


# Galaxy Evolution in the Mock Galaxy Factory

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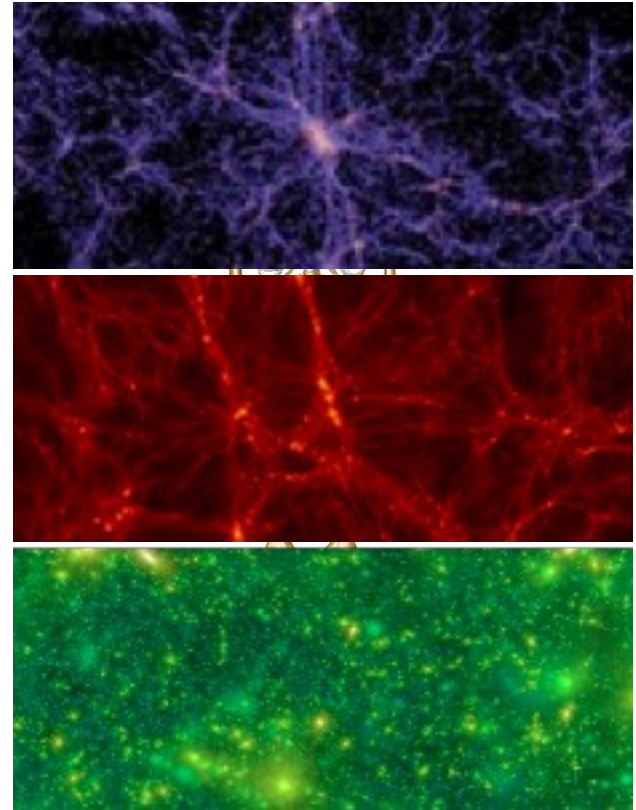
- 
- Mock catalogues production
  - Unified models of AGN in SAMs

# Four Steps to Produce a Mock Catalogue

1. Run a Dark Matter simulation
2. Apply a Galaxy Model
3. Apply Stellar Population Synthesis model
4. Remap and/or randomise simulation volume

# 1. Dark Matter Simulations

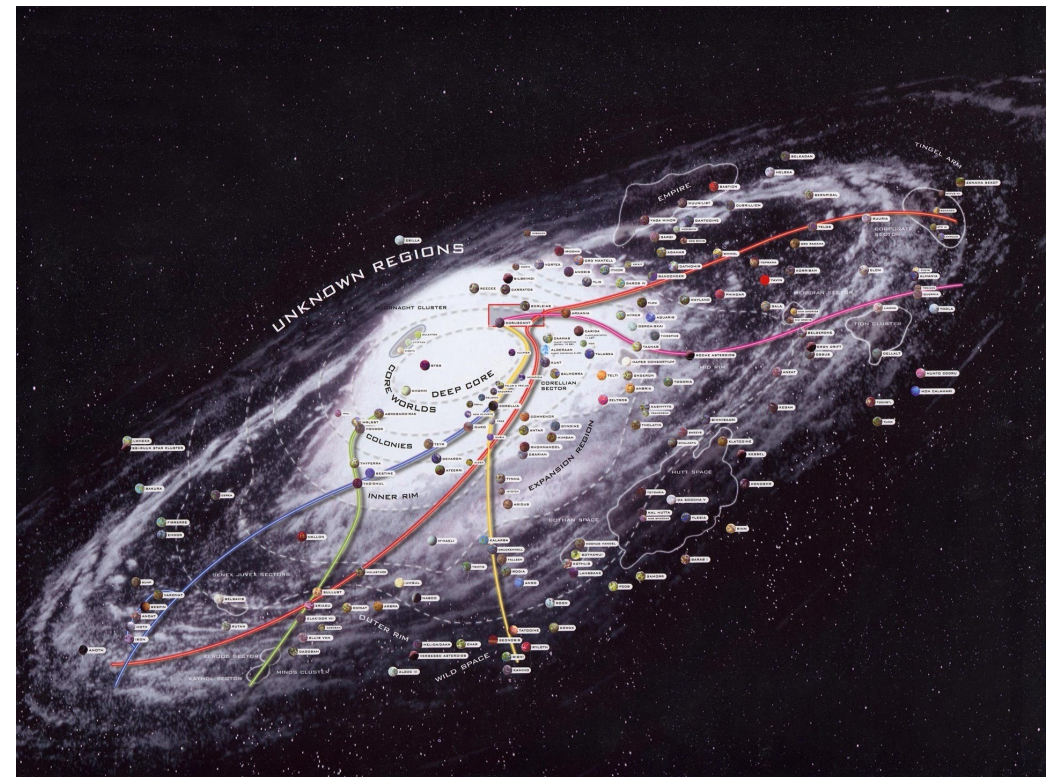
- Millennium (Springel et al. 2005)
- Bolshoi (Klypin et al. 2010)
- GiggleZ (Poole et al. in prep.)



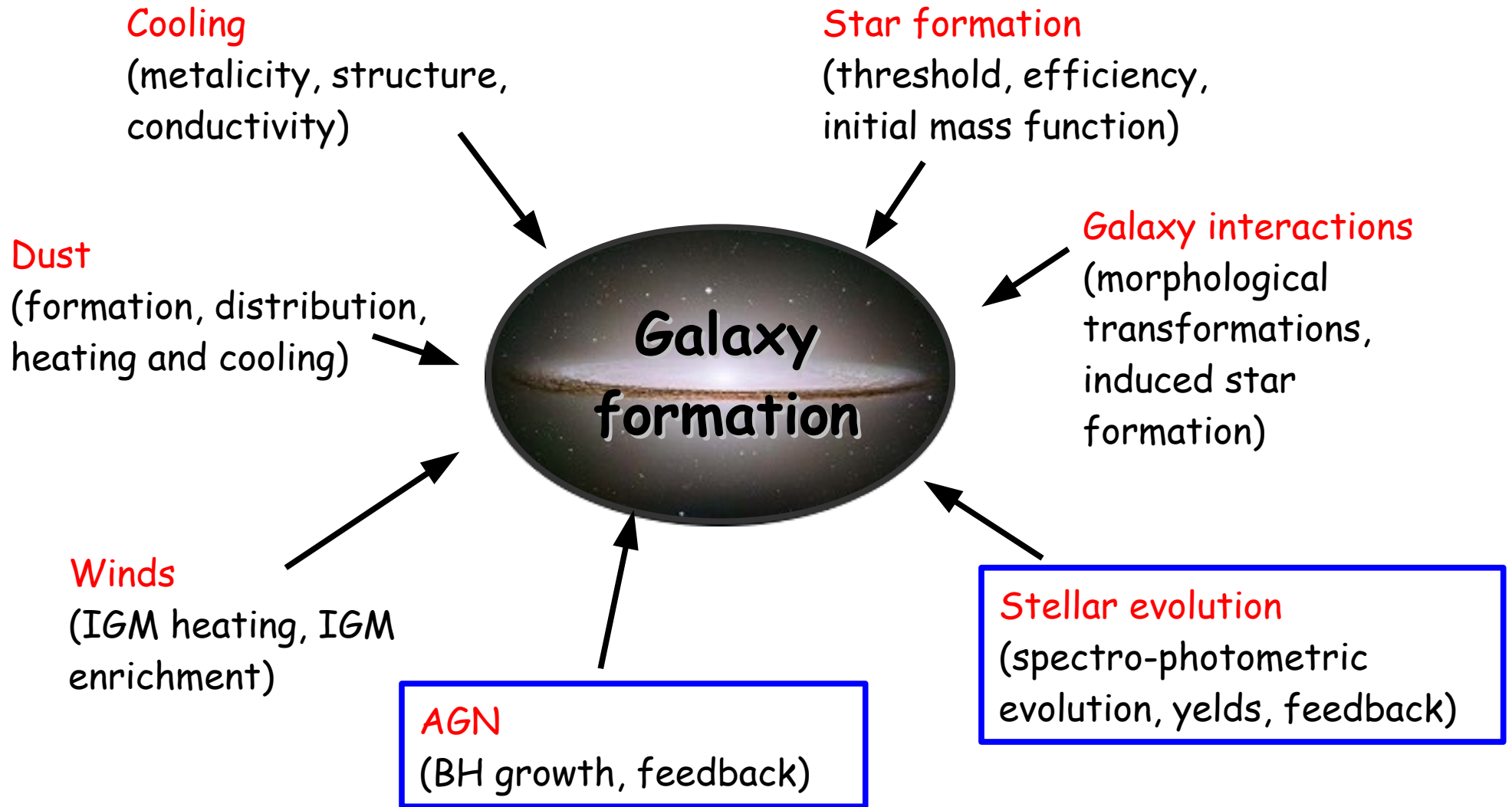


## 2. Galaxy Model

- Croton et al. 2006
- Somerville et al. 2008
- Benson et al. 2011



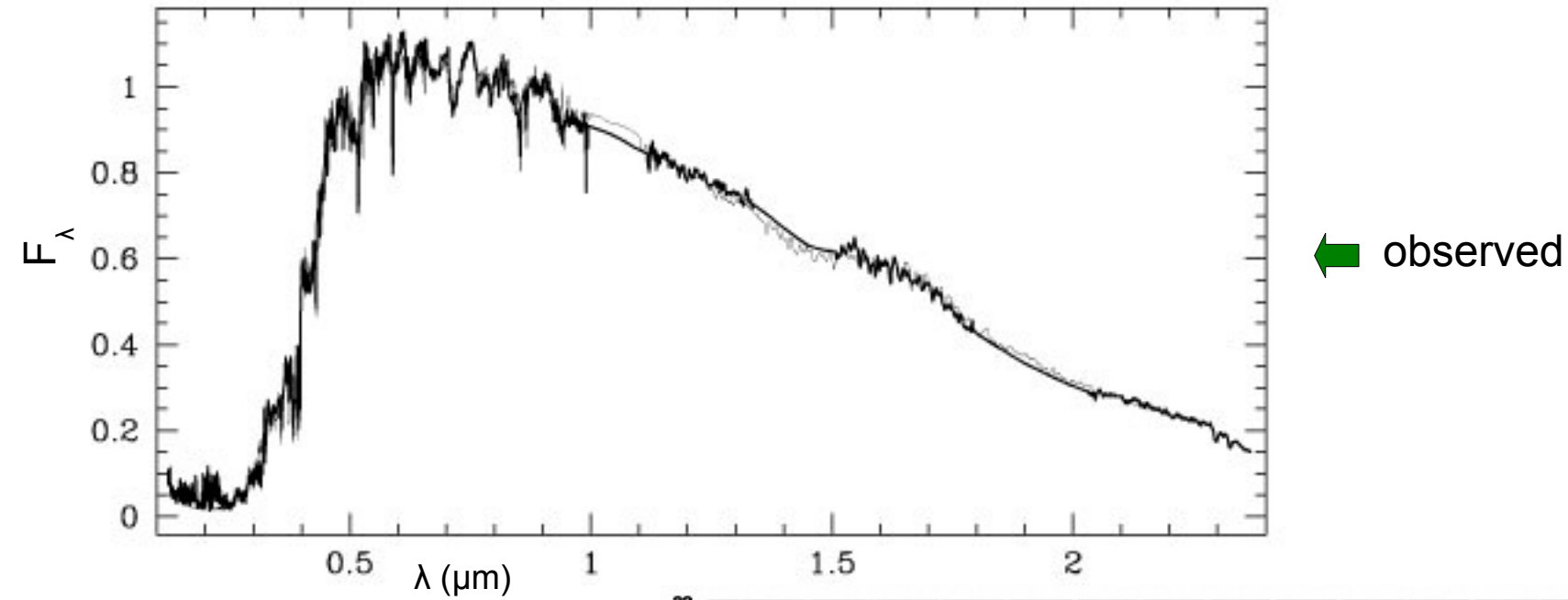
# Semi Analytical Galaxy Evolution (SAGE)



