

MILLIMETRE METHANOL MASERS

Dinah Cragg

School of Chemistry
Monash University

- Peter Godfrey (Monash U.)
- Andrej Sobolev (Ural State U.)

- Ed Sutton (U. Illinois)
- Simon Ellingsen (U.Tasmania)

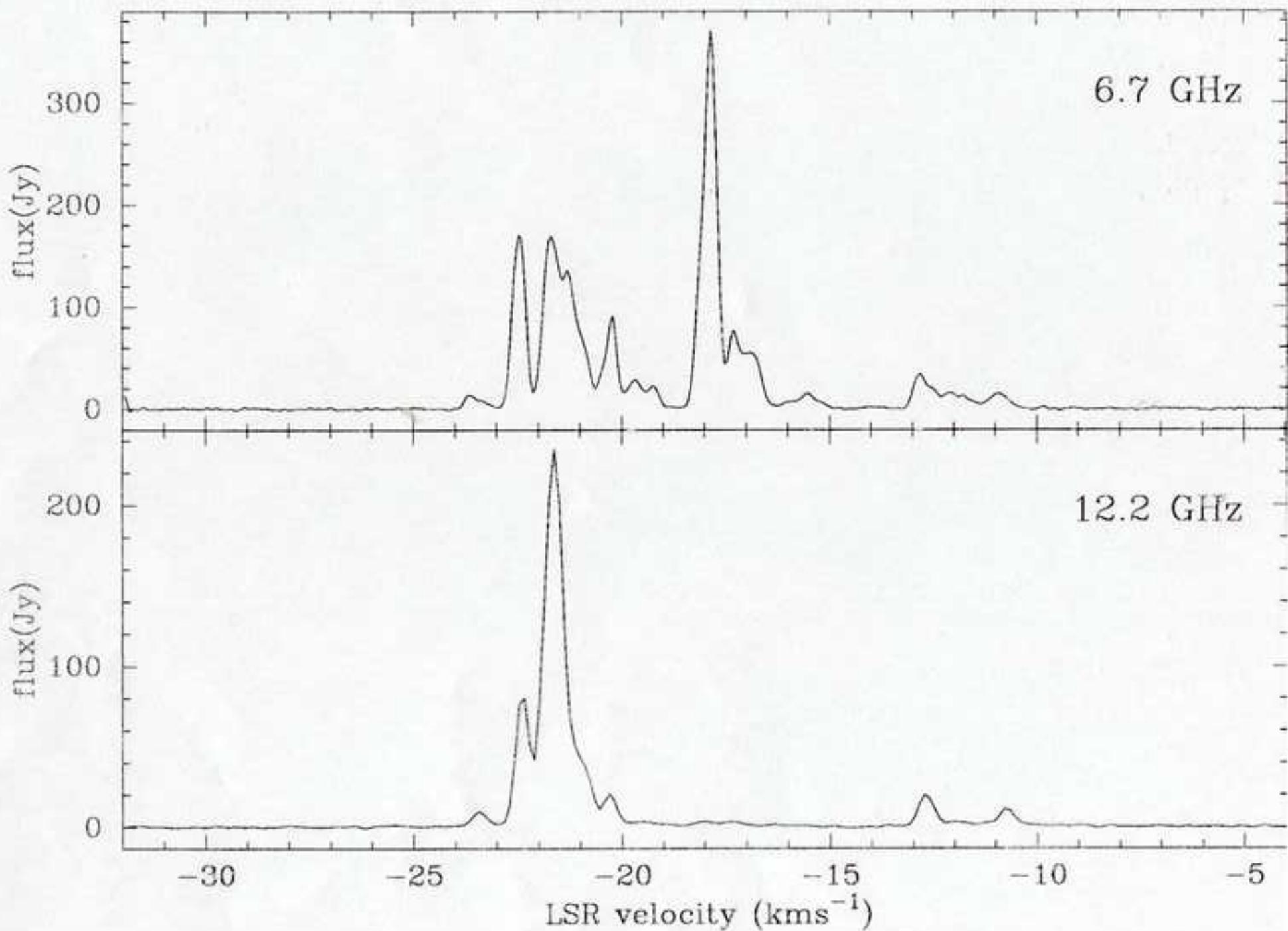
Class II Methanol masers

- Associated with recent massive star-formation
- ~ 500 Galactic sources at 6.6 and 12.1 GHz
- Spot size milliarcsec, clusters arcsec
- Brightness up to 10^{12} K
- 6.6 and 12.1 GHz masers coincide spatially
- ~ 30 sources also have weaker masers in mm range

