

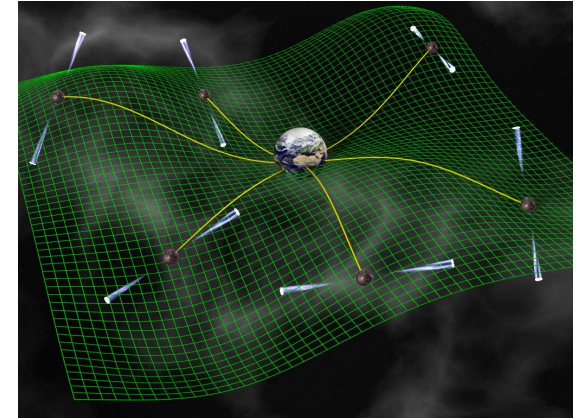
# *The Parkes Pulsar Timing Array*

*Co-learnium: 3 August 2023*

**Dick Manchester**

# The Parkes Pulsar Timing Array Project

- Concept of a PTA for nanoHertz GW detection first proposed by Romani (1989) and Foster & Backer (1990)
- PPTA project commenced in July 2003 with support from RNM's Australian Research Council Federation Fellowship and from CSIRO
- Two post-docs (George and Russell Edwards) employed, construction of 10cm/50cm receiver began
- Initial collaboration between ATNF and Swinburne University (Matthew Bailes' group); later collaborations formed with other groups, both Australian and international
- Scheduled observations at Parkes (aka "Murriyang") commenced in February 2004