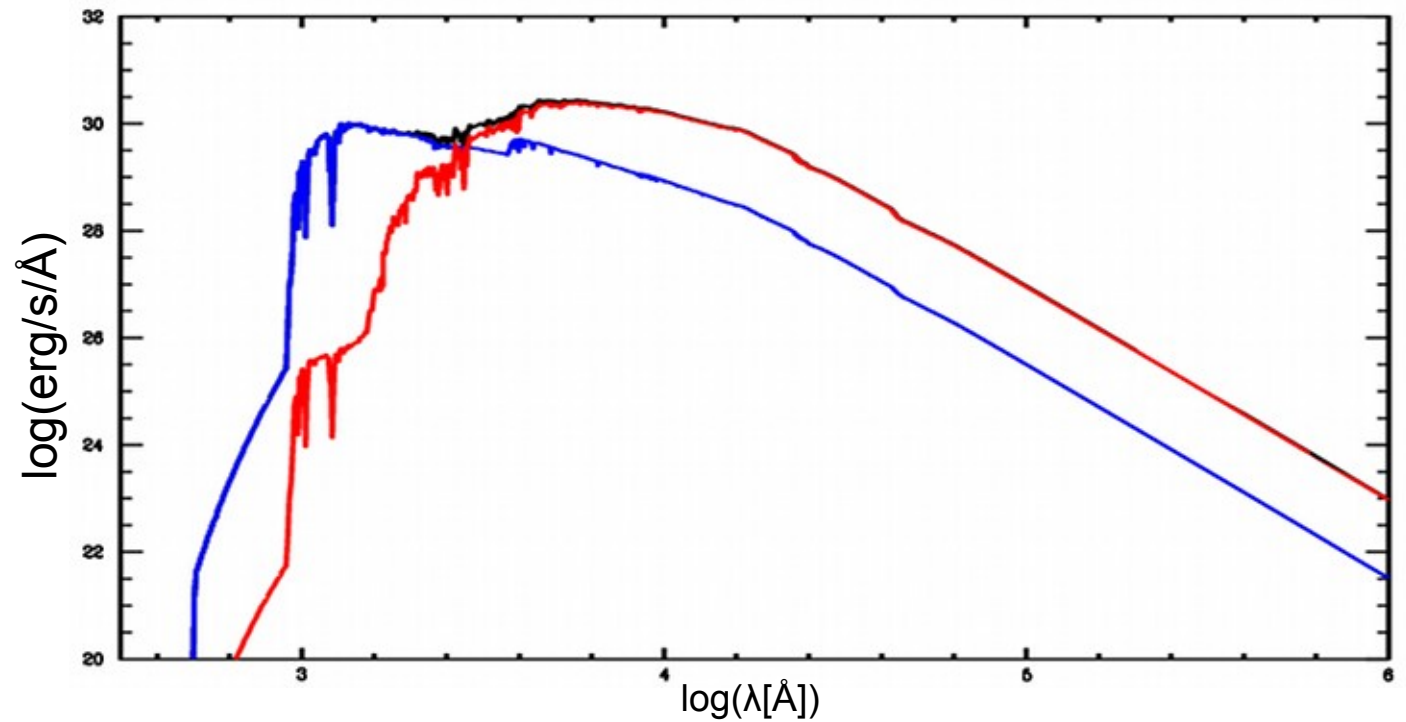
### 3. Stellar Population Synthesis models

- Bruzual & Charlot 2003
- Maraston 2005
- Conroy et al. 2009

# Spectral Energy Distribution (SED)



simulated →





# Stellar Population Synthesis models require Star Formation Histories

## Complications:

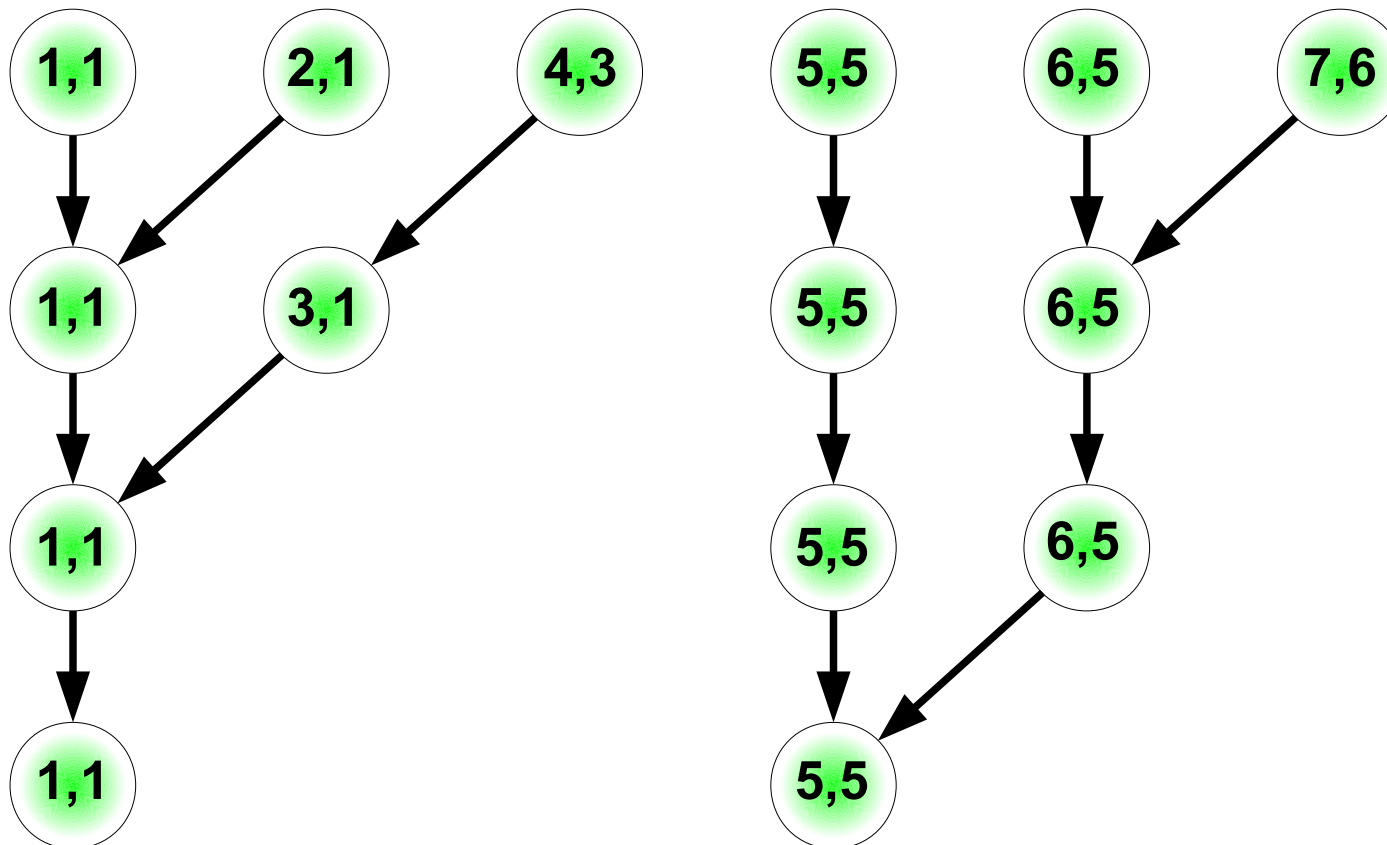
- Occupy large disk space
- Processing is time consuming
- Interpolation and re-binning