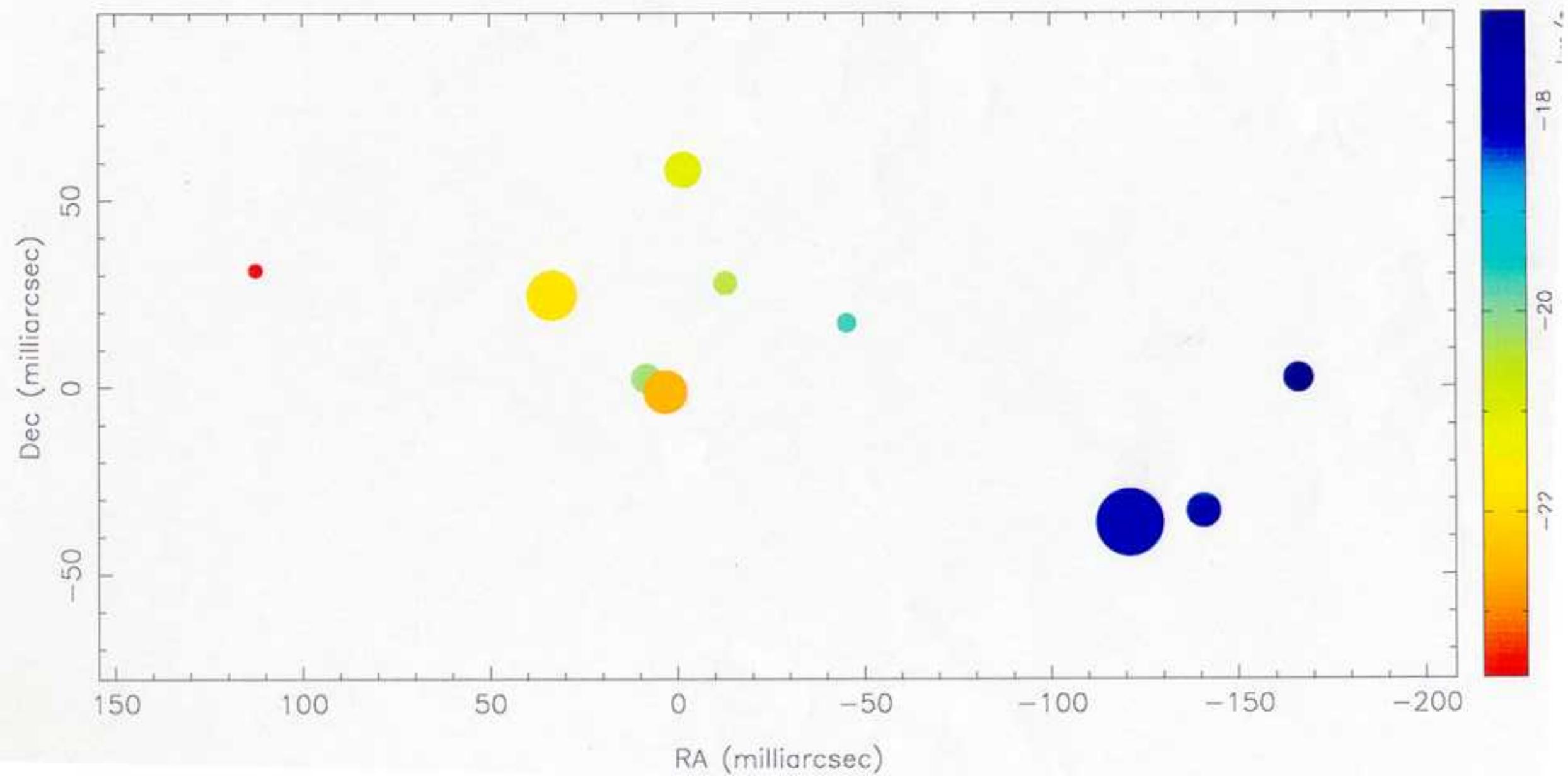
What is the spatial distribution of the mm masers?