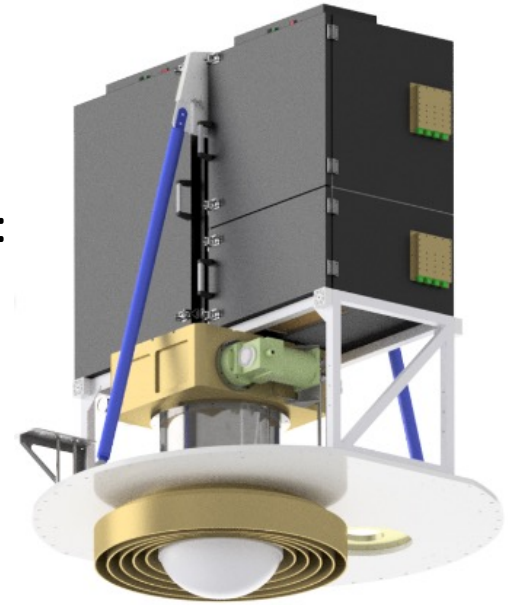


Credit: David Champion



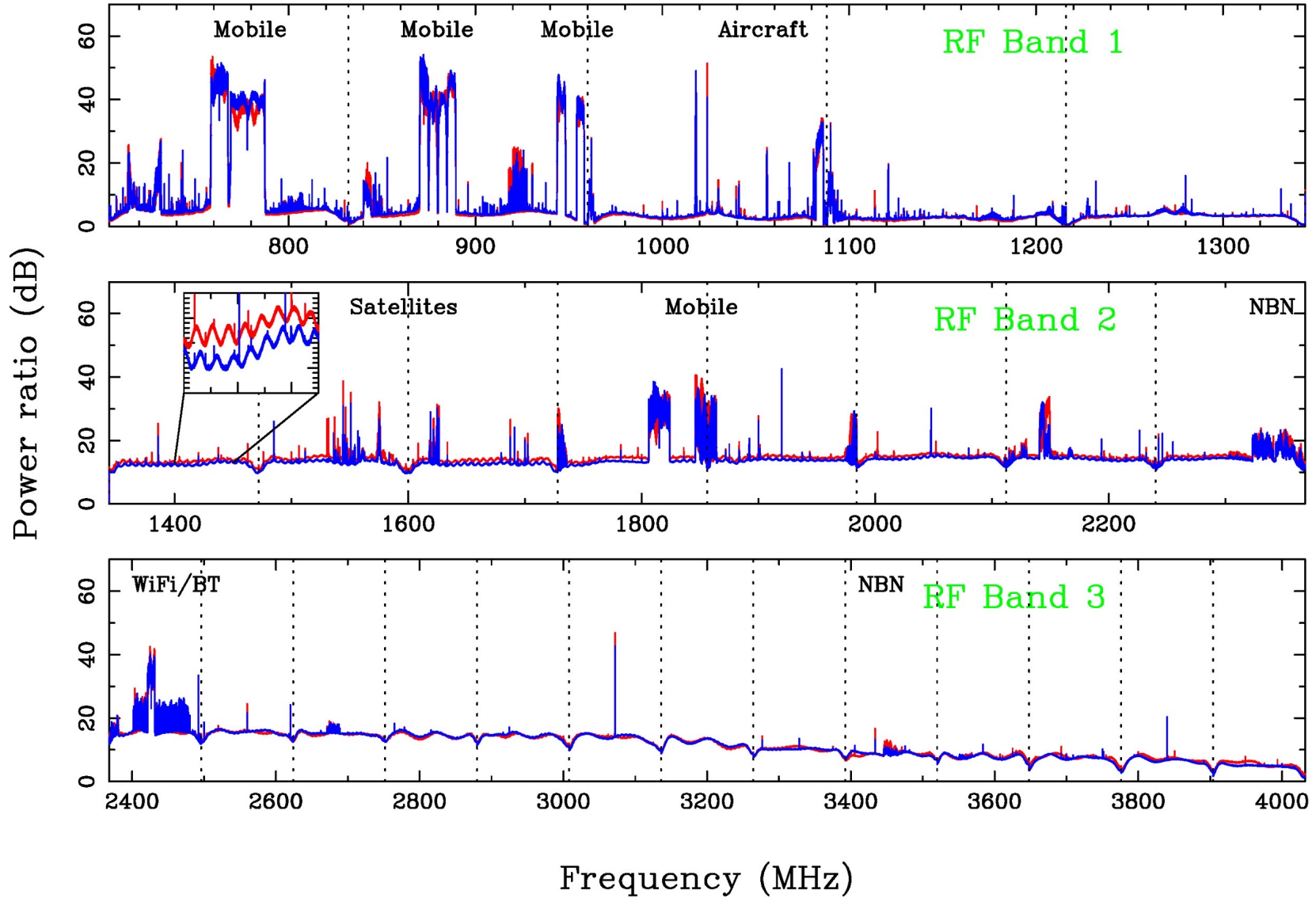
# PPTA Observations

- All observations use Parkes 64-m radio telescope
- Until 2018, all observations in three bands, two receivers:
  - 50/40cm band (700/64 MHz) 10cm/50cm receiver
  - 20cm band (1370/256 MHz) Centre beam of MB receiver
  - 10cm band (3100/1024 MHz) 10cm/50cm receiver
- Ultra-wide Low (UWL) receiver commissioned in 2018 (704 MHz – 4032 MHz in 26 x 128-MHz sub-bands). Now in regular use for PPTA observations
- 37 MSPs in PPTA sample; 20 - 25 regularly observed, 1hr for each pulsar
- On average, two observing sessions per month, some short (e.g., 8 hrs), some long (e.g., 2 days)
- ~15,500 hrs of observation time since 2003



The UWL Receiver

# UWL Observed Spectrum



# Sky Distribution of PPTA Pulsars

- PPTA pulsars are widely distributed in Galactic longitude (over the range visible to Parkes) and mostly at mid- to high-latitudes
- In Celestial coordinates, most of the PPTA pulsars are in Galactic time (16 – 20 hrs) which is in high demand
- This leads to very uneven observational coverage of the PPTA sample and ultimately to reduced sensitivity of the PPTA data set