## One of the solutions:

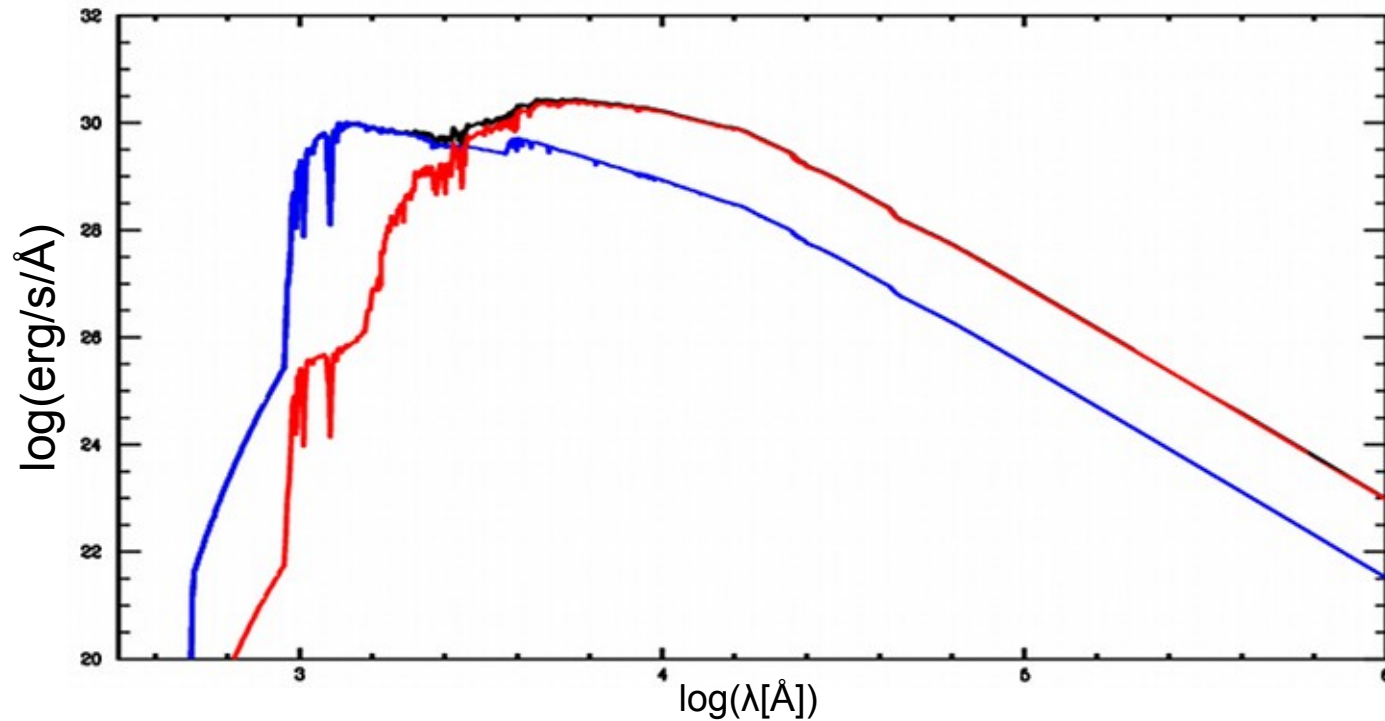
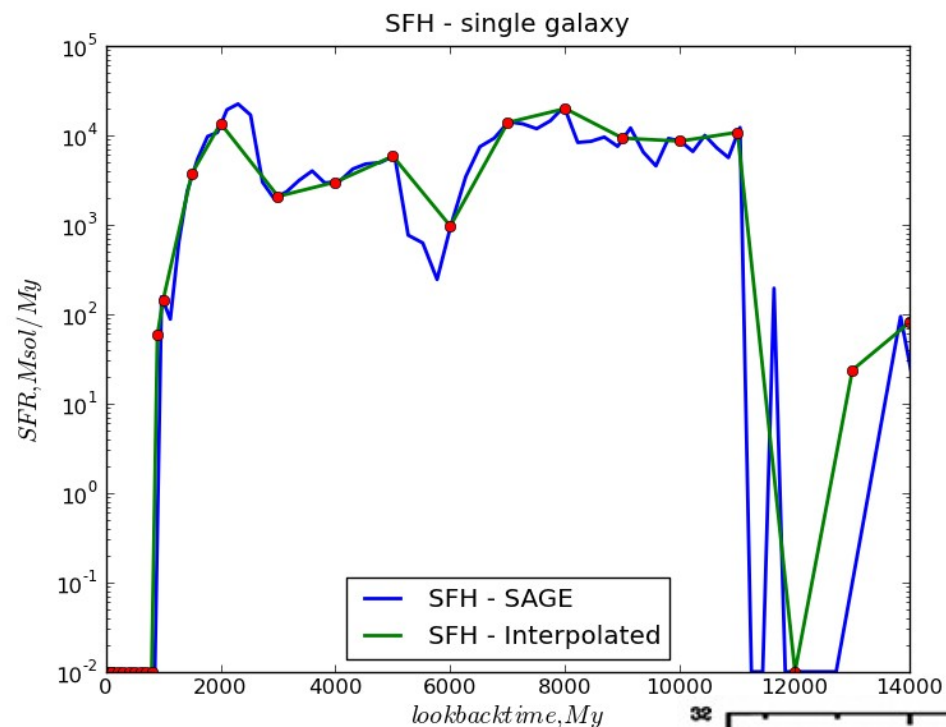
- Walking galaxies merger trees

# Galaxy Merger Trees

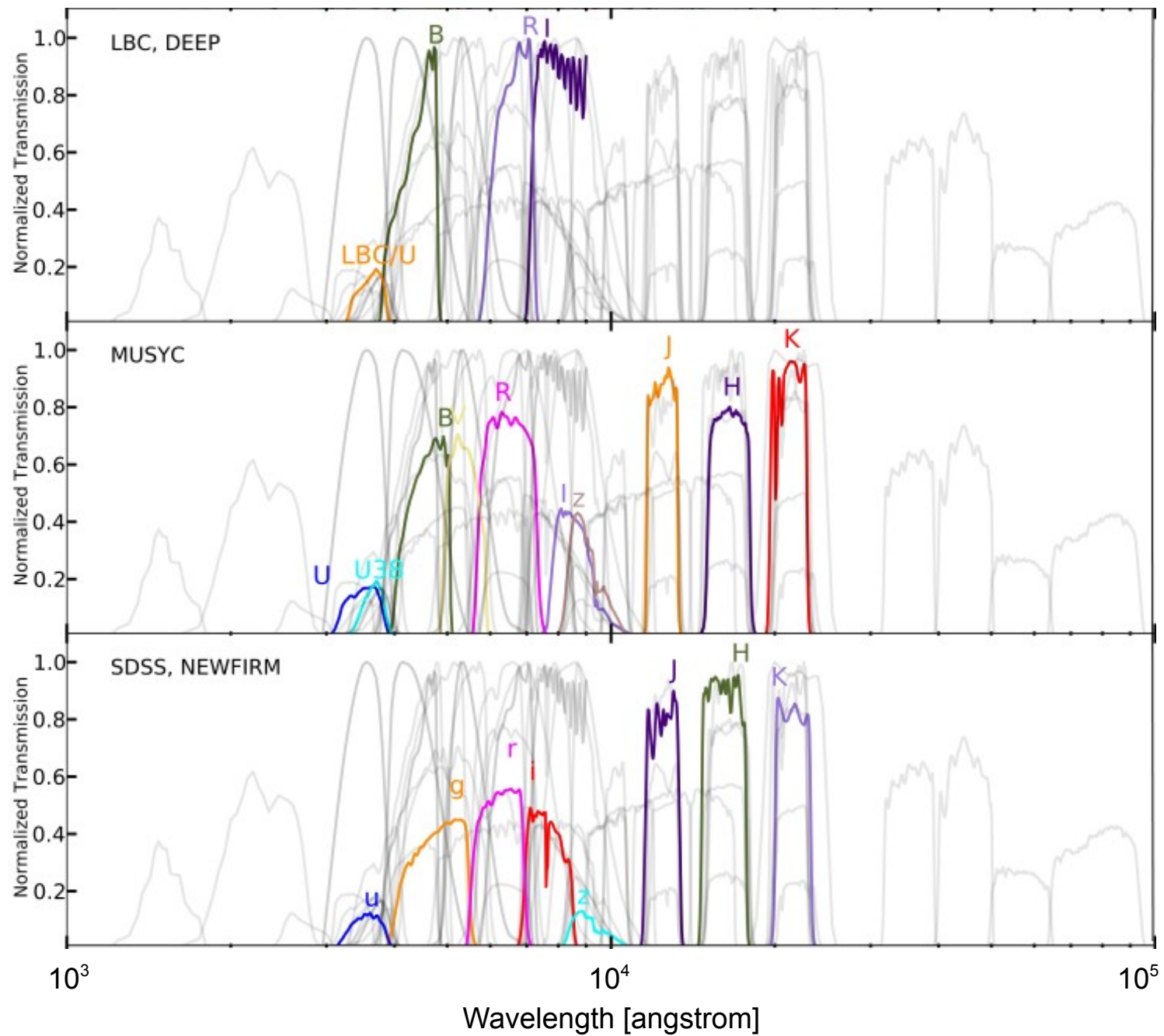
- Galaxy unique ID
- ID of a galaxy it will merge with



