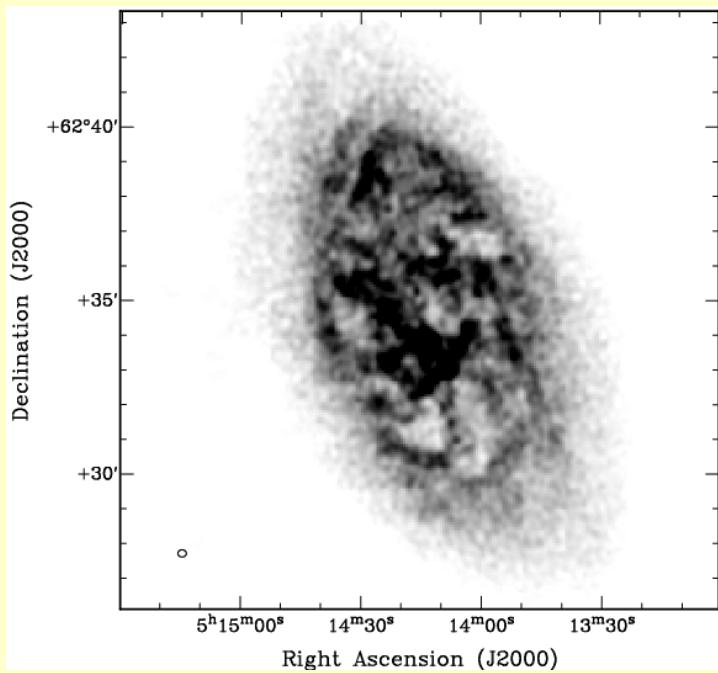




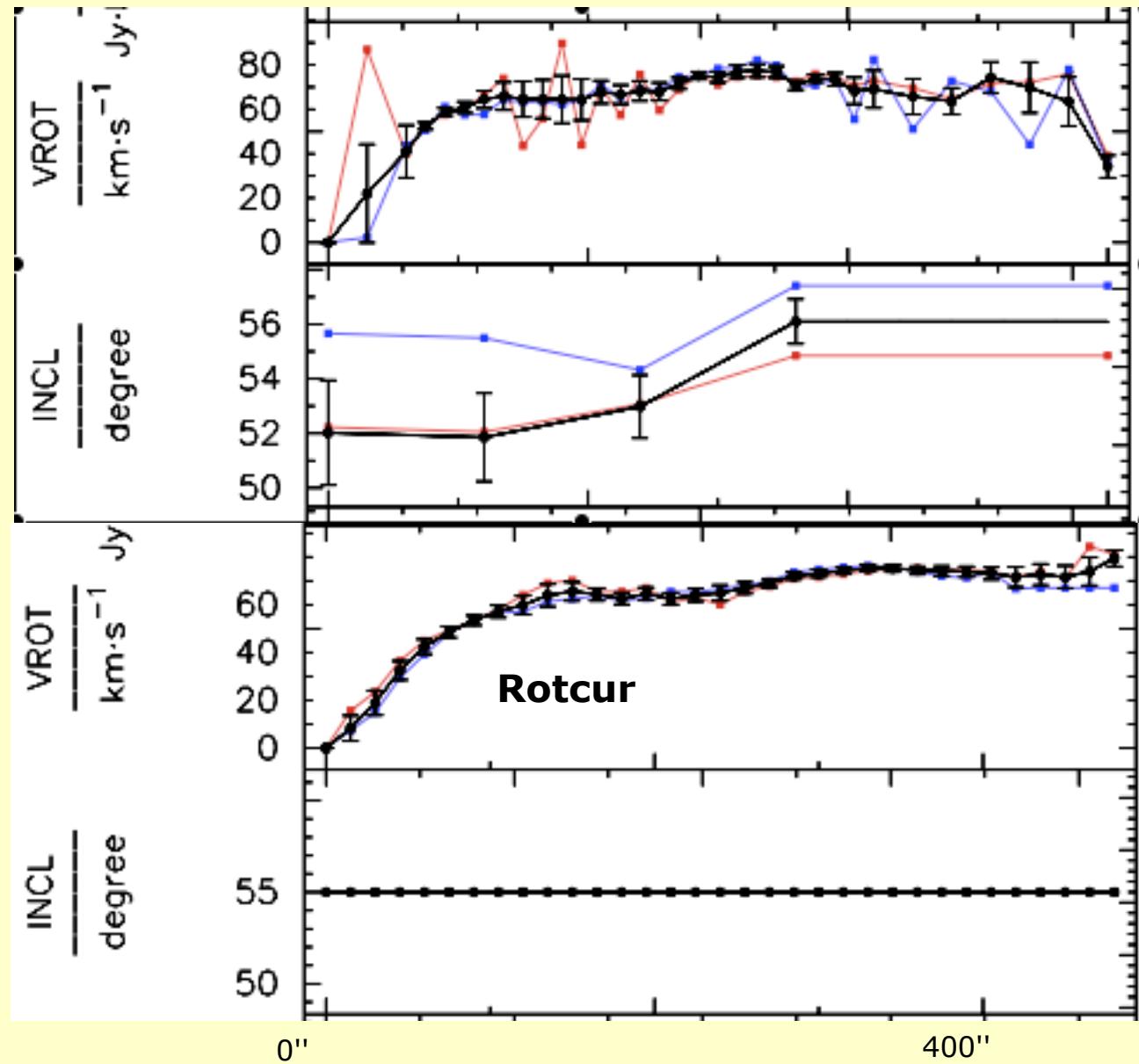
- $\chi^2$ -minimisation methods:
  - **Metropolis** (not successful)
  - **Downhill simplex** (not very successful, only for small number of parameters)
  - Simple one-by-one **golden-section** nesting:
    - i) Optimise the first parameter of the first ring
    - ii) Optimise the same parameter of the next ring
    - n) Optimise the last parameter of the last ring
    - n+1) return to start

- Establish fitting pipeline for moderately and marginally resolved galaxies
  - Script TiRiFiC to enable automated quality control
  - Starting, promising, but highly depending on
  - Finding the optimal  $\chi^2$ -minimisation scheme
    - HUGE room for improvement (**Suggestions welcome!**):
      - Minimiser (MCMC? Genetic?)
      - Errors

- Enable complex tilted-ring modelling for well-resolved galaxies ✓
- Speed-up ✓
- Publication (online description, download): not quite, but soon ✓
- User-interface ✓
- Establish fitting pipeline for moderately and marginally resolved galaxies
  - Script TiRiFiC to enable automated quality control ✓
  - Find optimal  $\chi^2$ -minimisation scheme XXX
  - Regularise parametrisation
  - Alternative parametrisation schemes to reduce number of parameters



**WSRT HI UGCA 105 (Schmidt et al.)**



- Unreliable fit results for irregular disks (not at low resolution)
- Regularisation
- So far not successful

- Enable complex tilted-ring modelling for well-resolved galaxies ✓
- Speed-up ✓
- Publication (online description, download): not quite, but soon ✓
- User-interface ✓
- Establish fitting pipeline for moderately and marginally resolved galaxies
  - Script TiRiFiC to enable automated quality control ✓
  - Find optimal  $\chi^2$ -minimisation scheme XXX
  - Regularise parametrisation XXX
  - If failing -> alternative parametrisation schemes to reduce number of parameters

- TiRiFiC is a tool to fit an extended tilted-ring model to data cubes
- Ready as a semi-automated application
- Possible part of WALLABY & WNSHS parametrisation pipeline
- First pipeline implementation promising
- Better minimisation scheme needed (**suggestions?**)
- Regularisation and alternative parametrisation to be implemented