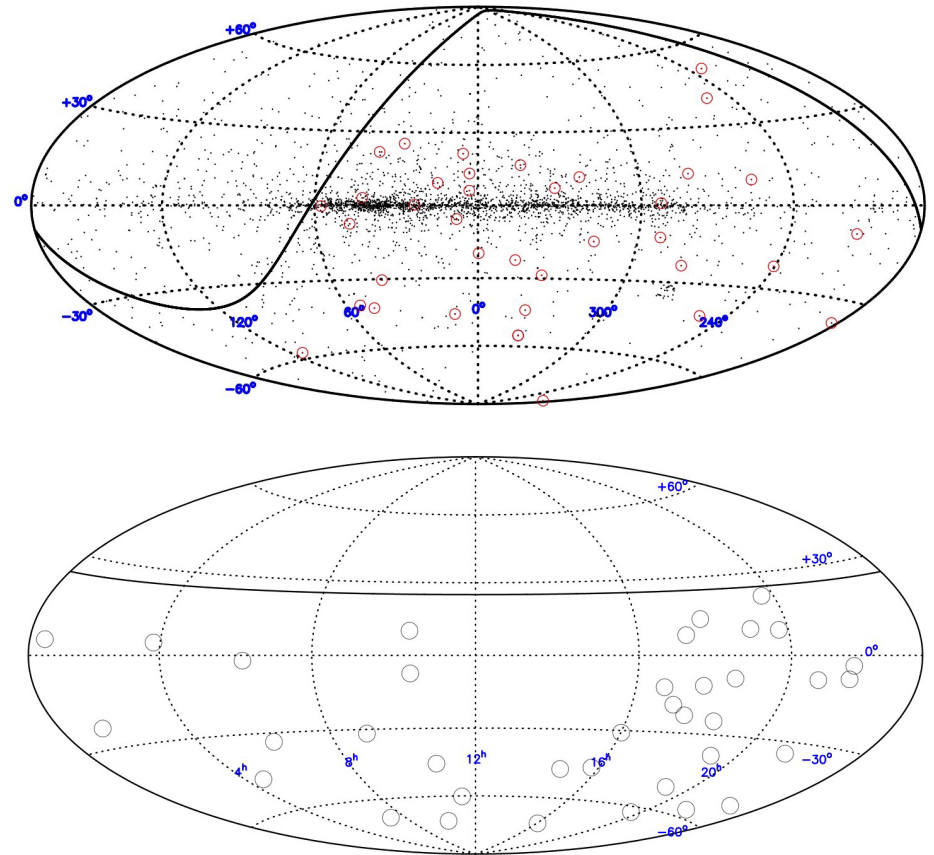


Image credits: George Hobbs

# PPTA Membership (2023)

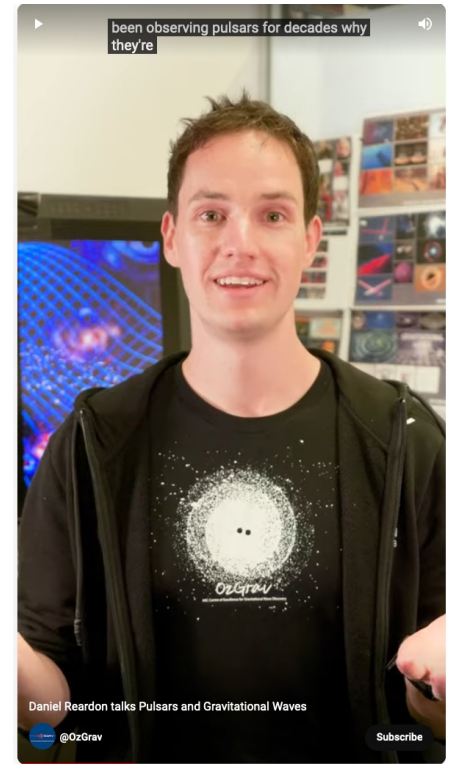
- Nine Australian institutions, 26 members (incl. 3 students)
- Ten international institutions, 10 members
- Members are co-authors on papers if they have made a significant contribution to the particular paper
- PPTA managed by a Steering Committee, currently consisting of eight people, including a student representative and a post-doc representative
- Current Chair: George Hobbs
- Two PPTA representatives are on the Steering Committee for the International Pulsar Timing Array (IPTA), currently Andrew Zic and Daniel Reardon
- The IPTA has various Working Groups and the PPTA is represented on most of them.