# Spectral Energy Distribution (SED)

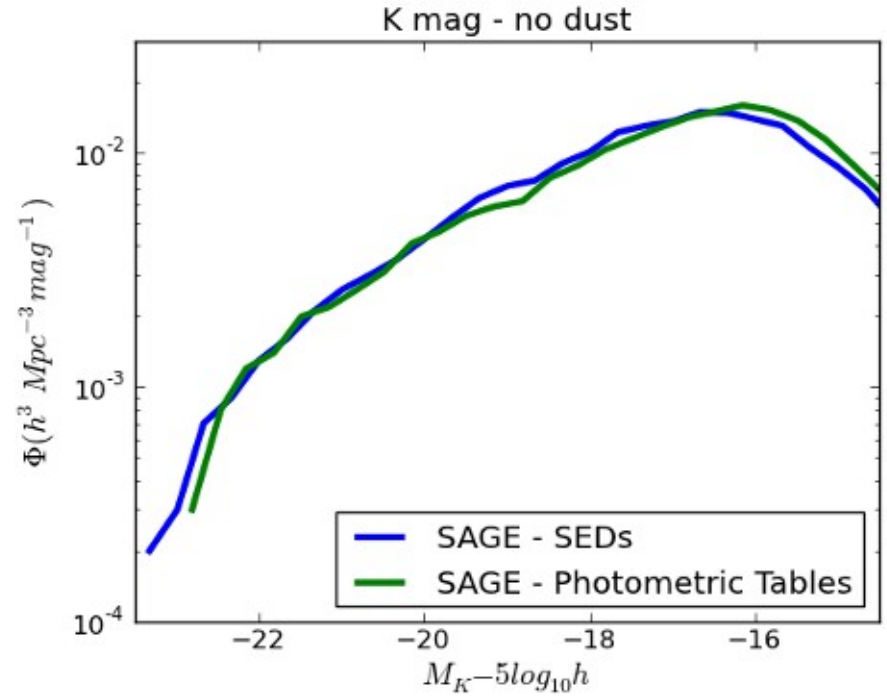
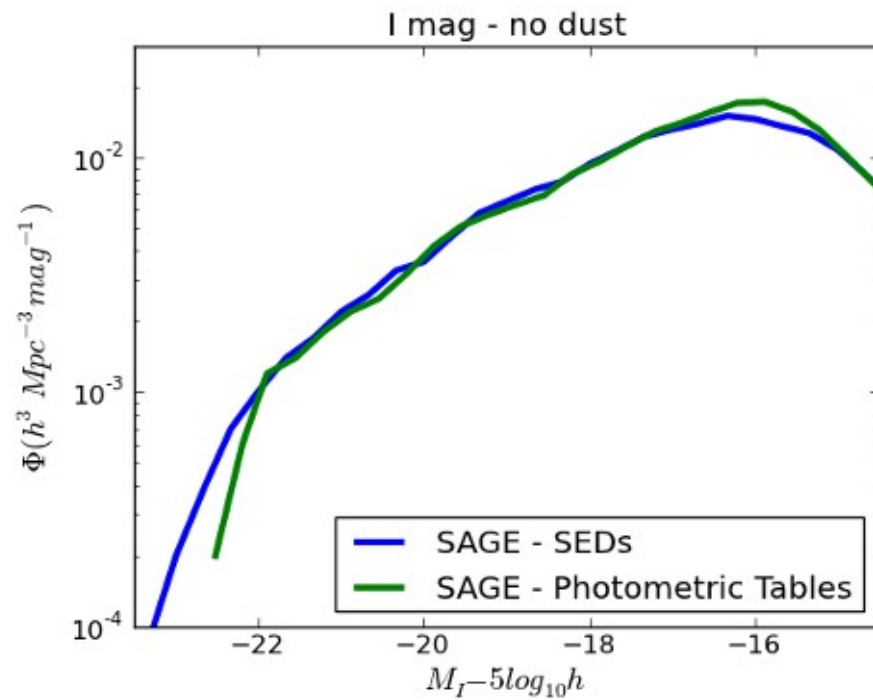
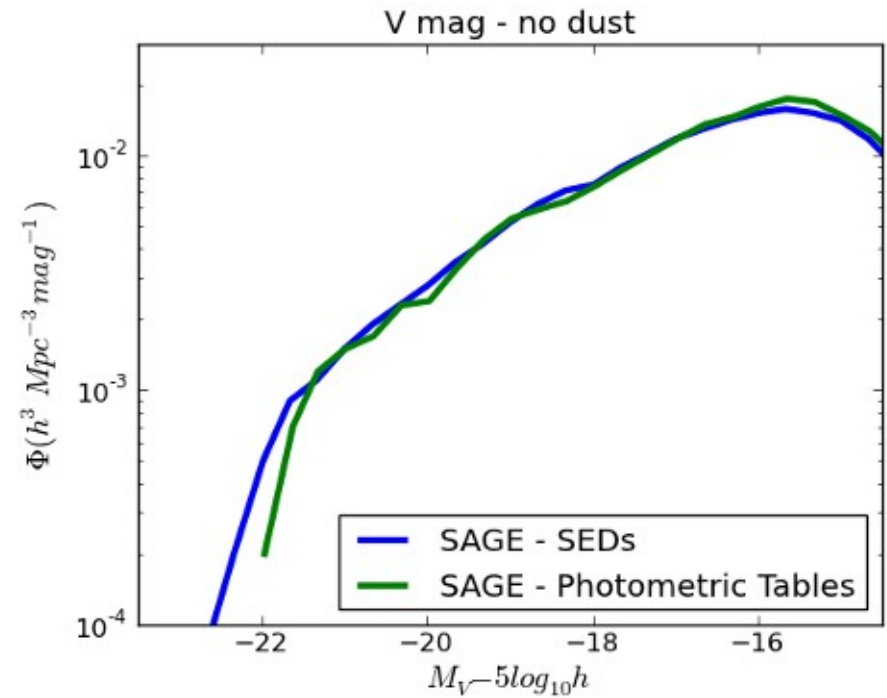
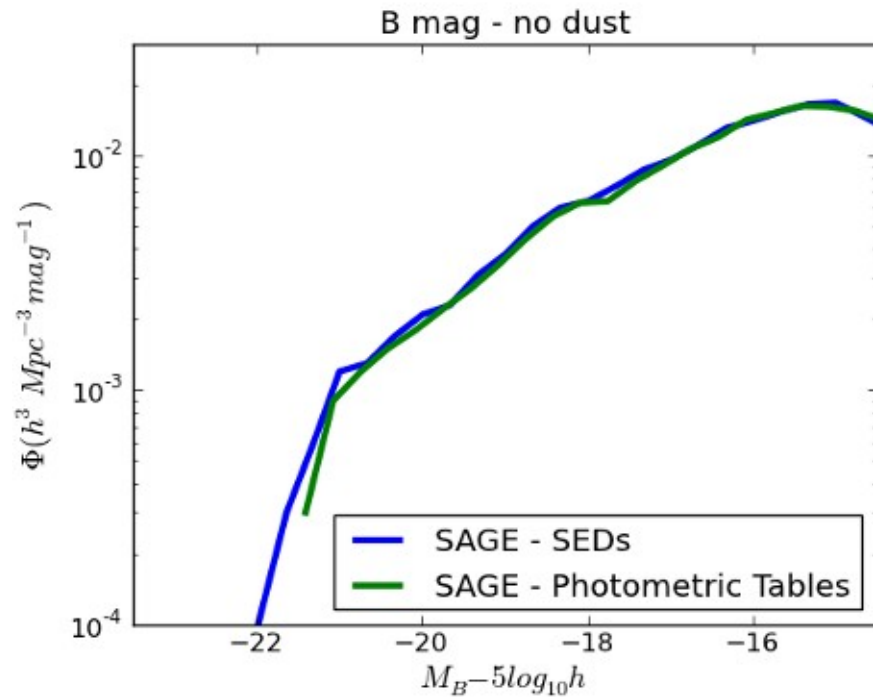


# Optical Filters

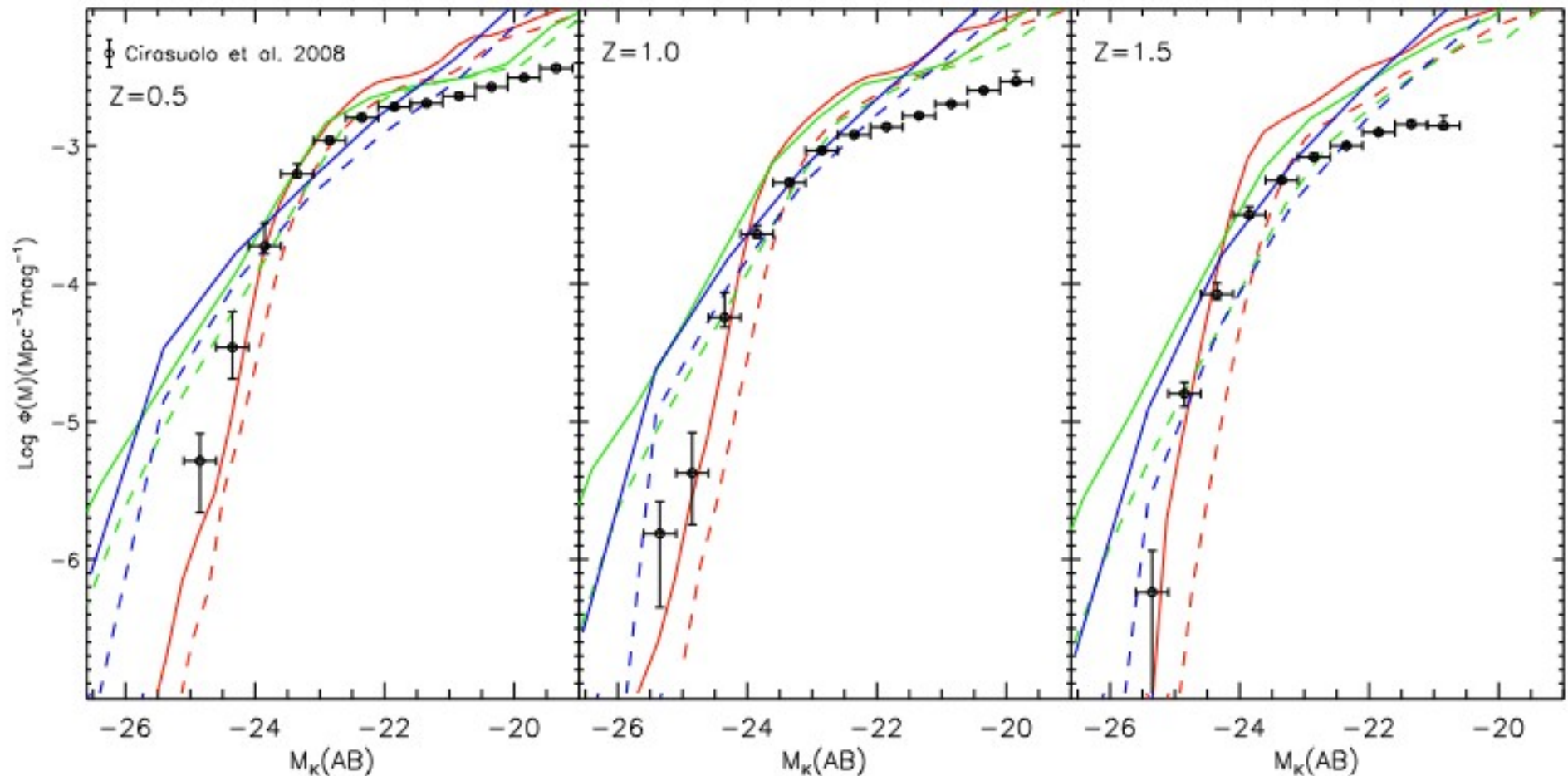




# Luminosity Function



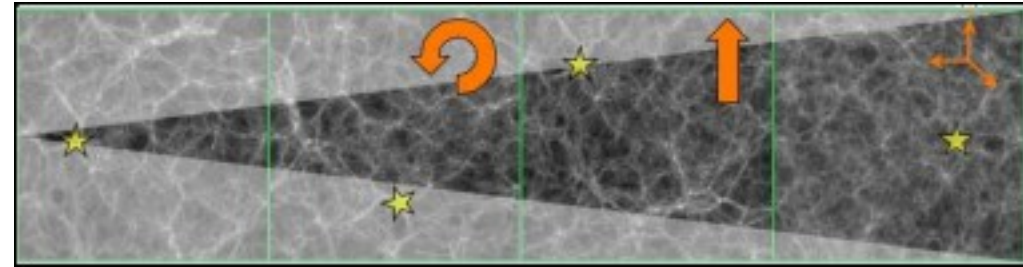
# K-Band Luminosity function with and without TP-AGB evolutionary phase



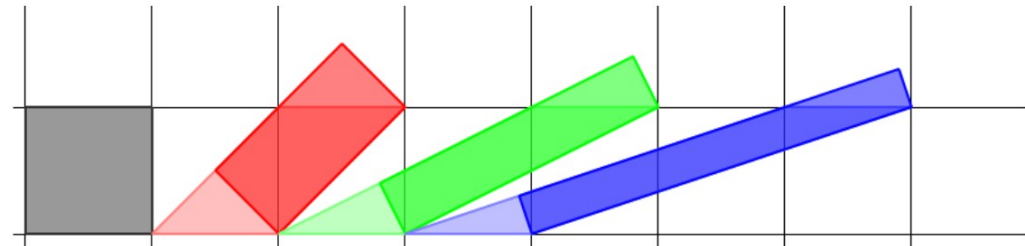
In red – semi-analytic model from De Lucia & Blaizot 2007, green – Monaco et al. 2007, and blue – Menci et al. 2006. Solid lines correspond to Maraston 2005 stellar population synthesis model, dashed – Bruzual & Charlot 2003. Data points from Cirasuolo et al. 2010 shown as black circles present data obtained from UKIDSS survey using UK Infrared Telescope in Hawaii.