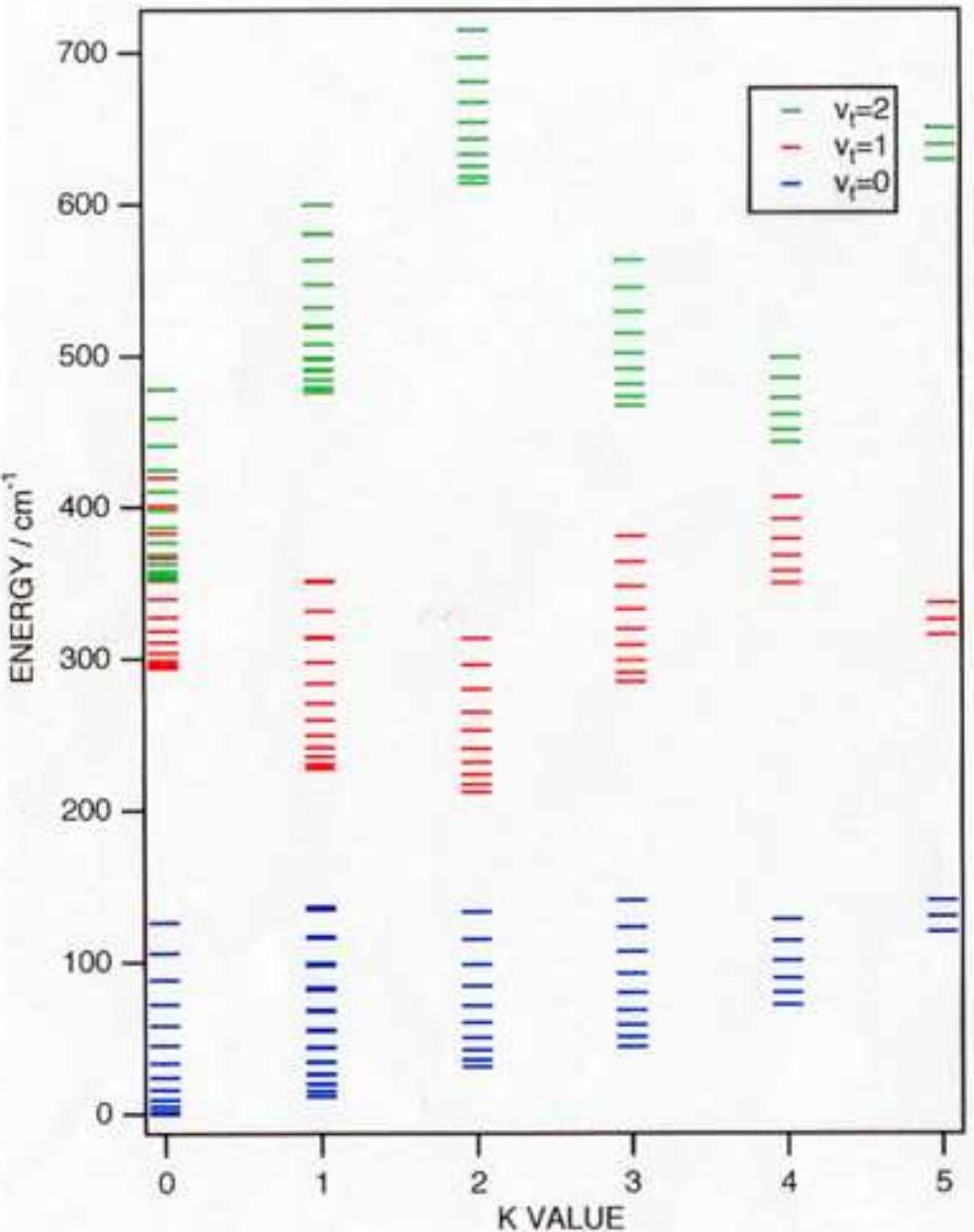
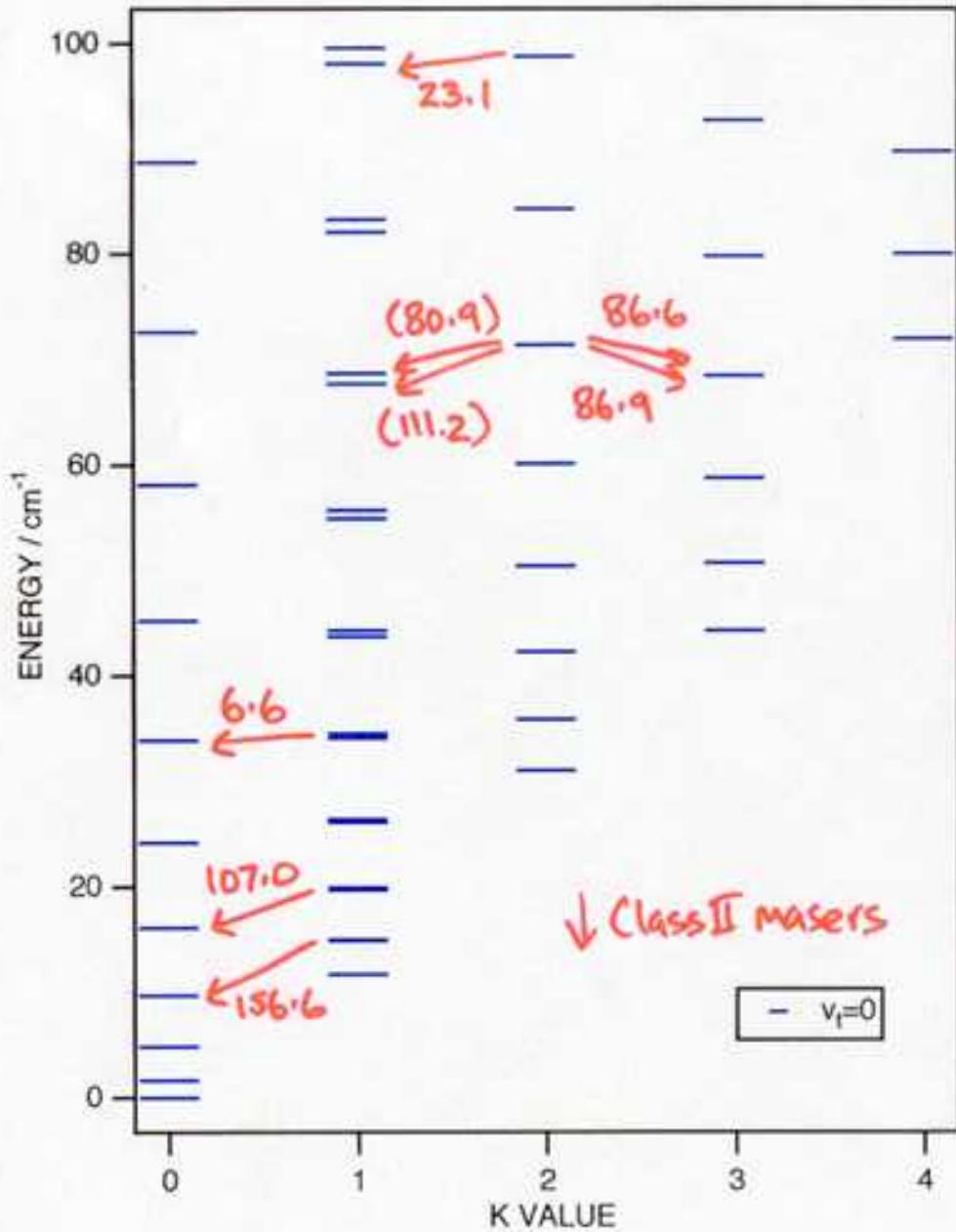
What are the physical conditions in the maser regions?