# Latest GWB Results

- Latest PPTA result on isotropic GWB published on July 1 as part of coordinated IPTA data release  
Reardon, Zic, Shannon et al. ApJL, July 1 2023
- Big PR event 29 June, led by Daniel and Andrew for the PPTA
- NANOGrav 15-yr data set, 68 pulsars > 3yr, gives  
~3.5 $\sigma$  result:  $h_c(f) = 2.4 \pm 0.7 \times 10^{-15}$ ; Agazie et al. ApJL, July 1 2023
- EPTA DR2 + InPTA – 24 yrs, 42 pulsars, gives  
 $h_c(f) = 2.5 \pm 0.7 \times 10^{-15}$ ; ~3 $\sigma$  (tentative) result – issues with combining different data sets  
Antoniadis et al. arXiv, June 28 2023
- CPTA DR1 – 3.5 yrs, 57 pulsars using FAST, gives  
 $h_c(f) = 1.0 \pm 0.3 \times 10^{-14}$ ; ~4.6 $\sigma$  (tentative) result – issues with signal-processing method  
Xu et al. RAA, July 1, 2023

***The most important factor giving high significance is the number of pulsars regularly observed***



OzGrav



# PPTA Isotropic GWB Paper

THE ASTROPHYSICAL JOURNAL LETTERS, 951:L6 (15pp), 2023 July 1






























<https://doi.org/10.3847/2041-8213/acdd02>

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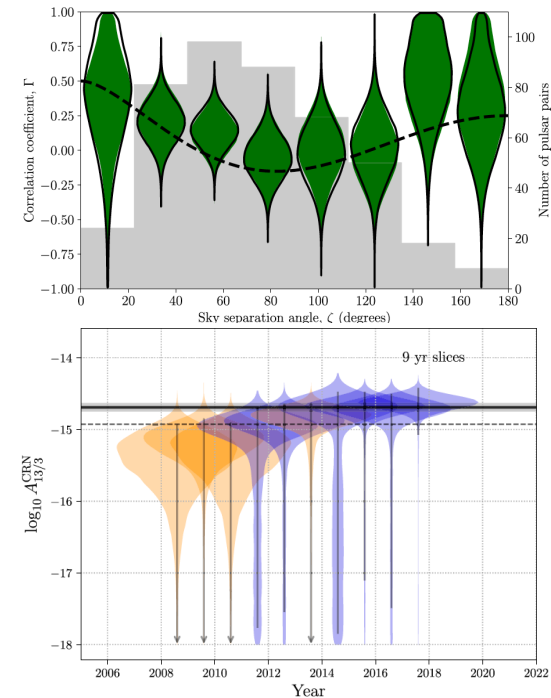
OPEN ACCESS



## Search for an Isotropic Gravitational-wave Background with the Parkes Pulsar Timing Array

Daniel J. Reardon<sup>1,2</sup> , Andrew Zic<sup>3,4</sup> , Ryan M. Shannon<sup>1,2</sup> , George B. Hobbs<sup>3</sup> , Matthew Bailes<sup>1,2</sup> ,  
Valentina Di Marco<sup>5,6</sup> , Agastya Kapur<sup>3,4</sup> , Axl F. Rogers<sup>7</sup> , Eric Thrane<sup>5,6</sup> , Jacob Askew<sup>1,2</sup> , N. D. Ramesh Bhat<sup>8</sup> ,  
Andrew Cameron<sup>1,2</sup> , Małgorzata Curyło<sup>9</sup> , William A. Coles<sup>10</sup> , Shi Dai<sup>11</sup> , Boris Goncharov<sup>12,13</sup> , Matthew Kerr<sup>14</sup> ,  
Atharva Kulkarni<sup>1,2</sup> , Yuri Levin<sup>5,15,16</sup> , Marcus E. Lower<sup>3</sup> , Richard N. Manchester<sup>3</sup> , Rami Mandow<sup>3,4</sup> ,  
Matthew T. Miles<sup>1,2</sup> , Rowina S. Nathan<sup>5,6</sup> , Stefan Osłowski<sup>17</sup> , Christopher J. Russell<sup>18</sup> , Renée Spiewak<sup>19</sup> ,  
Songbo Zhang<sup>3,20</sup> , and Xing-Jiang Zhu<sup>21</sup> 