## 4. Build a Lightcone

- Blaizot et al. 2004



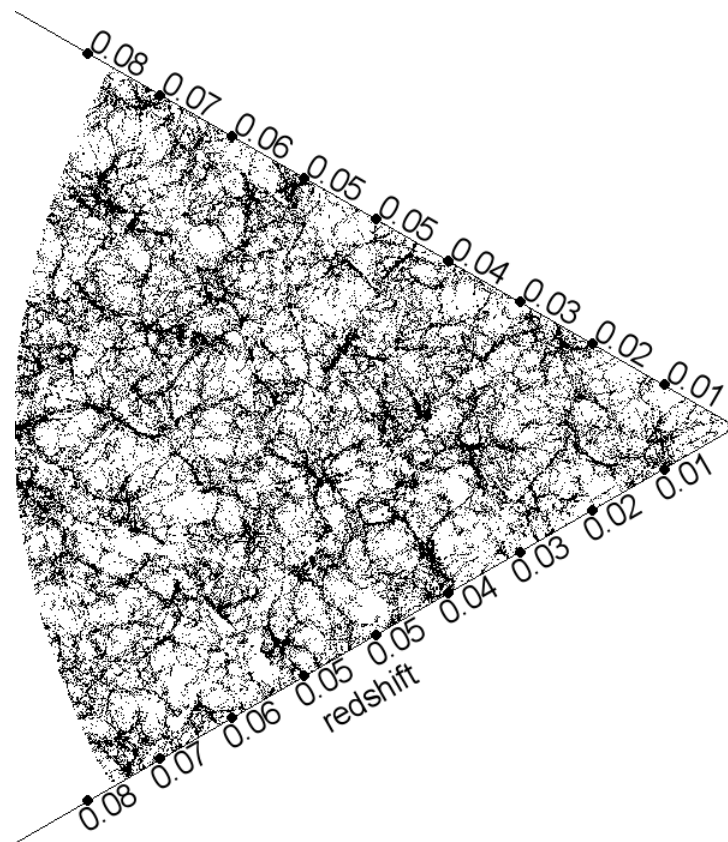
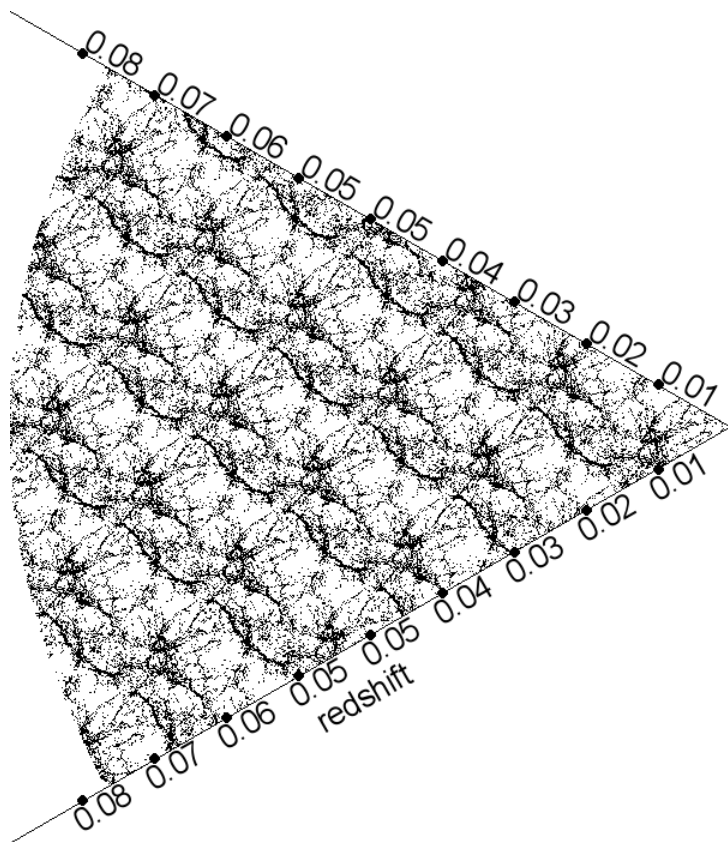
- Carlson & White 2010



- Bernyk et al. in prep.

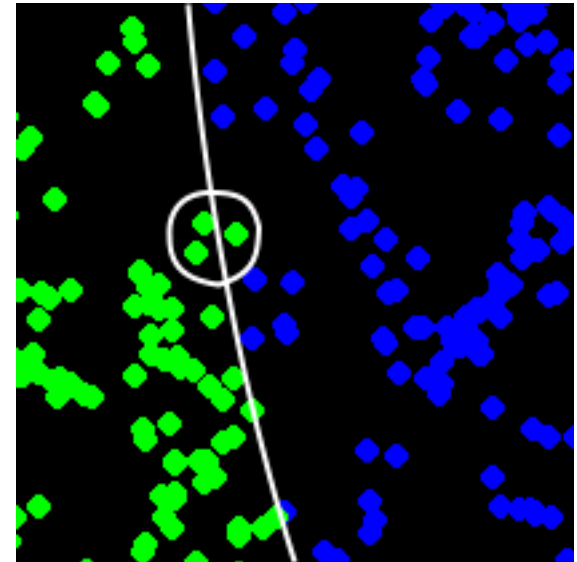
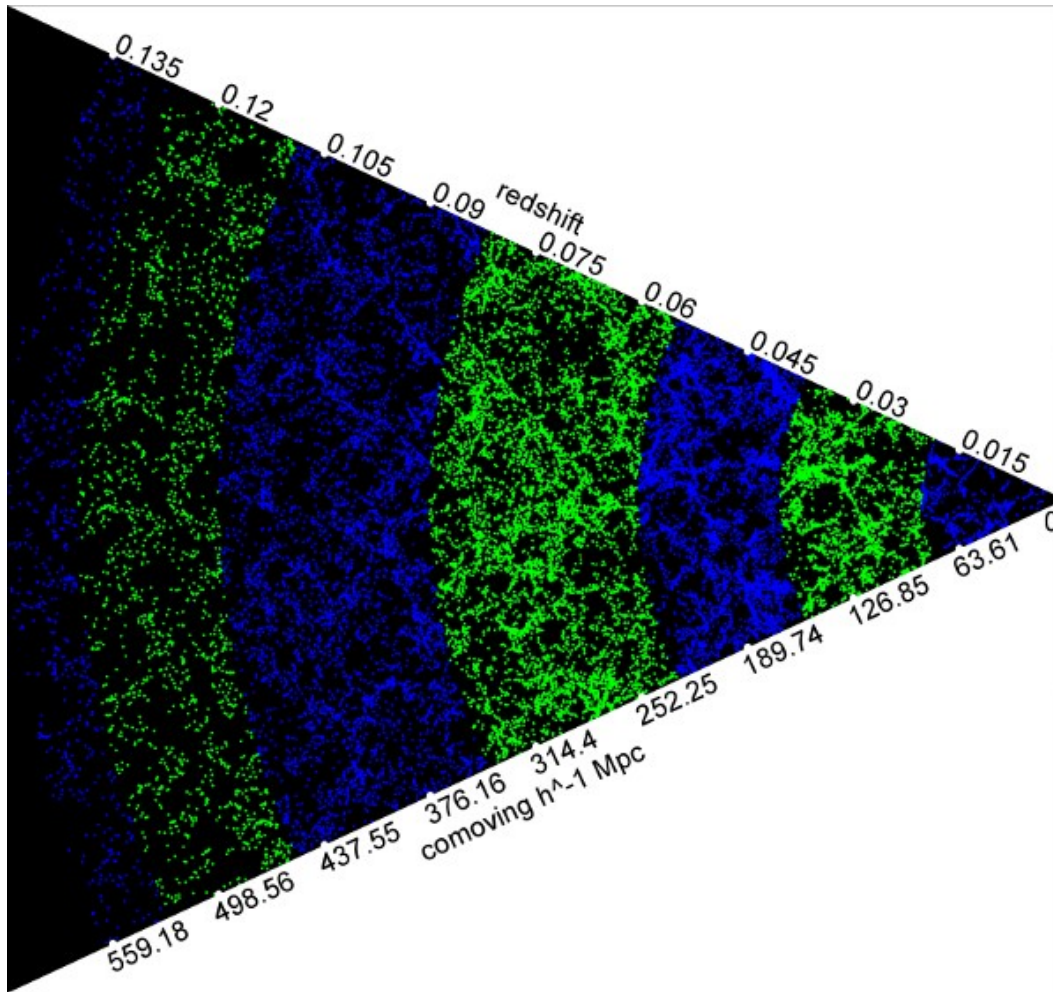
Using both techniques, and also extending Blaizot et al. 2004