G345.01+1.79



G345.01+1.79 (6.7 GHz ATCA)

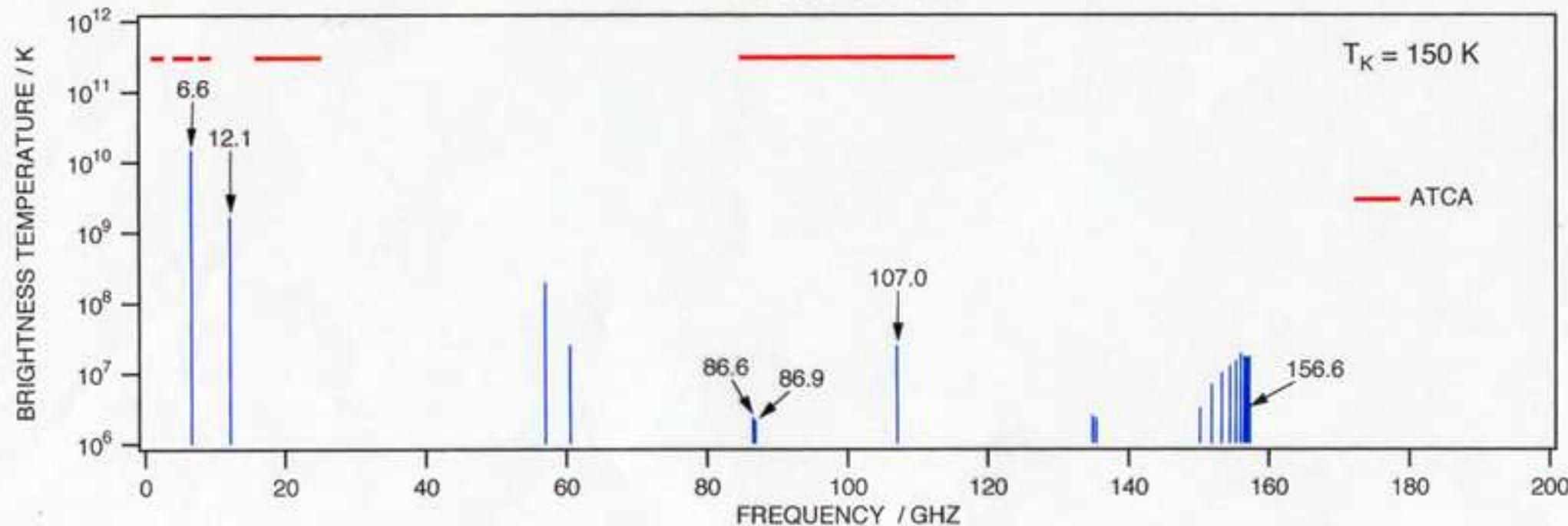
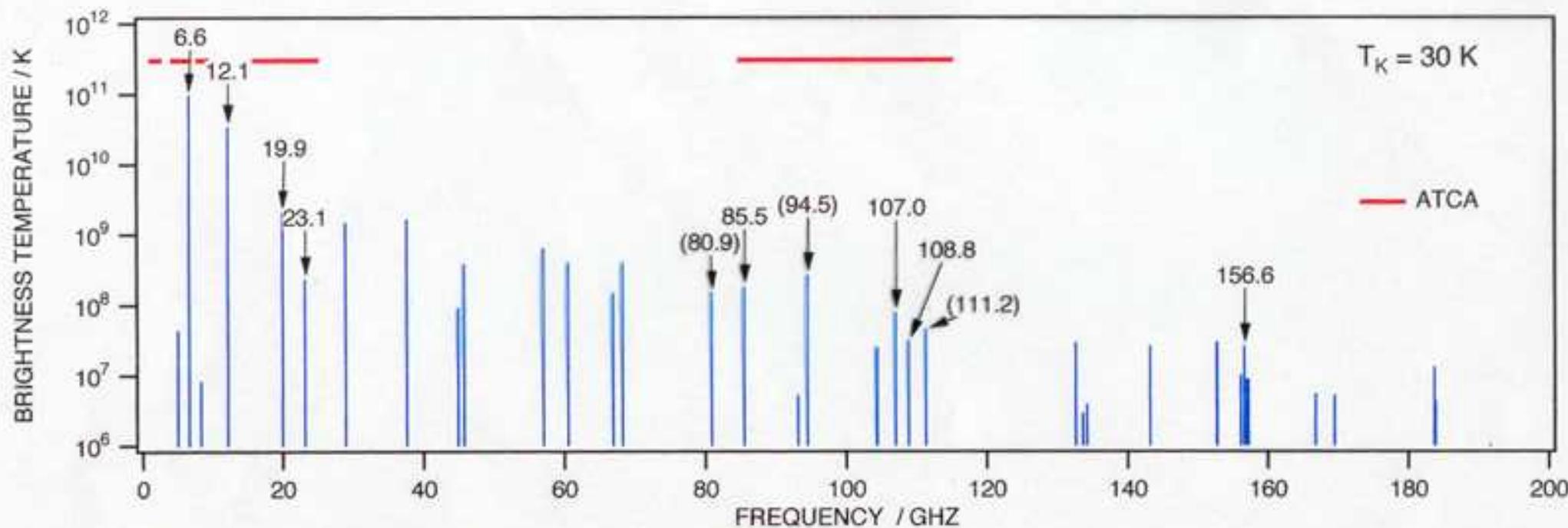


