

Mosaic(k)ing

ATNF Synthesis Workshop

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Overview

- ❖ Why mosaicking?
- ❖ Different Approaches
- ❖ Mosaicking in Practice
- ❖ Mosaicking at ATCA

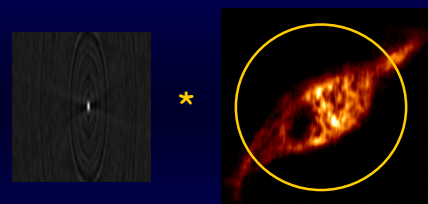
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Why Mosaicing?

- ❖ Source larger than primary beam: $\theta > \lambda/D$
 - o not limited by primary beam size
- ❖ Source structure at large scales: $\theta > \lambda/b_{\min}$
 - o better uv coverage

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Single Pointing



$$I(\vec{x}) = D(\vec{x}) * (P(\vec{x}) \times S(\vec{x})) + \sigma(\vec{x})$$

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Mosaicing: Individual Approach

- ❖ Observe single pointings
- ❖ Reduce them separately
- ❖ "Stitch" images together
- ❖ Disadvantages:
 - Deconvolution non-linear (cleaning bowl)
 - Overlap regions noisy (primary beam shape)
 - Expensive in observing time
 - Does not use all uv information
- ❖ OK for high-res where source \ll field

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$$M(\vec{x}) = W(\vec{x}) \frac{\sum_i P(\vec{x} - \vec{x}_i) \times I(\vec{x})}{\sum_i P^2(\vec{x} - \vec{x}_i)}$$

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The Special Stuff: Ekers & Rots (1979)

- ❖ Single dish consists of many sub-interferometers

diameter D

D

- ❖ Single pointing cannot recover information: one visibility, many "interferometers"
- ❖ More pointings (e.g. scanning) can image

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Ekers & Rots Trick 2

- ❖ An interferometer does not just measure baseline b
- ❖ It measures spacings between b-D and b+D

b

"Fourier coverage"

b-D b b+D

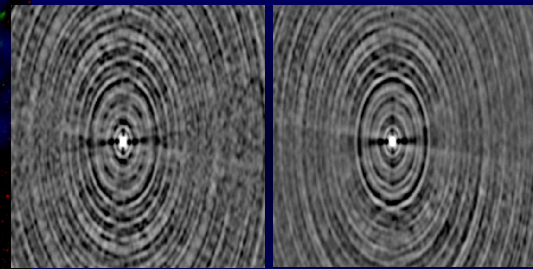
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Ekers & Rots 3

- ❖ Again, one pointing cannot recover this information (individual approach)
 - ❖ Many pointing together can recover this "extra" coverage (joint approach)
- ⇒ Use all uv data from all pointings simultaneously: better uv-coverage, better beam, better image

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Comparison of beams

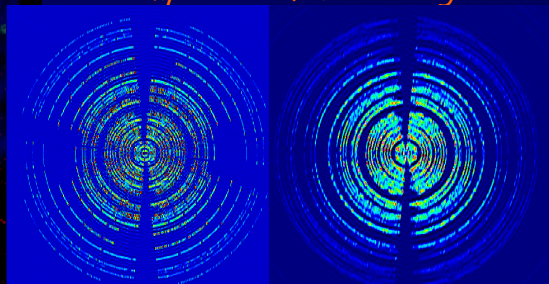


Individual

Joint

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Comparison of uv coverage



Individual

both radial (extra uv coverage)
and azimuthal (adjacent pointings)
improvement

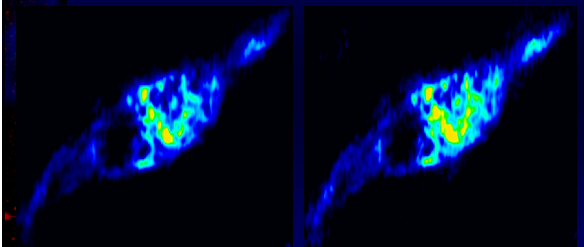
Joint

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Mosaicing differences

individual

joint

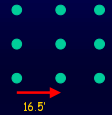


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Mosaicing Grid

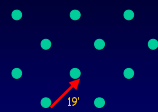
❖ Rectangular grid

- o Nyquist sampling $\theta = \lambda/2D$
- o = 16.5' at 21cm at ATCA



❖ Hexagonal grid

- o Nyquist sampling $\theta = (2/\sqrt{3})\lambda/2D$
- o = 19.0' at 21cm at ATCA
- o Solid angle coverage better