# Random rotation and shifting

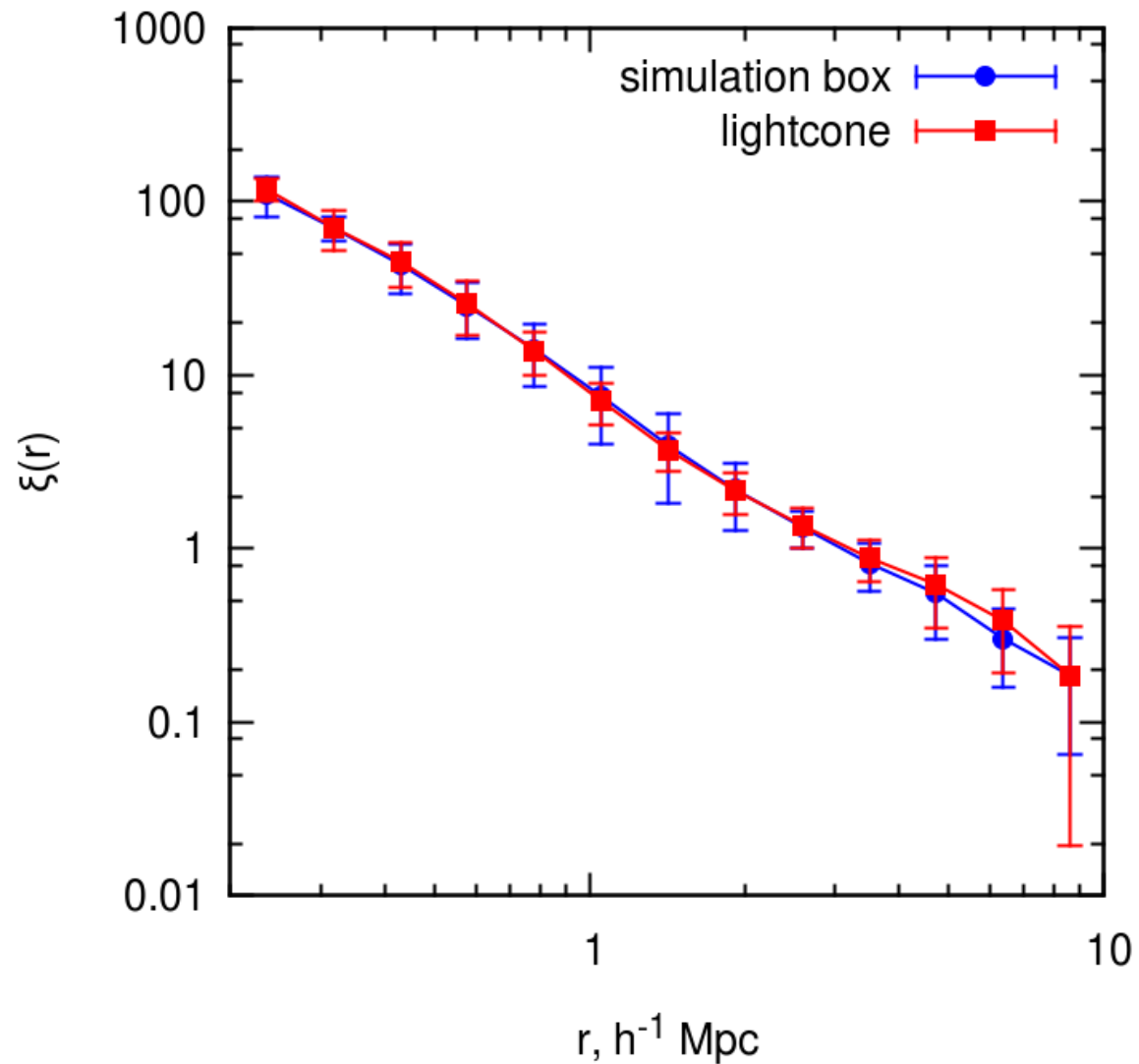




# Preserving Substructures



# Clustering Properties



# Mock Galaxy Factory

<http://tao.it.swin.edu.au>

- Cosmological Dark Matter simulations
- Galaxy models
- Catalogue shape: lightcone or cube
- Structure: unique or random
- Pregenerated mock catalogues

# Mock Galaxy Factory

## General -

Catalogue Type:	Light cone
Simulation:	Full Millennium
Galaxy Model:	Croton et al. 2006
Survey Preset:	Custom
Structure	Random - <input type="radio"/> Unique - <input checked="" type="radio"/>
Observer ID	0
Number of cones (Total unique cones = 7)	1

## Parameters -

Filter:	B (abs mag)	brightest:	
		faintest:	
$Z_{\min}$ :			0
$Z_{\max}$ :			.8
RA (arcmin, 0...5400):			81
Dec (arcmin, 0...5400):			81

## Selected simulation details:

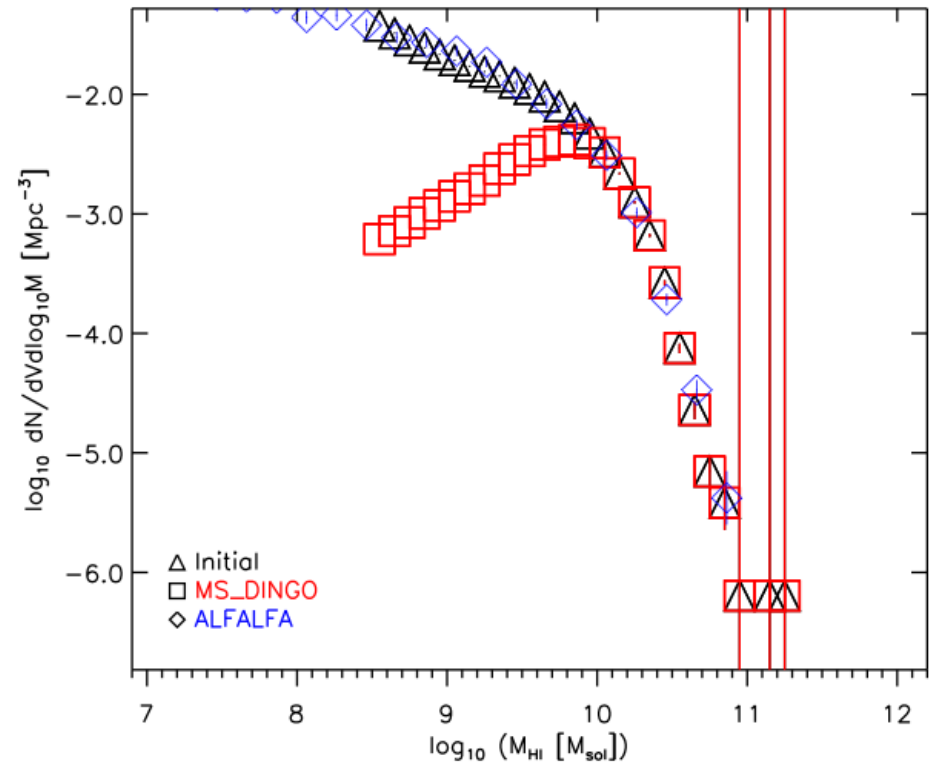
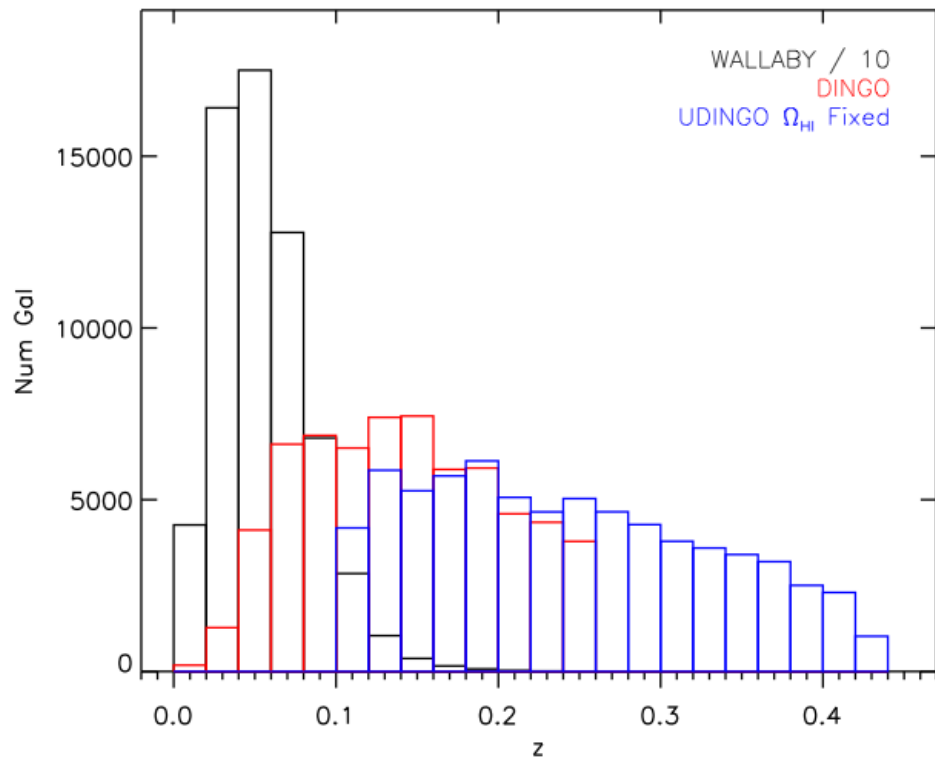
- **Full Millennium**  
Paper: [Springel et al. 2005](#)  
External link: <http://www.mpa-garching.mpg.de/galform/millennium/>  
Cosmology: WMAP-1  
Cosmological parameters:  $\Omega_m = 0.25$ ,  $\Omega_b = 0.045$ ,  $h = 0.73$ ,  $\Omega_\Lambda = 0.75$ ,  $n = 1$ ,  $\sigma_8 = 0.9$   
Box size: 500Mpc

## Selected galaxy model details:

- **Croton et al. 2006**  
Kind: semi-analytic model  
Paper: [Croton et al. 2006](#)