A SPECIES CH₃OH, v_t=0,1,2A SPECIES CH₃OH, TORSIONAL GROUND STATE

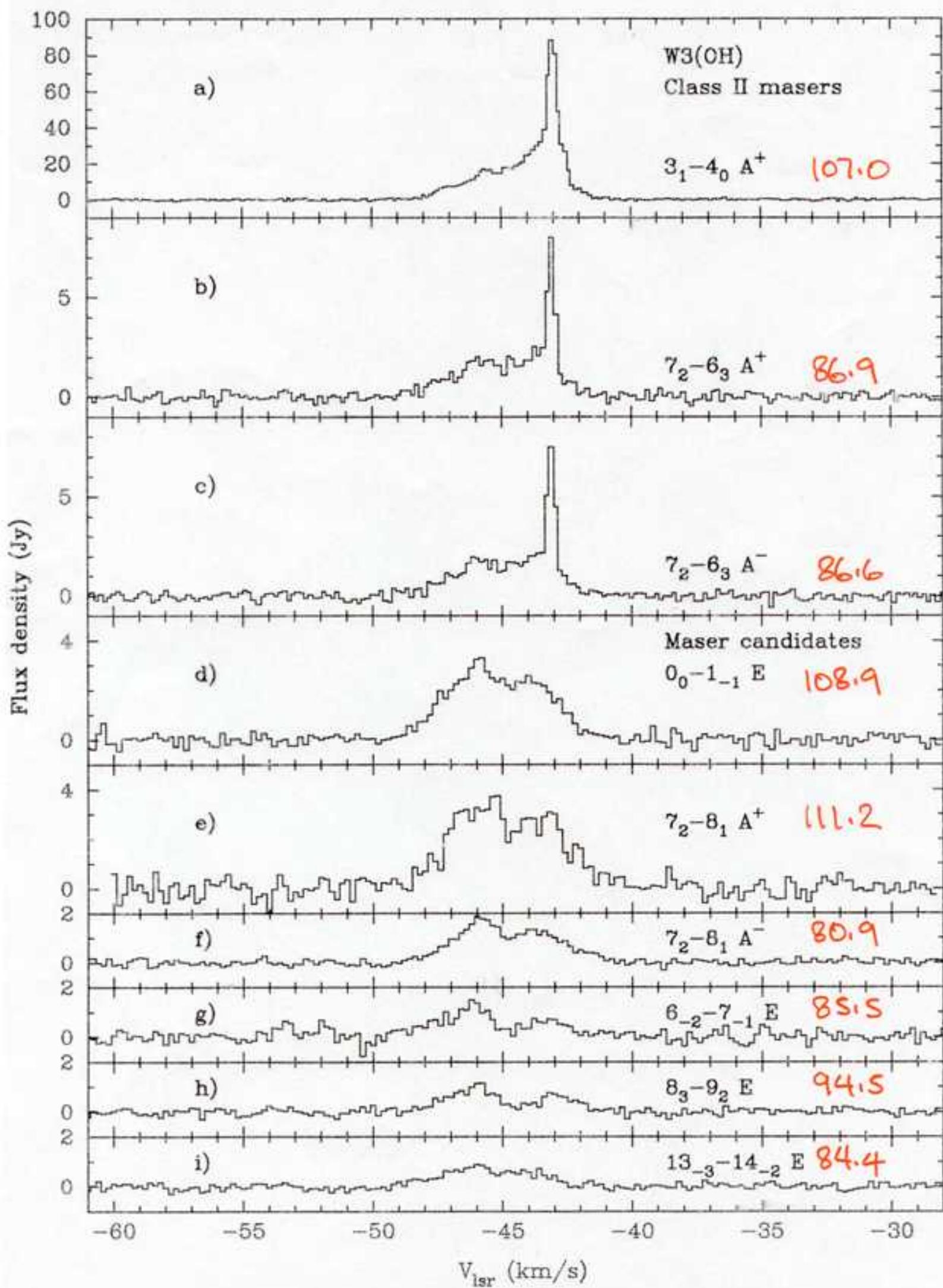
Sobolev-Deguchi maser pumping model

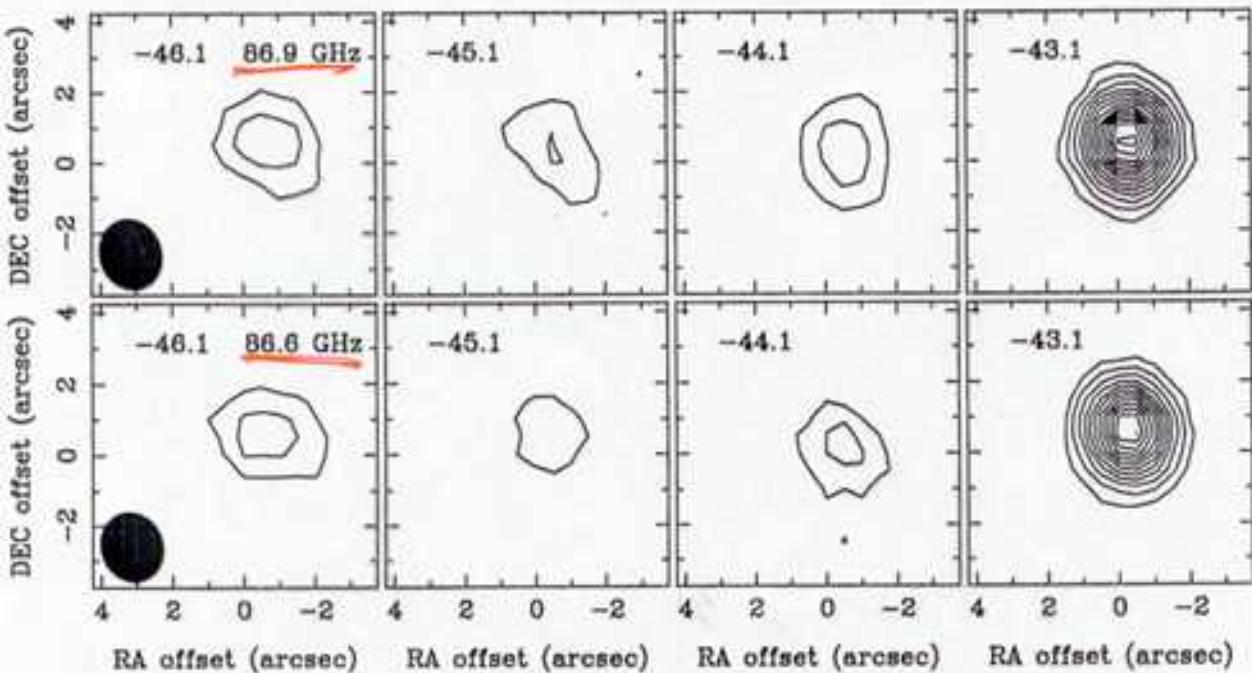
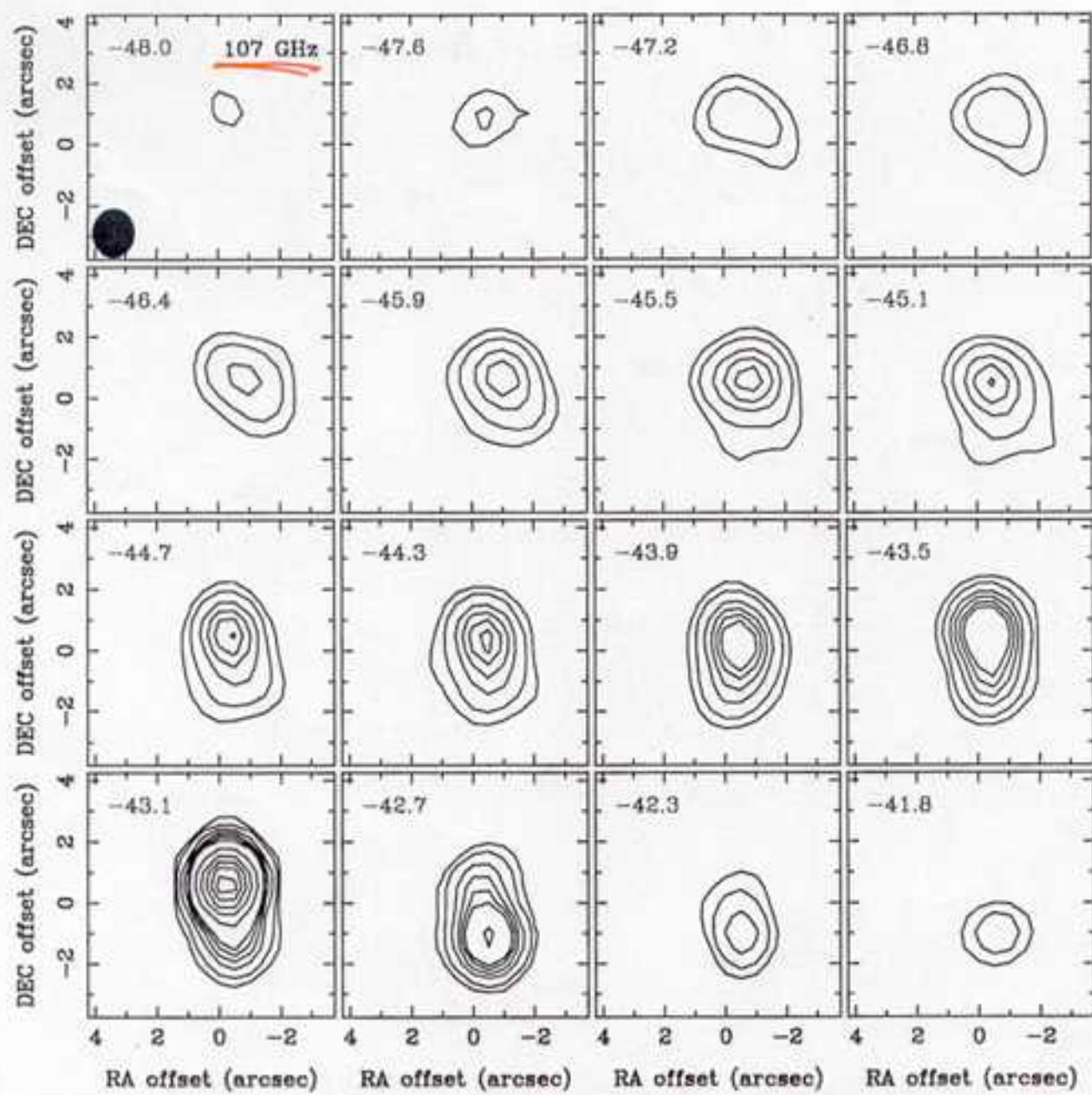
- IR radiation from warm dust (T_d) pumps methanol molecules to torsionally excited states
- Collisions with cool gas (T_k, n_H) determine radiative/collisional balance
- Large column density of methanol (N_m) for strong masers
- Beamed geometry (ϵ^{-1})
- Masers amplify ucHII background continuum (W_{HII})

PREDICTED CLASS II METHANOL MASER SPECTRA



BIMA Observations of W3(OH)



