



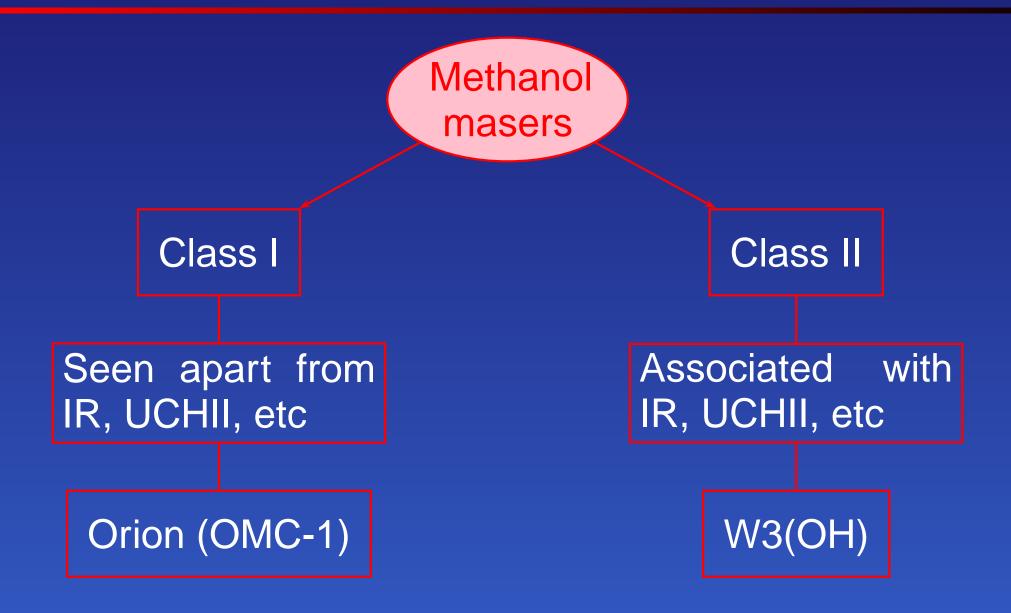
6.7 GHz and 25 GHz methanol masers in OMC-1

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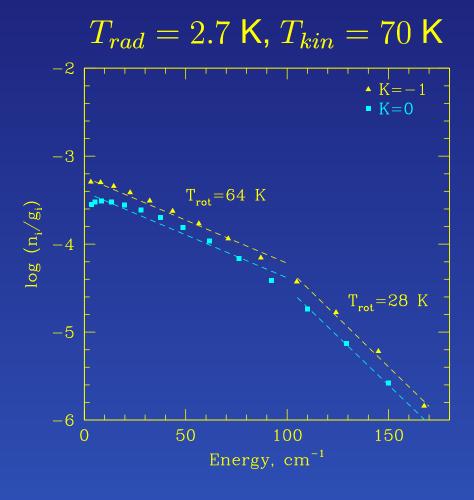
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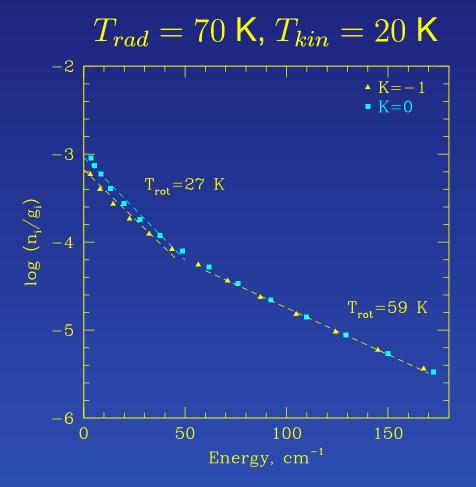
Two classes of methanol masers