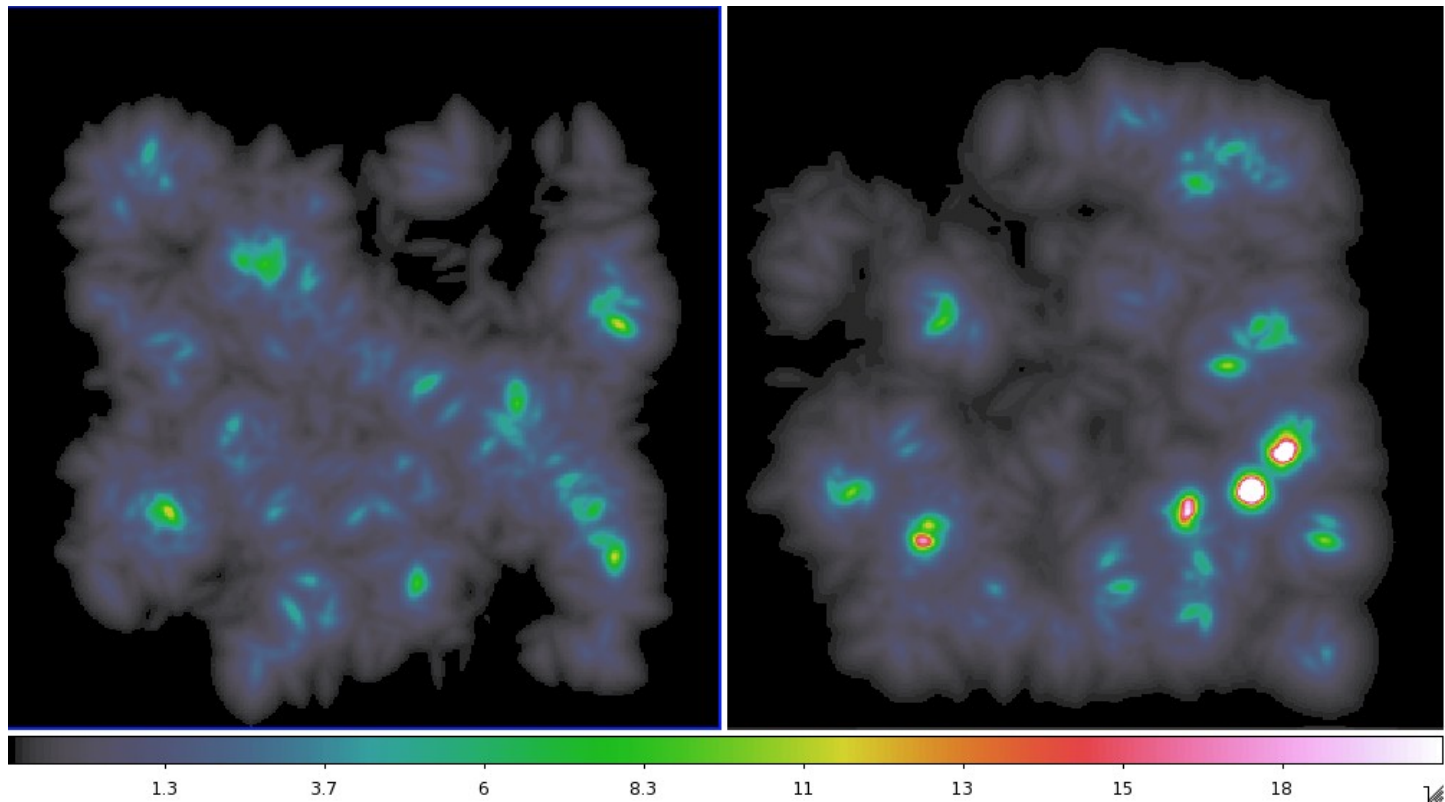


# Predictions for ASKAP Neutral Hydrogen Surveys



Duffy et al. - submitted

# First results FourStar: discovery of a candidate galaxy supercluster at $z \sim 2.2$



Spitler et al. in prep.

# Other Mock Surveys in progress

- ✓ PRIMUS
- ✓ GAMA
- ✓ CANDELS
- ✓ WiggleZ – coming early 2012



Image credit: NASA

# Types of AGN

Galaxy Type	Emission Lines		X-rays	Excess of		Strong Radio	Jets	Variable	Radio loud
	Narrow	Broad		UV	Far-IR				
<b>Seyfert I</b>	yes	yes	some	some	yes	few	no	yes	no
<b>Seyfert II</b>	yes	no	some	some	yes	few	yes	yes	no
<b>Quasar</b>	yes	yes	some	yes	yes	some	some	yes	10%
<b>Blazar</b>	no	some	yes	yes	no	yes	yes	yes	yes
<b>BL Lac</b>	no	none/faint	yes	yes	no	yes	yes	yes	yes
<b>OVV</b>	no	stronger than BL Lac	yes	yes	no	yes	yes	yes	yes
<b>Radio galaxy</b>	some	some	some	some	yes	yes	yes	yes	yes

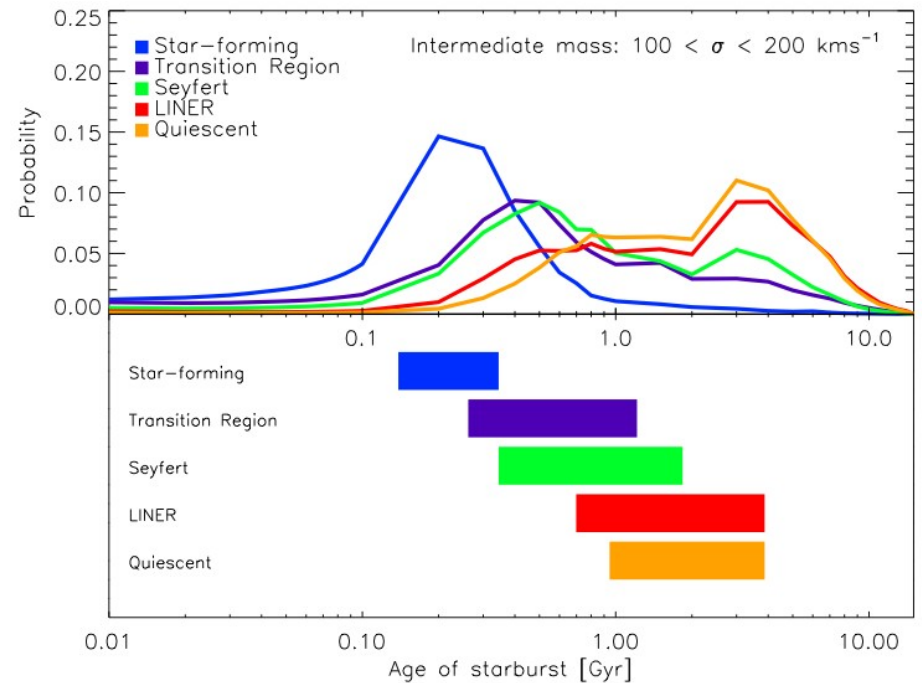
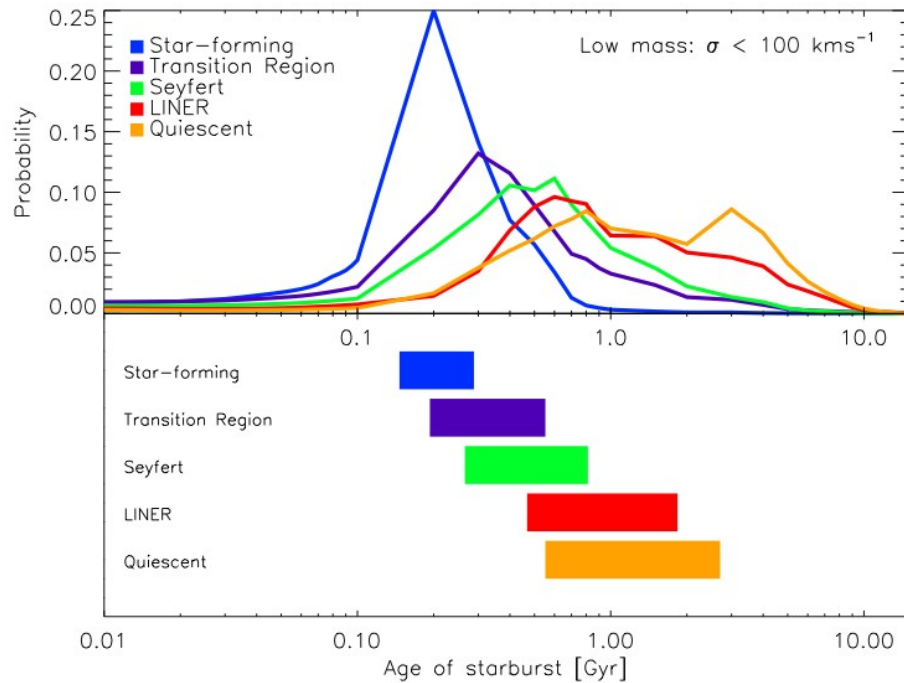


# AGN feedback

- Plays an important role in galaxy evolution
- No Galaxy Model which incorporates all types of AGN and accretion modes
- Observations are still at early stages

- Associate each major AGN class with a trigger and accretion model
- Couple this to the cosmological evolution of galaxies in the semi-analytic model

# Unified AGN model



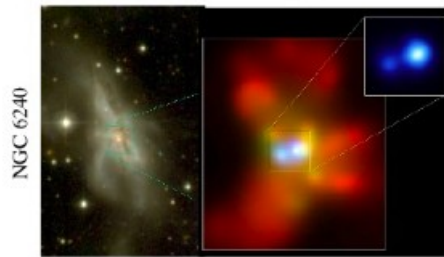
Schawinski et al. 2008

# Quasar light curve

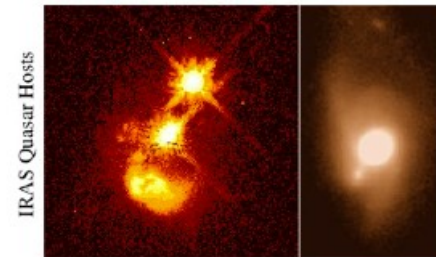
(c) Interaction/"Merger"



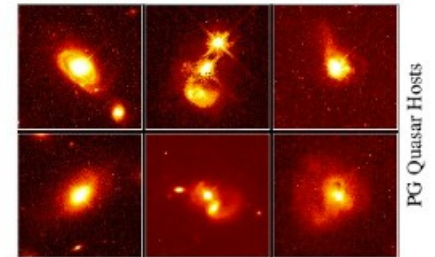
(d) Coalescence/(U)LIRG



(e) "Blowout"