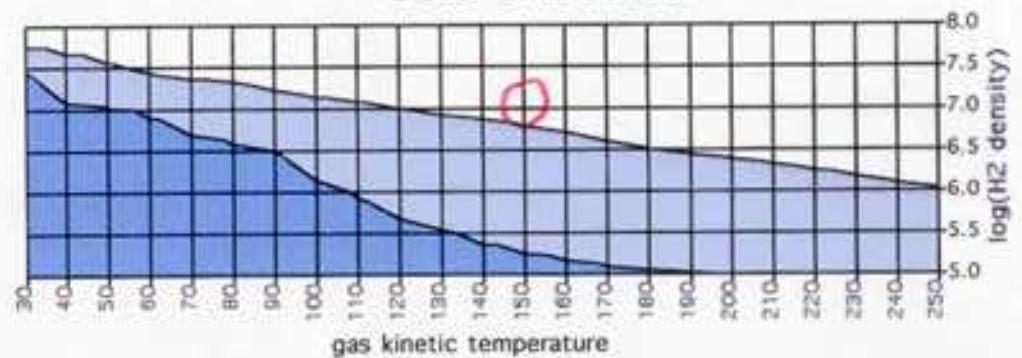


W3 (OH)

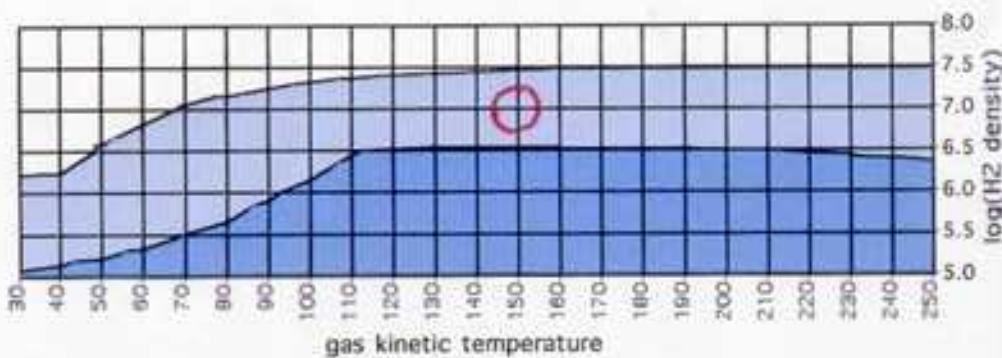
BIMA

CONTOURS OF MASER BRIGHTNESS FROM MODEL CALCULATIONS

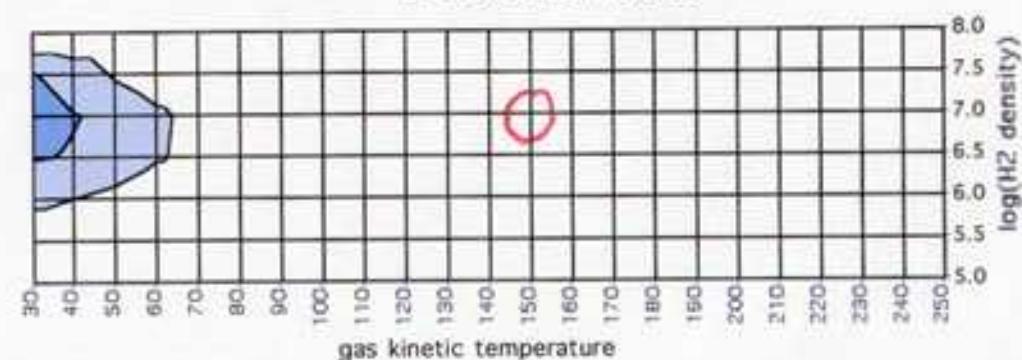
80.9 GHz maser



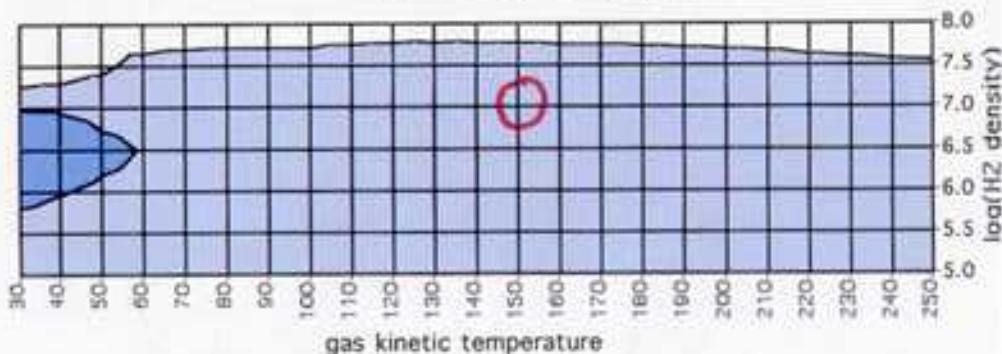
86.6 GHz maser



85.5 GHz maser



107.0 GHz maser



G345.01+1.79 SEST June 2000