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Drive times

- ❖ For short drives: $t = 2(s/a)^{1/2}$
- ❖ $a = 400 \text{ deg/min}^2$
- ❖ so for 0.5 degree drive, $t = 4.3 \text{ sec}$

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Complete uv sampling

- ❖ One baseline measures region in uv plane with size 2D
- ❖ At transit, time between independent points is $\tau = (86400/2\pi)(2D/L) \text{ sec}$ where D=diameter, L=longest baseline
- ❖ Nyquist sampling for N pointings: dwell time is $\tau/2N \text{ sec}$
- ❖ E.g. 375m: 13.4 min, 750m: 6.7 min

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Schedule file (at\$sched:c373 lmc 01.sch)

```
Source LMC A Observer LSS_DMCC_RJS Project_id C373
RA 04:05:57.400
Dec -73:37:54.41
Epoch 12000.0
CalCode
Scan length 00:20:00.00
Start 101
Sweep 0
Sweep 900
Scype MOSAIC
Pointing GLOBAL
Mode STANDARD
Averaging 1
Dual Freq: ON
PolarOffset: 0.0
Config file full_A_1004-128
```

Mosaic file (at\$mosaic:lmc_a.mos)

```
#DAY 1
#LMC grid first field at 04:05:57.4 -73:37:54.4 12000)
# ( 61.48903 -73.63179 deg RA DEC
# degrees
# d(RA) d(DEC) cyc Name RA DEC
# 0.0000 0.0000 1 SLMC_0001 04:05:57.4 -73:37:54.4
# 0.8280 0.2127 1 SLMC_0002 04:09:15.6 -73:26: 8.6
# 0.5905 0.3518 1 SLMC_0003 04:08:19.1 -73:06:35.9
# 1.3851 0.7367 1 SLMC_0004 04:11:29.8 -72:53:42.2
# 1.1485 1.0452 1 SLMC_0005 04:10:32.5 -72:35:11.7
# 1.9119 1.2622 1 SLMC_0006 04:13:36.2 -72:20:10.5
```

Sched files

Logs, until scan length filled

Degrees of polar rotation from REFERENCE position

(@LSS)©

Mosaicking with MIRIAD

- ❖ Linear mosaic with LINMOS (after CLEAN/RESTOR on indiv pointings)
- ❖ MIRIAD has multi-source capability: no need to split files
- ❖ Making cubes: use INVERT with option=mosaic
- ❖ Deconvolving: Maximum entropy MOSMEM or cleaning with MOSSDI
- ❖ Restoring with RESTOR

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Other mosaicing software

- ❖ Classic AIPS:
 - LTESS (linear)
 - VTESS, UTESS (non-linear)
- ❖ AIPS++
 - part of IMAGER tool
 - has MEM and clean variations
 - also has "mosaicwizard"

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Summary

- ❖ Mosaicing is essential for widefield imaging
- ❖ Mosaicing is easy
- ❖ Use individual approach for high angular resolution observations where $b \gg D$: minimize primary beam errors
- ❖ Use joint approach where $b \approx D$: extra uv information

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Literature

- ❖ Ekers & Rots (1979), IAU Coll 49, 61
- ❖ Cornwell (1988), A&A, 202, 316
- ❖ Cornwell, Holdaway & Uson (1993), A&A, 271, 697
- ❖ Sault, Staveley-Smith, Brouw (1996), A&A Sup, 120, 375
- ❖ Holdaway (1998), ASP Conf 180, Ch. 20
- ❖ Sault & Killeen, Miriad manual, Ch. 21

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