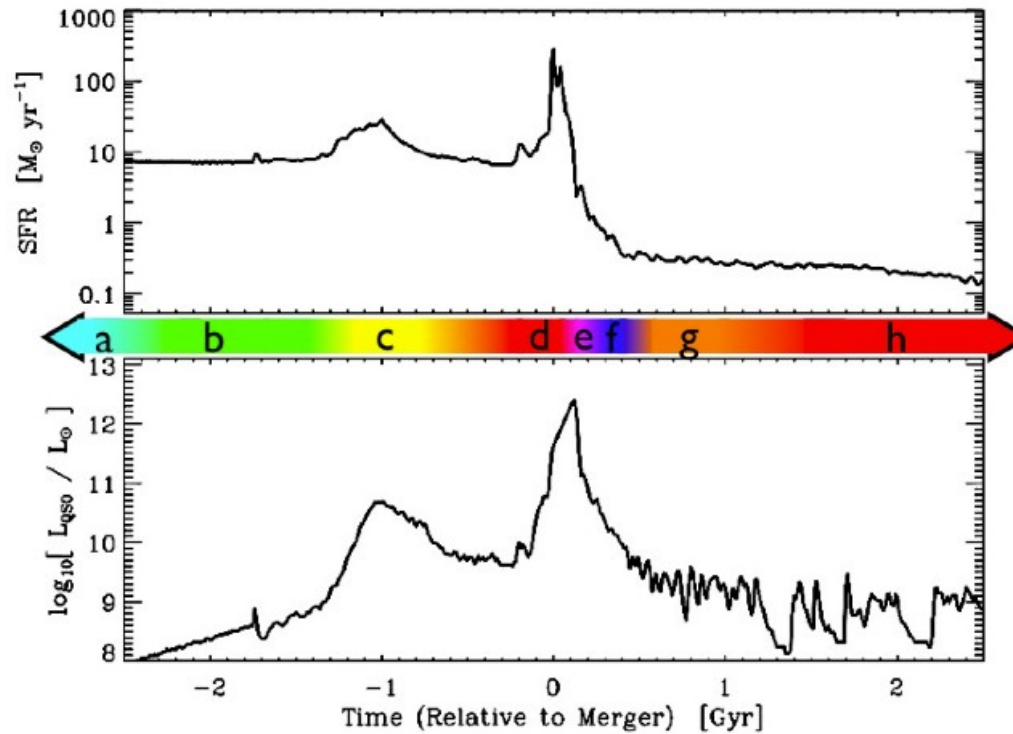
(f) Quasar



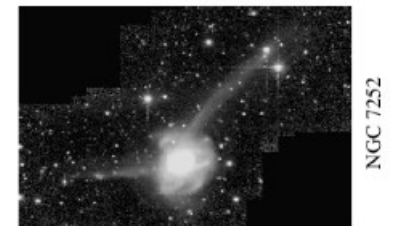
(b) "Small Group"



(a) Isolated Disk



(g) Decay/K+A

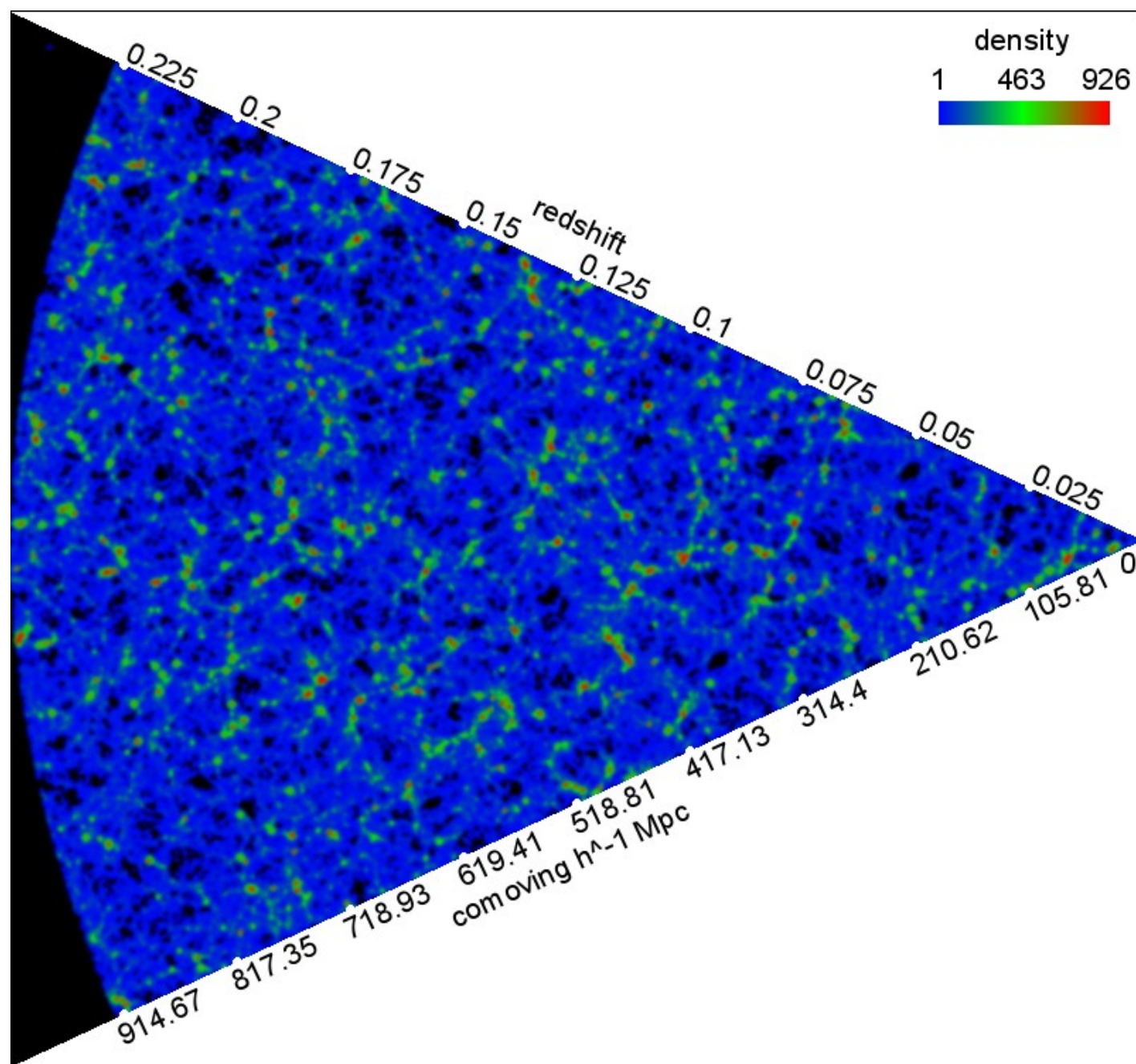


(h) "Dead" Elliptical



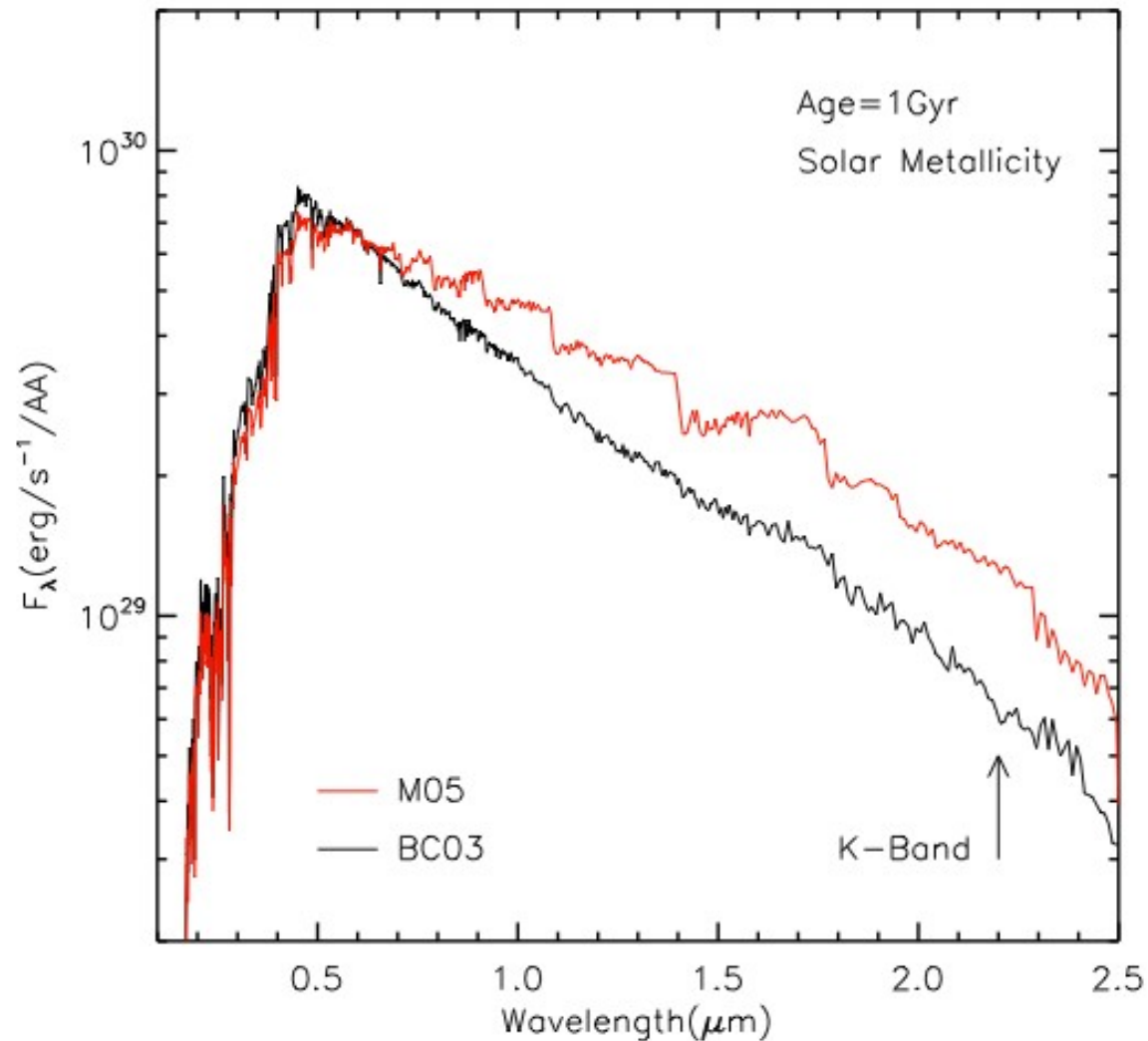
Hopkins et al. 2008

- The Mock Galaxy Factory is public
- SEDs are in progress
- More DM simulations and Galaxy Models are in progress
- Unified semi-analytic prescription for different AGN regimes is required for better understanding of the SMBH and galaxy co-evolution





# SSP SED with and without TP-AGB evolutionary phase



## Output -

### Core Properties:

- ☒ Redshift real
- ☐ Redshift observed
- ☒ Coordinates (RA, Dec)

### Halo Properties:

- ☐ Central Halo Mvir
- ☐ Halo Mvir
- ☐ Halo Vvir
- ☐ Halo Velocity dispersion

### Intrinsic Galaxy Properties:

- ☐ Galaxy Type
- ☐ Stellar Mass
- ☐ Bulge mass
- ☐ DiskRadius

### Galaxy Luminosities:

- ☐ B (abs mag)
- ☐ V (abs mag)

[select](#) / [deselect](#) all

## Miscellaneous -

Output format:

HDF5

Email results to:

[Submit](#)