Two classes: pumping difference



Class I, e.g. $4_{-1} - 3_0$ E at 36 GHz

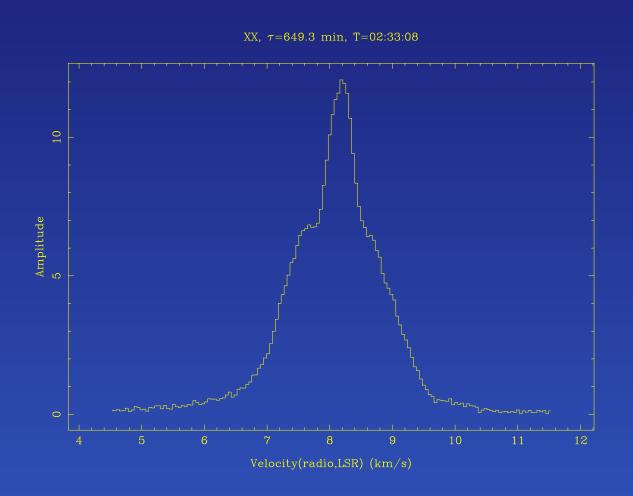


Class II, e.g. $2_0 - 3_{-1}$ E at 12 GHz

Theory: they may coexist

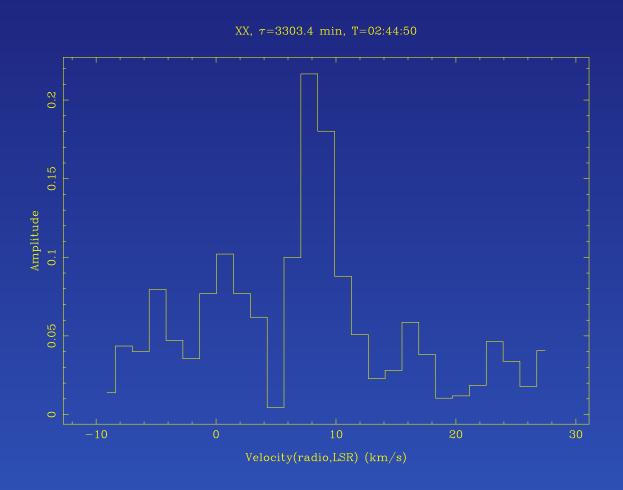
- Masers of different classes are often seen in the same sources, but apart from each other.
 This is not an interesting case.
- Modern models of the methanol maser pumping predict a weak 6.7 GHz maser (Class II) under the same conditions when the 25 GHz transition (Class I) becomes a maser.
- The OMC-1 has the brightest known 25 GHz methanol maser. It is worth to search a 6.7 GHz emission in this source.

OMC-1: 25 GHz spectrum



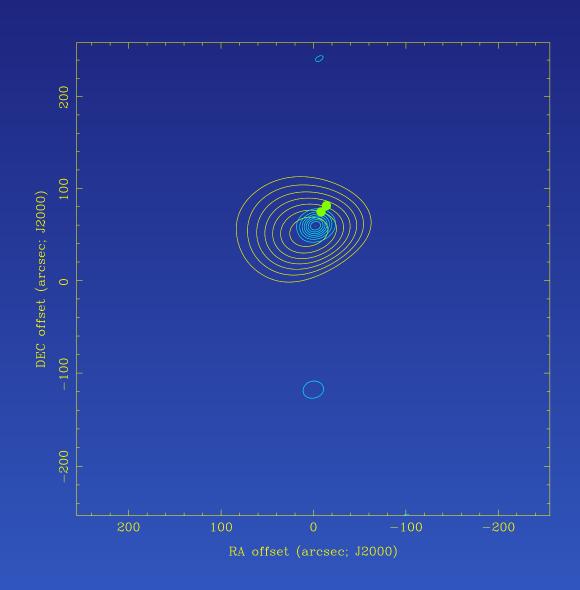
The 25 GHz spectrum (observed the first time with ATCA). the All baselines have been averaged together.

OMC-1: 6.7 GHz spectrum



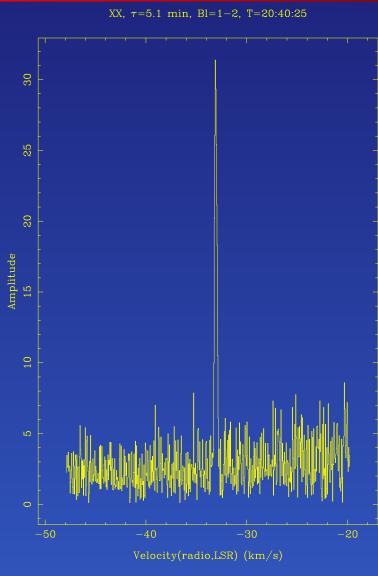
The 6.7 GHz spectrum. All baselines have been averaged together.

OMC-1: 6.7 GHz and 25 GHz

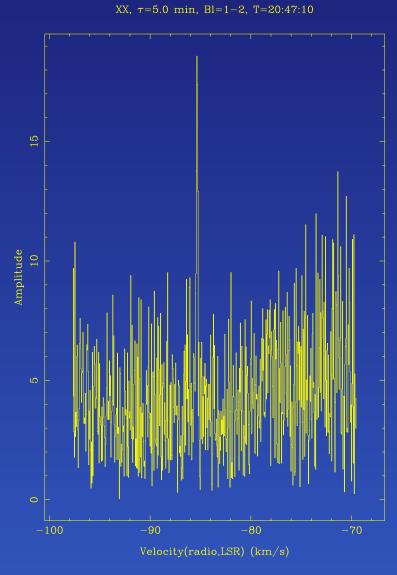


The 6.7 GHz emission may originate from the same place, where the 25 GHz maser spots were found.

25 GHz: new detections



305.36+0.2



333.23-0.05

Conclusions

- A 6.7 GHz emission has been detected towards OMC-1.
- The 6.7 GHz emission may be associated with the 25 GHz maser spots.
- No line has been detected for long baselines in the recent observations although the line was seen in archival data with a similar integration time.
- New 25 GHz methanol masers were discovered in 305.36+0.2 and 333.23-0.05.