- Part of IPTA “3P+” coordinated release of GWB search and data results from major PTAs
- Tentative detection of GWB at  $2\sigma$  level,  
$$h_c(f) = 2.04 \pm 0.25 \times 10^{-15} \text{ for } \alpha = -2/3$$
- Evidence for time variation in signal – weaker at earlier times; consistent with earlier PPTA limit of  $h_c(f) < 1 \times 10^{-15}$  Shannon et al., Science, 2015





# PPTA Ultra-Light Dark Matter Paper

PHYSICAL REVIEW D **106**, L081101 (2022)

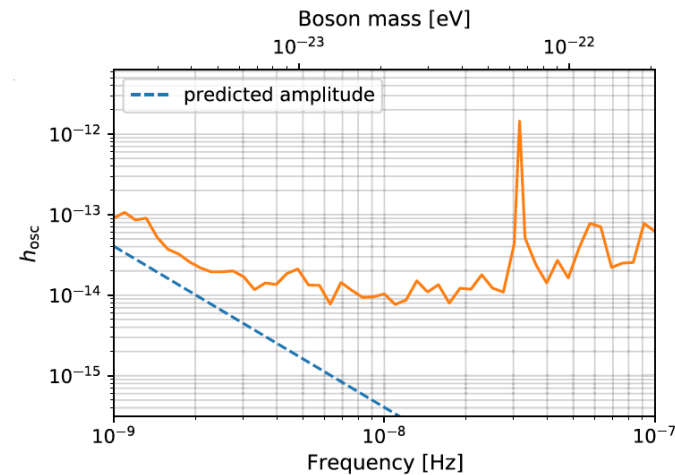
Letter

## Constraining ultralight vector dark matter with the Parkes Pulsar Timing Array second data release

Yu-Mei Wu,<sup>1,2,3,\*</sup> Zu-Cheng Chen<sup>4,5,†</sup> Qing-Guo Huang,<sup>1,3,2,‡</sup> Xingjiang Zhu,<sup>5,§</sup> N. D. Ramesh Bhat,<sup>6</sup> Yi Feng,<sup>7</sup>  
George Hobbs,<sup>8</sup> Richard N. Manchester,<sup>8</sup> Christopher J. Russell,<sup>9</sup> and R. M. Shannon<sup>10,11</sup>

(PPTA Collaboration)

- PTAs can constrain local density of ultra-light bosons ( $m \sim 10^{-23} \text{ eV} \sim 10^{-56} \text{ g}$ ) (!)
- PPTA DR2 used to set limit  $h_c(f) = \sim 10^{-14}$
- At low frequencies ( $1 - 5 \times 10^{-9} \text{ Hz}$ ), PPTA limit is  $\sim$  five times expected value
- Limit is similar to that set by PPTA for scalar dark matter (Poraykov et al., Phys Rev D, 2018)



# PPTA Outreach

- Led by Rob Hollow (ATNF)
- Rob's tireless promotion of the ATNF and, in particular, Parkes and its pulsar research has brought awareness of our activities in these areas to many groups both in Australia and internationally
- An important component of ATNF's Outreach is the PULSE@Parkes project
- Initiated by George and Rob about 17 years ago
- Groups of 20-30 high-school students (Yr 10-12) observe pulsars using Parkes, 2hr sessions
- Some international groups, China, Japan, Thailand
- Analyse Parkes data, measure DMs; data freely available on ATNF archive
- Over 3500 students from more than 300 schools have participated
- About 100 papers published in refereed journals



PULSE@Parkes observing session

# Summary

## In summary:

- Over its 20-year lifetime, the PPTA has had ~15,000 hrs of observation time on the Parkes Radio Telescope (“Murriyang”)
- Using these data, the project has produced a large number (many hundreds) of important papers, mostly in refereed journals
- It has fostered an extensive Outreach program that has reached thousands of school students and the general public both in Australia and internationally

***Overall, an extremely successful project!***